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# What are the optimum discrete angles to use in thermal-infrared radiative transfer calculations?

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#### Abstract

As computer power increases there is a need to investigate the potential gains of using more than two streams in the radiative transfer calculations of weather and climate models. In this paper, seven quadrature schemes for selecting the zenith-angles and weights of these streams are rigorously evaluated in terms of the accuracy of thermalinfrared radiative transfer calculations. In addition, a new method is presented for generating 'Optimized' angles and weights that minimize the thermal-infrared irradiance and heating-rate errors for a set of clear-sky training profiles. It is found that the standard approach of applying Gauss-Legendre quadrature in each hemisphere is the least accurate of all those tested for two and four streams. For clear-sky irradiance calculations, 'Optimized' quadrature is between one and two orders of magnitude more accurate than Gauss-Legendre for any number of streams. For all-sky calculations in which scattering becomes important, a form of Gauss-Jacobi quadrature is found to be most accurate for between four and eight streams, but with Gauss-Legendre being the most accurate for ten or more streams. The fact that no single quadrature scheme performs best in all situations is because computing irradiances involves two different integrals over angle, and the relative importance of each integral depends on the amount of scattering taking place. Additional optimized quadratures for clear-sky and all-sky calculations with 4–8 streams are presented that constrain the relationships between angles in a way that reduces the number of exponentials that need to be computed in a radiative transfer solver.

#### **KEYWORDS:**

two-stream approximation, discrete ordinate method, heating rate, numerical convergence, infrared radiative transfer

#### | INTRODUCTION

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The discrete ordinate method (Chandrasekhar 1960) is widely used for 1D plane-parallel radiative transfer problems, and involves discretizing the diffuse radiation field into 2N zenith angles. The simplest two-stream (N = 1) version was originally proposed by Schuster (1905), and over a century later the two-stream approach still underpins almost all weather and climate models worldwide. At the European Centre for Medium-Range Weather Forecasts (ECMWF), the radiation scheme accounts for only 3.5% of the computational cost of the 9-km-resolution model (Hogan and Bozzo 2018), and this is likely to decrease with the planned upgrade of gas optical properties (Hogan and Matricardi 2022). It is therefore important to investigate the additional accuracy, and possibly reduction in regional temperature biases, that could be obtained by increasing the number of streams from two to four or maybe more. In the thermal infrared (hereafter 'longwave'), this could be

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facilitated by the finding of Fu et al. (1997) that by approximating the treatment of scattering, the additional cost of the radiative transfer solver in moving from two to four streams can be reduced from a factor of 9.0 to a factor of 1.8.

This paper is concerned with determining the optimal disrete angles to use for the modest number of streams that could he afforded in the longwave part of the radiation scheme in a weather or climate model. One might think that this matter yould be settled: the undisputed standard for performing refernce plane-parallel radiation calculations in the shortwave and longwave is the Discrete Ordinate Radiative Transfer package DISORT' (Stamnes et al. 1988), which only offers 'doublefauss' quadrature (Sykes 1951), whereby the cosine of the zenith angle,  $\mu$ , is discretized separately in each hemisphere sing Gauss-Legendre quadrature. Yet in the two-stream case this results in the single quadrature point  $\mu_1 = 1/2$  corresponding to a discrete zenith angle of  $\theta_1 = 60^\circ$ . Despite having been used in two-stream schemes for thermal radiative transfer roblems (e.g., Schuster 1905, Toon et al. 1989, Hogan 2019), it has been found (e.g., Rodgers and Walshaw 1966) that much more accurate irradiances and heating rates are achieved usng the Elsasser (1942) value of  $\mu_1 = 1/1.66$ , corresponding  $\theta_1 = 53^\circ$ , and indeed this value is used in most weather and climate models worldwide. If Gauss-Legendre quadrature suboptimal for two-stream longwave radiative transfer, it is legitimate to question its use for larger numbers of streams.

Li (2000) made an important contribution in his investigation of alternative longwave Gaussian quadrature schemes, and tested them with up to six streams using a single atmospheric in the absence of scattering. This is a springboard for the present study. In section 2 we show how the two angular integrals in longwave radiative transfer (one to represent scattering and the other to convert the radiance distribution to an tradiance) lead to conflicting requirements on the 'optimal' choice of quadrature angles. We give a physical explanation why Li's 'Gaussian quadrature with different moment powers' eads to better performance than Gauss-Legendre.

An alternative to Gaussian quadrature was proposed by acis and Oinas (1991), who used three values per hemisphere of  $\mu_1 = 0.1$ ,  $\mu_2 = 0.5$  and  $\mu_3 = 1$ , with hand-tuned weights. In a no-scattering longwave solver, the transmittance of a layer of optical depth  $\tau$  for stream *i* is  $T_i = \exp(-\tau/\mu_i)$ . The computational advantage of using  $\mu$  values that are multiples of each other is that the *N* exponentials may then be replaced by one exponential plus a few multiplications (in this case four:  $T_{tmp} = T_1 \times T_1$ ,  $T_2 = T_{tmp} \times T_{tmp} \times T_1$  and  $T_3 = T_2 \times T_2$ ). This quadrature is still used in the no-scattering longwave solver of the NASA Goddard Institute for Space Studies (GISS) climate model. 477870x, ja, Downloa

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In section 3 we demonstrate a new method in which angles and weights are chosen that formally minimize the meansquared error in irradiances and heating rates for a set of 50 training profiles. A variant of this method is to constrain the  $\mu$  values to be in a certain ratio, enabling the Lacis and Oinas (1991) optimization to be applied. In section 4, seven quadrature schemes are evaluated on 50 independent clear-sky (i.e. no-scattering) evaluation profiles for up to 32 streams. Then in section 5 a global model snapshot is used to evaluate the schemes in cloudy (scattering) situations, which involved modifying DISORT to use user-supplied quadrature angles. The conclusions in section 6 provide recommendations for the appropriate quadrature depending on number of streams and whether or not scattering is to be represented.

#### 2 | THEORETICAL BASIS

#### 2.1 | Scattering and irradiance integrals

The azimuthally-averaged longwave radiative transfer equation for a plane-parallel (i.e. horizontally homogeneous) atmosphere may be written as (e.g., Fu et al. 1997)

$$\mu \frac{dI_{\nu}(\tau_{\nu},\mu)}{d\tau_{\nu}} = I_{\nu}(\tau_{\nu},\mu) - (1-\omega_{\nu})B_{\nu} - \frac{\omega_{\nu}}{2} \int_{-1}^{1} p_{\nu}(\mu',\mu)I_{\nu}(\tau,\mu')d\mu', \quad (1)$$

where  $I_v$  is the radiance at frequency v,  $\tau_v$  is the optical depth of the atmosphere measured downwards from top-ofatmosphere (TOA) and acting as our vertical coordinate,  $\mu$  is the cosine of the zenith angle of the beam and is positive for upward propagating radiation and negative for downward,  $\omega_v$ is the single-scattering albedo of the medium,  $B_v$  is the Planck function and  $p_v(\mu', \mu)$  is the azimuthally averaged scattering phase function representing the probability of light travelling in direction  $\mu'$  being scattered into direction  $\mu$ . Thus the three terms on the right-hand side represent respectively loss of radiation by extinction out of the beam, gain of radiation by emission into the beam and gain of radiation by scattering into the beam. For the remainder of the paper we drop the vsubscript for brevity.

The discrete ordinate method for approximating and efficiently solving (1) involves discretizing the radiation field into 2N discrete directions  $\mu_1$  to  $\mu_{2N}$  such that:

$$\mu_{i} \frac{dI(\tau, \mu_{i})}{d\tau} = I(\tau, \mu_{i}) - (1 - \omega)B$$
$$- \frac{\omega}{2} \sum_{j=1}^{2N} w_{j}' p(\mu_{j}, \mu_{i}) I(\tau, \mu_{j}), \qquad (2)$$

where  $w'_{i}$  is the weight to be applied to direction  $\mu_{i}$  in the numerical quadrature scheme. So how should we choose the optimum quadrature angles and weights? The scattering integral in (1) is unweighted by  $\mu'$ , suggesting that if the integrand an be approximated by a polynomial in  $\mu'$ , Gauss-Legendre vould be optimal. Indeed, Chandrasekhar (1960) proposed Gauss-Legendre quadrature across the full range  $-1 \le \mu' \le 1$ , which in the case of N = 1 results in quadrature angles of  $t_{1,2} = \pm 3^{-1/2}$ . One problem with this choice is that there is nvariably a discontinuity in the radiance field at the horizon that is poorly sampled by Gauss-Legendre quadrature because it places angles more densely at the ends of the range than in he middle. This led Sykes (1951) to propose 'double-Gauss' quadrature whereby the two  $\mu'$  ranges of -1 to 0 and 0 to 1 are iscretized separately using Gauss-Legendre quadrature. This the approach taken in DISORT (Stamnes et al. 1988).

It is important to recognise that the scattering integral in (1) is not the only consideration when discretizing  $\mu$  space, esecially for longwave radiative transfer where scattering tends not to be the dominant process, and in clear skies it is not important at all. In a weather or climate model we are concerned with irradiances, i.e. the power passing through a horizontal lane at a particular height, which for upwelling irradiance may be written in continuous form as

$$F(\tau) = 2\pi \int_{0}^{1} \mu I(\tau, \mu) d\mu,$$
 (3)

and similarly for the downwelling irradiance. We see immediately that the integral weights the radiance by the cosine of the zenith angle  $\mu$ ; indeed, this equation is a form of Lamert's Cosine Law. The Sykes (1951) approach to discretizing (3) is to use the same Gauss-Legendre angles and weights as previously, resulting in

$$F(\tau) \simeq 2\pi \sum_{j=1}^{N} w'_{j} \mu_{j} I(\tau, \mu_{j}), \qquad (4)$$

where the indices j = 1 to N correspond to the upward ropagating radiances at different angles. However, the Gauss-Legendre placement of  $\mu$  values symmetrically in the 0–1 interval is not likely to be optimal for computing longwave irhadiances; the weighting by  $\mu$  in (3) means that the optimal angles should be weighted more towards zenith ( $\mu = 1$ ) and nadir ( $\mu = -1$ ) than the horizon ( $\mu = 0$ ). This goes some way to explaining why in the two-stream case the Elsasser (1942) value of  $\mu_1 = 1/1.66 = 0.602$  performs better than the Gauss-Legendre value of 0.5 (see section 1).

Effectively, the Sykes (1951) approach treats (3) as an unweighted integral of the function  $\mu I(\tau, \mu)$ . If we treat it as the integral of  $I(\tau, \mu)$  weighted by  $\mu$ , then this suggests we should use alternative quadratures designed for weighted integrals,

leading to a discretization of the form

$$F(\tau) \simeq \pi \sum_{j=1}^{N} w_j I(\tau, \mu_j).$$
<sup>(5)</sup>

The angles  $\mu_j$  must be the same as used for discretizing the scattering integral in (2), even if calculated using an alternative method to Gauss-Legendre, but the weights are different since  $w_j$  folds in the  $\mu$  dependence, yet like  $w'_j$  must be normalized to satisfy

$$\sum_{j=1}^{N} w_j = 1.$$
 (6)

From this point until the end of section 4 we consider only clear-sky atmospheres in which (3) is the only integral to be discretized; then in section 5 we consider more realistic profiles containing clouds in which the scattering integral becomes important as well. At that point the relationship between the weights  $w'_i$  and  $w_j$  is provided.

#### 2.2 | Optimizing transmittance

In order to establish a theoretical basis for alternative quadrature schemes optimized for longwave irradiance calculations, consider the transmittance of a slab of optical thickness  $\tau$ illuminated from one side by isotropic radiation:

$$T(\tau) = 2 \int_{0}^{1} \exp\left(-\frac{\tau}{\mu}\right) \mu d\mu, \qquad (7)$$

which is the exponential integral of the third kind. The black solid line in Fig. 1a depicts  $T(\tau)$  computed numerically with extremely fine resolution in  $\mu$ . In the absence of scattering and emission, the transmittance of a slab is equivalent to the ratio of irradiances at the far and near ends of the slab. Therefore, optimizing for transmittance is similar to optimizing for irradiances in the full longwave radiative transfer problem. In the two-stream approximation, (7) reduces to

$$T_{\rm TS}(\tau, D) = \exp(-\tau/\mu_1) = \exp(-D\tau), \qquad (8$$

where all radiation is treated as propagating with zenith angle  $\theta_1 = \cos^{-1} \mu_1$ . This angle is usually expressed in terms of a diffusivity factor  $D = 1/\mu_1$ . In Fig. 1a we see that the Gauss-Legendre value of D = 2 underestimates transmittance for all optical depths, although is correct in the limit of very small optical depth. Elsasser's value of 1.66 performs much better; indeed, he derived it simply by fitting these two curves by eye. Nonetheless, it does tend to overestimate transmittance at low optical depth and underestimate it at high optical depth.

We can be more systematic than Elsasser by computing the root-mean-squared error (RMSE) in transmittance for different values of D. To do this requires a distribution of optical depths to be specified. We do this by assuming a uniform distribution



**Figure 1** (a) Transmittance *T* of a slab of varying optical depth  $\tau$  to isotropic radiation, and two approximations of the form  $T = \exp(-D\tau)$ . (b) Root-mean-squared error (RMSE) in transmittance versus *D* for this approximation, along with specific values of *D* discussed in the text (where 'Legendre', 'Laguerre' and 'Jacobi' refer to Gauss-Legendre, Gauss-Laguerre and Gauss-Jacobi quadrature) and the corresponding effective zenith angle is  $\theta_1 = \cos^{-1}(1/D)$ .

f transmittances, leading to the RMSE being given by

$$\text{RMSE}(D)^2 = \int_0^\infty \left[ T_{\text{TS}}(\tau, D) - T(\tau) \right]^2 \frac{dT(\tau)}{d\tau} d\tau, \quad (9)$$

which is shown versus D [and equivalently  $\theta_1 = \cos^{-1}(1/D)$ ] ific schemes discussed in this paper. We see that the Gausslegendre value of D = 2 is very far from optimal: the minimum error in fact occurs for D = 1.6145, although Elsasser's alue is close to the minimum. As will be demonstrated in later sections of this paper, there is no single 'correct' value for D ecause one value does not perfectly reproduce the transmitance for all optical depths in Fig. 1a, and the optical depth of the atmosphere varies strongly with wavelength and absorber mount. Even at a single wavelength, the spectral heating rate at a particular altitude depends on the transmittance to all ther layers as well as to the surface and TOA. This has not prevented proposals for parametrizing D as a function of optical depth (Zhao and Shi 2013, DeSouza-Machado et al. 2020) or using a different value for different parts of the spectrum (Feng and Huang 2019).

#### 2.3 | Gaussian quadrature schemes

Finally in this section we present the theoretical basis for alternative Gaussian quadrature schemes for longwave radiative transfer. While the schemes proposed are similar to those examined by Li (2000), some important additional insights are presented on why they should work better than Gauss-Legendre quadrature. An introduction to alternative Gaussian quadrature schems is provided by Press et al. (2007); the corresponding nodes and weights can be derived from the values in Tables 25.8 and 25.9 of Abramowitz and Stegun (1972), although we compute them using a free-software package implementing the algorithm described by Kautsky and Elhay (1982).

The choice of Gaussian quadrature scheme depends firstly on the functional form of the weight term, which in (5) and (7) is simply  $w(\mu) = 2\mu$  for integration in the interval  $0 \le \mu \le 1$ and is illustrated by the shaded area in Fig. 2a. Gauss-Jacobi quadrature deals with weights of general form  $w(x) = (1 - 1)^{1/2}$  $(x)^{\alpha}(1+x)^{\beta}$  for integration in the interval  $-1 \le x \le 1$ , so is appropriate for our problem with the substitutions  $\mu = (x+1)/2$ ,  $\alpha = 0$  and  $\beta = 1$ . For the two-stream problem (i.e. a single angle in the  $\mu$  interval from 0 to 1), this quadrature proposes  $\mu_1 = 2/3$ . This is shown as 'Jacobi 1' (because  $\beta = 1$ ) in Fig. 1b where despite being better than unweighted Gauss-Legendre quadrature, it is still clearly not the best choice, and indeed for more than two streams the same is found when applied to real atmospheric profiles. The problem is that to be well approximated by a Gaussian quadrature rule, the integrand should be well approximated by a polynomial, with N-point Gaussian quadrature being exact for polynomials up to degree 2N-1. Figure 2a shows the transmittance of a single

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Figure 2 (a) Transmittance T of a slab of optical depth  $\tau$  to a beam propagating with a zenith-angle cosine of  $\mu$ ; the red shading ndicates the weighting when integrating over  $\mu$  to compute the transmittance to isotropic incident radiation (see Eq. 7). (b) The same but after a change of variables  $s = \mu^{1/3}$  to make the transmittance T(s) more amenable to numerical quadrature; the resulting weighting by  $w = 6s^5$  leads to the Gauss-Jacobi-5 quadrature rule. (c) As panel a but with an alternative change of variables whose exponential weight function leads to the Gauss-Laguerre quadrature rule.

beam of radiation, i.e. the function  $\exp(-\tau/\mu)$ , for four different optical depths, and it would clearly not be well fitted by a low-order polynomial, especially at low optical depths.

The situation is improved with a change of variables. For example, if we use a variable of integration  $s = \mu^{1/\gamma}$  then (7) secomes

$$T(\tau) = 2\gamma \int_{0}^{1} s^{2\gamma - 1} \exp\left(-\frac{\tau}{s^{\gamma}}\right) ds.$$
(10)

igure 2b shows that the integrand  $\exp(-\tau/s^{\gamma})$ , here for the case of  $\gamma = 3$ , varies somewhat more smoothly with s and herefore we should expect the quadrature scheme to be more ccurate. The weight function now has the form w(s) = $2\gamma s^{2\gamma-1}$ , which integrates to unity in the s interval 0–1 and shown by the shaded area in Fig. 2b. Gauss-Jacobi quadrature may still be used but with  $\beta = 2\gamma - 1$ , leading to the reight function taking the form  $w(s) = (\beta + 1)s^{\beta}$ . Table 25.8 of Abramowitz and Stegun (1972) provides the nodes and weights for  $\beta$  up to 5 and N up to 8, or an off-the-shelf softvare package for computing Gauss-Jacobi quadrature can be used for any value of  $\beta$  and N. The resulting set of nodes  $s_1$ to  $s_N$  need to be transformed back to  $\mu$  space with  $\mu_i = s_i^{\gamma}$ . Figure 1b shows the two-stream performance of Gauss-Jacobi quadrature for a change of variables with  $\gamma = 2$  ('Jacobi 3') and  $\gamma = 3$  ('Jacobi 5'), both of which are significantly nearer the minimum in the error curve. Li (2000) also explored the use of various Gauss-Jacobi quadrature schemes, although what we refer to as 'Gauss-Jacobi- $\beta$ ' quadrature, he referred to as 'Gaussian quadrature with moment power  $\beta$ ' (but note that he used the symbol *m* in place of  $\beta$ ). Li (2000) did not explain why

values of  $\beta$  larger than 1 should perform better, so hopefully the explanation here in terms of a change of variables resulting in the integrand being closer to a polynomial, for which Gaussian quadrature is designed, is valuable.

An alternative change of variables involves the use of a logarithmic scale for  $\mu$ , i.e.  $t = -2 \ln \mu$ , resulting in (7) becoming

$$T(\tau) = \int_{0}^{\infty} \exp(-t) \exp\left[-\frac{\tau}{\mu(t)}\right] dt.$$
 (11)

It is apparent from Fig. 2c that this makes the integrands even better behaved, although it does increase the domain of integration from 0–1 to 0–∞, which will be shown later in the paper to be a disadvantage for large orders of quadrature N. The appropriate quadrature scheme for a weight function of  $w(t) = \exp(-t)$  is Gauss-Laguerre, whose nodes need to be tranformed back to  $\mu$  space with  $\mu_i = \exp(-t_i/2)$ . For N = 1, Gauss-Laguerre produces  $D = 1/\mu_1 = 1.6487 = e^{1/2}$ , which can be seen in Fig. 1b to be very close to the Elsasser value but a little closer to the minimum of the error curve.

Table 1 presents the angles and weights of the Gauss-Laguerre and Gauss-Jacobi-5 quadratures for N up to 4. We show Gauss-Jacobi only in the  $\beta = 5$  case because in section 4 it is found to be superior to all other values of  $\beta$ .

Li (2000) extended Gauss-Jacobi quadrature up to a moment power of  $\beta = \infty$ , in which limit off-the-shelf numerical schemes for computing nodes and weights no longer work. Nonetheless, he was able to compute the nodes and weights analytically up to N = 3. Intriguingly, his values (see his Eqs. A7, A9 and A13) match exactly those computed from 477870x, ja, Downlc

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**Table 1** The cosine-angles  $\mu$  and weights w for five quadrature schemes: Gauss-Laguerre, Gauss-Jacobi with  $\beta = 5$  and three 'Optimized' schemes, where the number of angles per hemisphere N ranges from 1 to 4. The suffix 'IR' indicates that a prescribed 'integer ratio' applies between each  $\mu$  value and the first (indicated in brackets), enabling the number of exponentials in downstream calculations to be be reduced. The suffix 'JP' indicates that an additional 'Jacobi prior' has been added to he cost function penalizing the difference between optimized  $\mu$  and w values and their Gauss-Jacobi-5 equivalents, making optimized-IRJP well suited for scattering atmospheres.

		N	Variable	Gauss-Laguerre	Gauss-Jacobi-5	Optimized	<b>Optimized-IR</b>	<b>Optimized-IRJP</b>
		1	μ	0.6065306597	0.6297376093	0.6096748751		
			w	1.0000000000	1.0000000000	1.0000000000		
		2	μ	0.1813898346	0.2509907356	0.1976969570	0.1828926897	0.2669139064
				0.7461018061	0.7908473988	0.7419416274	0.7315707589 (×4)	0.8007417192 (×3)
			w	0.1464466094	0.2300253764	0.1520985621	0.1352478522	0.2509036055
				0.8535533906	0.7699746236	0.8479014379	0.8647521478	0.7490963945
		3	μ	0.0430681066	0.1024922169	0.0661385934	0.0675169363	0.1073702810
				0.3175435896	0.4417960320	0.3440369508	0.3375846814 (×5)	0.4294811240 (×4)
				0.8122985952	0.8633751621	0.8156973793	0.8102032354 (×12)	0.8589622480 (×8)
			w	0.0103892565	0.0437820218	0.0197413567	0.0197437659	0.0445786516
				0.2785177336	0.3875796738	0.2857816420	0.2746853796	0.3679447208
				0.7110930099	0.5686383044	0.6944770013	0.7055708545	0.5874766276
	-	4	μ	0.0091177205	0.0454586727	0.0259142819	0.0263733596	0.0468366244
				0.1034869099	0.2322334416	0.1420093170	0.1318667980 (×5)	0.2341831222 (×5)
				0.4177464746	0.5740198775	0.4312455503	0.4219737537 (×16)	0.6088761177 (×13)
				0.8510589811	0.9030775973	0.8441789463	0.8439475074 (×32)	0.9367324887 (×20)
			w	0.0005392947	0.0092068785	0.0030584329	0.0028332575	0.0093955477
				0.0388879085	0.1285704278	0.0539378694	0.0476214091	0.1353113093
				0.3574186924	0.4323381850	0.3332755640	0.3349230090	0.5081423593
				0.6031541043	0.4298845087	0.6097281337	0.6146223244	0.3471507838

Sauss-Laguerre quadrature, so we conclude that the latter is equivalent to Gauss-Jacobi quadrature with  $\beta = \infty$  and the apropriate changes of variables. It is beyond the scope of this paper to prove this equivalence mathematically, but it is conenient for those who might want to try the  $\beta = \infty$  quadrature proposed by Li (2000) for N > 3 to know that it can be comruted easily using an off-the-shelf Gauss-Laguerre algorithm, cr taken from Table 12.9 of Abramowitz and Stegun (1972).

Li (2000) argued that the most accurate quadrature schemes for longwave radiative transfer arise from using the highest moment power, i.e. the largest value of  $\beta$ ; the logic of this paragraph would suggest that this is Gauss-Laguerre quadrature. This was partially based on how close its two-stream diffusivity was to Elsasser's value of 1.66, although since the latter was fitted to the transmittance curve by eye we would argue that there is nothing particularly special about it, and we can see from Fig. 1b that in terms of transmittance at least, values of D in the range 1.57–1.66 are at least as accurate as Elsasser's value and there is little to choose between them. Ultimately, the test of a quadrature scheme should be in its performance on real

atmospheric profiles in both clear-sky and cloudy conditions, and this is pursued rigorously in sections 4 and 5.

#### **3 | OPTIMIZED QUADRATURE FOR CLEAR-SKY RADIATIVE TRANSFER**

Figure 1b demonstrated that, in the N = 1 case, Gaussian quadrature is not the only basis on which to select angles and weights: we can instead seek the value or values that minimize some scalar measure of overall error with respect to a reference calculation. In this section we extend this idea to more than one angle and use real atmospheric profiles. Hogan and Matricardi (2022) showed that the gas-optics part of a radiative transfer scheme could be significantly improved by optimizing the absorption coefficients of the look-up tables in order to improve the agreement with line-by-line calculations for a set of reference profiles. Here we take a similar approach but optimize the angles and weights of a quadrature scheme. We use the 50 present-day profiles of the 'Evaluation-1' dataset from

the Correlated K-Distribution Model Intercomparison Project (CKDMIP; Hogan and Matricardi 2020), which are defined on 54 atmospheric layers extend up to a pressure of 0.01 hPa and cover a wide range of temperature, humidity and ozone oncentrations. The well-mixed gases are N<sub>2</sub>, O<sub>2</sub>, CO<sub>2</sub>, CH<sub>4</sub>,  $l_2O$ , CFC-11 and CFC-12, and neither clouds nor aerosols are represented. The gas-optics model used is 'FSCK-32' described by Hogan and Matricardi (2022), which divides the entire longwave spectrum into 32 non-contiguous spectral inervals. Here and throughout this paper we neglect the effects of Earth curvature.

Following Hogan and Matricardi (2022), we minimize a ost function of the form

$$J = \sum_{k=1}^{n} \left\{ \sum_{j=1}^{m} h_{j} \left( H_{j}^{\text{quad}} - H_{j}^{\text{ref}} \right)^{2} + f \left[ \left( F_{\uparrow \text{TOA}}^{\text{quad}} - F_{\uparrow \text{TOA}}^{\text{ref}} \right)^{2} + \left( F_{\downarrow \text{surf}}^{\text{quad}} - F_{\downarrow \text{surf}}^{\text{ref}} \right)^{2} \right] \right\},$$
(12)

where the outer summation is over the *n* profiles, the inner summation is over the m layers,  $H_i$  denotes the heating rate of ayer j in K d<sup>-1</sup>,  $F_{\uparrow TOA}$  denotes upwelling irradiance at TOA h W m<sup>-2</sup>,  $F_{\downarrow surf}$  denotes downwelling irradiance at the surface, the superscript 'quad' represents values computed using he quadrature scheme and the superscript 'ref' represents reference values. Following Hogan (2010) we define the layer veights as proportional to the difference in the square-root of pressure across them, such that the contribution of the tropophere and stratosphere is approximately equal:  $h_j = (p_{j+1/2}^{1/2} \frac{1}{j-1/2} \frac{1}{2} \frac{1}{p_{m+1/2}}$ , where  $p_{j+1/2}$  is the pressure of the interface beween layers j and j + 1 (increasing from TOA towards the urface), and  $p_{m+1/2}$  is the surface pressure. The user-specified value f balances the irradiance and heating-rate errors, and we ave found that  $f = 0.02 \,(\text{K d}^{-1})^2 / (\text{W m}^{-2})^2$  provides a satisactory balance between the two in resulting radiative transfer alculations.

Since scattering can be neglected, the radiative transfer consists of projecting N beams of radiation up and down prough the atmosphere, and for simplicity we set the surface emissivity to unity. Thus the initial radiance of each upward-propagating beam at the surface is equal to the Planck function at the surface,  $B_{m+1/2}$ , while the initial radiance of each downward-propagating beam at TOA is zero. Assuming the Planck function varies linearly with optical depth across a layer, the solution to (2) for the downwelling radiance at angle

$$I_{i,j+1/2} = I_{i,j-1/2} \exp\left(-\frac{\tau_j}{\mu_i}\right) + \left[1 - \exp\left(-\frac{\tau_j}{\mu_i}\right)\right] \left(B_{j-1/2} - \frac{\mu_i \Delta B_j}{\tau_j}\right) + \Delta B_j,$$
(13)

where  $\tau_j$  is the zenith optical depth of the layer and  $\Delta B_j = B_{j+1/2} - B_{j-1/2}$  is the difference in Planck function across the layer. This may be applied recursively from TOA down to the surface, and similarly for the upward propagating beams. The irradiances may then be computed by applying (5) to the radiances at the interface between each layer. The reference calculations are performed with many hundreds of evenly spaced angles.

We define the 'state vector', **x**, containing variables to be optimized as the N angles  $\mu_1$  to  $\mu_N$  and all but one of the corresponding 'normalized' weights  $W_1$  to  $W_{N-1}$ , where  $W_i = w_i/2\mu_i$ . The use of normalized weights in the minimization makes each element of the state vector more similar in magnitude. The final normalized weight is computed from the others to ensure that (6) holds:

$$W_N = \frac{1}{\mu_N} \left( 1 - \sum_{j=1}^{N-1} W_j \mu_j \right).$$
(14)

The initial values of the state vector are such that the angles are evenly spread in  $\mu$  space and have equal normalized weight. The cost function is minimized by coding the radiative transfer (primarily Eq. 13) in C++, and using the automatic differentiation and optimization library 'Adept' (Hogan 2014) to compute the gradient of the cost function with respect to all elements of the state vector,  $\partial J / \partial x$ . The 'L-BFGS' minimization algorithm of Liu and Nocedal (1989) calls the radiative transfer repeatedly with different values for x, using  $\partial J / \partial x$  to find the x that minimizes J. The software is freely available at https://github.com/rjhogan/optimize-angles. The resulting 'Optimized' angles and weights are shown in Table 1 for N up to 4, and are evaluated against the Gaussian quadrature schemes using real atmospheric profiles in the next section.

Lacis and Oinas (1991) proposed an N = 3 quadrature in which the second and third  $\mu$  value was a multiple of the first, which meant that three exponentials could be replaced by one, thereby substantially speeding up radiative transfer calculations whose computational cost is often dominated by the exponential function. We implement this idea by choosing the integer ratios closest to those between the  $\mu$  values emerging from Optimized quadrature in Table 1, and then repeating the optimization but retrieving only the smallest  $\mu$  value and computing the others by applying these integer ratios. The resulting quadrature schemes up to N = 4 are shown in the penultimate 'Optimized-IR' column of Table 1. We may also find the closest integer-ratio quadrature scheme to an existing Gaussian scheme by adding a term to the cost function of the form

$$J_p = f_p \left[ \sum_{j=1}^{N} (\mu_j - \mu_j^p)^2 + (W_j - W_j^p)^2 \right], \quad (15)$$

where  $\mu_j^p$  and  $W_j^p$  are the 'prior' nodes and normalized weights of an existing Gaussian quadrature scheme, and  $f_p$  is the weight applied to the term. We find a value of 0.001 provides ne most accurate scheme. The final column of Table 1 shows the resulting 'Optimized-IRJP' quadrature scheme obtained using the Gauss-Jacobi-5 scheme as a prior. The motivation to emain close to Gauss-Jacobi-5 is that this scheme is found to be most accurate when scattering is introduced in section 5.

#### - | CLEAR-SKY EVALUATION OF QUADRATURE SCHEMES

In this section we evaluate various quadrature schemes in terms of their ability to predict clear-sky irradiances and heating ates. We use the 50 present-day profiles of the 'Evaluation-' dataset from CKDMIP (Hogan and Matricardi 2020), which are a different set of profiles from the Evaluation-1 dataset sed in section 3. Again we use the 'FSCK-32' gas-optics model described by Hogan and Matricardi (2022), and the ame clear-sky radiative transfer algorithm.

Figure 3a depicts the RMSE in surface downwelling and IOA upwelling irradiance versus the number of streams from 32, for seven different quadrature schemes. The least accurate is Gauss-Legendre quadrature, also known as doubleiauss (Sykes 1951), followed by Gauss-Laguerre which is typically 5 to 8 times more accurate for a given number of treams. Both exhibit approximately fourth-order convergence, .e. when the number of streams is increased by a factor of f, RMSE reduces by around a factor of  $f^4$ .

As explained in section 2, Gauss-Jacobi quadratures may be constructed via a change of variables  $s = \mu^{1/\gamma}$  resulting in the ntegration being weighted by  $s^{\beta}$  where  $\beta = 2\gamma - 1$ . We have tested a range of integer values of  $\gamma$ , and while the performance of values of 2, 3 and 4 is quite similar, we find  $\gamma = 3$  (i.e.  $\beta = 5$ ) is superior, and this is shown as 'Gauss-Jacobi-5' in Fig. 3a and the remaining plots in this paper. This quadrature scheme is superior to Gauss-Laguerre for all numbers of streams except 2; for 32 streams it is more than 40 times more accurate than Gauss-Laguerre and 400 times more accurate than Gauss-Legendre. Recall that Gauss-Laguerre quadrature is the same as Gauss-Jacobi in the limit of  $\gamma \rightarrow \infty$ . The 'Optimized' quadrature scheme generated using the method described in section 3 performs a little better than Gauss-Jacobi-5 up to 24 streams, while the 'Optimized-IR' scheme (enforcing integer 477870x, ja, Downloa

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ratios between the retrieved  $\mu$  values) for N between 2 and 4 performs only very slightly worse.

The reason that Gauss-Laguerre performs worse than Gauss-Jacobi-5 and Optimized quadrature for larger numbers of streams can be explained by the change of variables of  $t = -2 \ln \mu$  described in section 2, which transforms an integral in the range 0–1 to one in the range 0–∞. For large *N*, Gauss-Laguerre quadrature places many quadrature points at large values of *t* but with vanishingly small weights. In angular space this wastes quadrature points on zenith angles very close to 90°, so the performance of the scheme in terms of RMSE is as if it had fewer quadrature points overall. This can already be see for N = 4 in Table 1: Gauss-Laguerre has a node with a zenith angle of 89.5° but only 0.05% of the energy.

Figure 3a also shows the performance of Elsasser's value for N = 1, and the Lacis and Oinas (1991) coefficients for N = 3 ( $\mu_1 = 0.1$ ,  $\mu_2 = 0.5$ ,  $\mu_3 = 1$ ,  $w_1 = 0.0432$ ,  $w_2 = 0.5742$  and  $w_3 = 0.3826$ ). The latter is better than Gauss-Legendre but worse than all the other quadrature schemes.

We next evaluate the heating rates computed by the various quadrature schemes. Figures 3b and 3c depict the RMSE in tropospheric and stratospheric/mesospheric heating rates, respectively, according to whether the pressure is greater than or less than 100 hPa. In each case we follow section 3 and weight each layer by the difference in the square-root of pressure between the bottom and top of the layer. This time the order of convergence of Gauss-Legendre quadrature with increasing Nis closer to 2.8. Again, Gauss-Laguerre is more accurate than Gauss-Legendre, with Gauss-Jacobi-5 and then the Optimized quadrature even more accurate especially for larger numbers of streams. The exception is for stratospheric/mesospheric heating rates where for four streams we find that Gauss-Laguerre is clearly the most accurate. This again highlights that one set of quadrature points is not necessarily optimum for both irradiances and heating rates.

Figure 4 reveals that stratospheric/mesospheric heatingrate errors are largely associated with biases centred on the stratopause at around 1 hPa, and scale with the shape of the CO<sub>2</sub>-dominated cooling-rate profile (e.g. Fig. 5 of Hogan and Matricardi 2022). Gauss-Legendre quadrature over-predicts the cooling rate in all cases (negative heatingrate bias), while Gauss-Jacobi-5 tends to under-predict the cooling but with an amplitude somewhat smaller. For the twostream (N = 1) case, very similar performance is found for Gauss-Laguerre, Optimized quadrature and Elsasser's D =1.66. This is because, as shown in Fig. 1b, these schemes all have very similar values of D. For 4 streams (N = 2), Gauss-Laguerre quadrature is clearly the best with virtually no bias in stratospheric heating rates, followed by Optimized and Optimized-IR. For 6 streams, these three schemes perform similarly well.



**Tigure 3** Root-mean-squared error (RMSE) in clear-sky (a) surface downwelling and top-of-atmosphere (TOA) upwelling longwave irradiances, (b) heating rate for pressure, p, greater than 100 hPa and (c) heating rate for p less than 100 hPa, for the 50 CKDMIP 'Evaluation-2' profiles, as a function of the number of angular streams (equal to 2N) using seven different quadraure schemes. The reference calculations use Gauss-Jacobi-5 quadrature (with 64 streams) since this is the most accurate of the Gaussian schemes.



**Figure 4** (Solid lines) Atmospheric heating rate biases for various quadrature schemes with (a) 2, (b) 4 and (c) 6 angular streams, sing the 50 CKDMIP 'Evaluation-2' clear-sky profiles. The shaded areas for Gauss-Legendre and Gauss-Jacobi-5 quadrature encompass 95% of the profiles (estimated as 1.96 times the standard deviation). Note that Elsasser is only for 2 streams, Lacis only for 6 streams and Optimized-IR only for 4-6 streams.

#### 5 | ALL-SKY EVALUATION OF QUADRATURE SCHEMES

In this section we extend the evaluation to cloudy skies in which scattering becomes important. A modified version of the offline ecRad radiation scheme (Hogan and Bozzo 2018) has been used in which DISORT (Stamnes et al. 1988) has been embedded as a solver. This way we take advantage of ecRad's treatment of the optical properties of gases and clouds, and again use the fast FSCK-32 longwave gas-optics scheme. We have confirmed that the results presented in this section are insensitive to the gas-optics scheme used, having performed sensitivity tests (not shown) with the slower RRTMG (Rapid Radiative Transfer Model for Global Circulation Models; Mlawer et al. 1997) scheme also available in ecRad. We use Mie theory to generate phase functions for liquid droplets and the Baum et al. (2014) 'general habit mixture' phase functions for ice particles. We run this ecRad-DISORT model on

a 3D global snapshot of the latest ECMWF reanalysis (ERA5) at 12 UTC on 11 July 2019 with 137 vertical levels and 1° resolution. Meyer et al. (2022) extracted a 2D slice from this exact scene to demonstrate features of ecRad, although here ve neglect aerosols. Since DISORT cannot handle horizontal ub-grid cloud heterogeneity, we compute the total cloud cover assuming the same exponential-random overlap as in the operational ECMWF model, and assume clouds fill the cloudy fraction of each column homogeneously. The gridbox-mean rradiances are then computed as the weighted average of a clear-sky and a cloudy column. While this would not be suitable for an operational radiation scheme because it neglects ub-grid heterogeneity of cloud optical depth, it is adequate for estimating the relative accuracy of different quadrature chemes in all-sky conditions.

DISORT solves the coupled differential equations represented by (2), now with the summation term representing scattering processes that was neglected in section 4. DISORT las been modified so that the angles and weights may be configured by the user, rather than always being computed using Gauss-Legendre quadrature. The weights  $w'_j$  in (2) and  $w_j$  in 5) both satisfy the normalization given by (6). In the case of iauss-Legendre quadrature, the symmetry in the placement of  $\mu$  values in the interval 0–1 means that we can simply equate 4) and (5) to obtain  $w'_j = w_j/(2\mu_j)$ , which is implicitly assumed by the DISORT code. For the other quadratures conidered here, we have made a modification to DISORT so that the weights used in the scattering summation in (2) satisfy

$$w'_{j} = \frac{w_{j}/\mu_{j}}{\sum_{i=1}^{N} w_{i}/\mu_{j}}.$$
 (16)

Vhile it may seem unsatisfactory that a different scaling is required for the quadrature weights when used to approximate two different integrals, it is simply a consequence of one of ne integrals being weighted by  $\mu$  and the other not. The fact that quadratures symmetric in the range 0–1 (such as Gaussegendre) can use the same set of weights for both integrals by applying (4) instead of (5) is not a strong argument for them, ince energy can be conserved without this symmetry. Indeed, the widespread use of Elsasser's value demonstrates this in the two-stream case. Fu et al. (1997) stated that the two-stream renationship between irradiance and radiance is  $F = 2\pi\mu_1 I_1(\mu_1)$ , but that for choices of  $\mu_1$  other than 1/2 this should be replaced by  $F = \pi I_1(\mu_1)$ . Essentially the replacement of (4) with (5) is the 2*N*-stream generalization of Fu et al.'s statement.

Figure 5 depicts the RMSE in irradiances and heating rates for the ERA5 scene, but this time for all-sky conditions with scattering. The results for irradiances and tropospheric heating rates are very different from the clear-sky results in Figs. 3a and 3b. As before, the Gauss-Laguerre, Gauss-Jacobi-5 and Optimized quadrature schemes all outperform Gauss-Legendre 477870x, ja, Downloa

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with two and four streams, but for larger N they tend to approach an asymptotic RMSE regardless of N. By contrast, Gauss-Legendre quadrature continues to converge with N at much the same rate as for clear skies, and is the most accurate all-sky scheme for 16 streams. Gauss-Jacobi-5 is clearly better than Gauss-Laguerre and Optimized quadrature, still outperforming Gauss-Legendre quadrature up to and including the eight-stream case. Optimized-IRJP quadrature is intended to provide the closest quadrature to Gauss-Jacobi-5 but with  $\mu$  values in fixed integer ratios, and indeed it is only a little less accurate than Gauss-Jacobi-5. The heating-rate errors in the cloud-free stratosphere/mesosphere are, by contrast, virtually identical between Figs. 3c and 5c.

The explanation for the all-sky behaviour in Figs. 5a and 5b is as follows. The optimum quadrature schemes for the scattering integral in (1) and the irradiance integral in (3) are clearly different due to only one being weighted by  $\mu$ . In the longwave, gaseous absorption and emission is fundamental while scattering is of secondary importance. This means that errors in calculations of irradiances and heating rates using a small number of streams are dominated by errors in vertical transmission. Using a quadrature scheme optimized for the transmission problem means that errors reduce rapidly with increasing N, until errors due to scattering begin to dominate, for which such a quadrature scheme is less well suited. One might expect convergence to continue, albeit at a slower rate, but this is not observed in Fig. 5a. The reason is believed to be the fact that the phase function of clouds has a strong peak in the forward direction that cannot be resolved by discrete-ordinate radiative transfer calculations unless they have a much larger number of streams than considered here. DISORT addresses this with 'delta-M' scaling of the phase function (Wiscombe 1977): the number of streams is used to determine what scale of angular structures in the phase function can be adequately resolved, and the part of the forward-scattering peak that cannot be resolved is treated as if it was not scattered at all. This works well for Gauss-Legendre quadrature, because its more even spacing of quadrature points in  $\mu$  space ensures that the part of the forward-scattering peak deemed to be 'resolvable' by the delta-M method is indeed adequately resolved. This is less the case for the other quadrature schemes. It is beyond the scope of this paper to see whether modifying the delta-M method would be advantageous for other quadrature schemes.

We stress that the primary purpose of this paper is to explore the appropriate quadrature scheme to use in weather and climate models if 4 or even 6 streams could be afforded, and it is clear from Fig. 5a that Gauss-Jacobi-5 incurs longwave irradiance errors around four times less than Gauss-Legendre, even when clouds are present. But for using DISORT to perform reference calculations, for which 128 streams is common, Gauss-Legendre quadrature should still be used.

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**Figure 5** As Fig. 3 but computed using ERA5 fields at 12 UTC on 11 July 2019, for all-sky conditions with atmospheric scattering. The reference calculations use the most accurate Gaussian scheme with 32 streams: in panels a and b this is Gauss-Legendre quadrature and in panel c it is Gauss-Jacobi-5 quadrature.

Finally, we investigate the spatial pattern of irradiance erors and their dependence on column water vapour, for various uadrature schemes. Figures 6 and 7 show the instantaneous errors in TOA upwelling and surface downwelling irradiance, espectively, for the ERA5 scene considered in this section. The quadrature schemes shown are Gauss-Legendre, Gaussaguerre, Gauss-Jacobi-5 and Optimized for 2-6 streams, except that Elsasser's D = 1.66 scheme has been used in lace of Gauss-Laguerre for the two-stream case because it is hore widely used, and is actually very similar to Gauss-Laguerre (which uses D = 1.6487). In the two-stream case, auss-Legendre (D = 2) performs very poorly with irradiance biases of around 6 W m $^{-2}$ . The other three are much better; Opmized quadrature (D = 1.6402) is a little better than Elsasser, although the errors in different parts of the world tend to be larger than the global-mean bias. The most striking contrast in hese figures is between cloud-free subtropical Africa and the rest of the world. The cloud-free regions are particularly poorly epresented by Gauss-Legendre quadrature, but these tend to be the most accurate regions for the other quadrature schemes in the six-stream case.

The bias (*B*) and error standard deviation (*E*) are shown above each panel of Figs. 6 and 7. To put these values in perspective, when longwave scattering is turned off in the standard ECMWF longwave scheme for this scene, at TOA we have B = 1.3 W m<sup>-2</sup> (very similar to the value reported by Hogan and Bozzo 2016) and E = 1.0 W m<sup>-2</sup>, while at the surface we have B = -0.4 and E = 0.3 W m<sup>-2</sup>. These values are of similar magnitude to the two-stream errors for D = 1.66shown above Figs. 6d and 7d. Many climate models still neglect longwave scattering so these results suggest there would be little point in increasing the number of streams without also turning on longwave scattering. Note that Lacis and Oinas (1991) used a six-stream scheme yet neglected longwave scattering.

Figure 8 depicts the dependence of clear-sky TOA upwelling and surface downwelling irradiances on column water vapour for the same quadrature schemes. While it is again clear that Gauss-Legendre quadrature performs worst up to six streams, the surface irradiance errors of all schemes have a dependence on water vapour that could be important for estimating the impact of changed water vapour on surface temperature. For example, Fig. 8d suggests that as water vapour is increased above 20 kg m<sup>-2</sup>, Elsasser's scheme predicts that the rate at which surface downwelling irradiance increases will be too large by around 0.1 W m<sup>-2</sup> (kg m<sup>-2</sup>)<sup>-1</sup>. Fig. 8e shows that this is reduced to around -0.01 W m<sup>-2</sup> (kg m<sup>-2</sup>)<sup>-1</sup> for the Gauss-Jacobi-5 and Optimized quadratures in the four-stream case.

#### 6 | CONCLUSIONS

The two-stream equations have been the mainstay of atmospheric radiative transfer for many decades, but with ever more resources being found to increase model resolution (e.g. Satoh et al. 2019), now is the time to investigate the potential benefits of increasing the number of streams in weather and climate models. The radiative transfer problem involves two integrals over the cosine of zenith angle,  $\mu$ , but the integral to obtain irradiances from radiances is weighted by  $\mu$ , whereas the integral representing scattering from one stream to another



Figure 6 Errors in computed top-of-atmosphere upwelling irradiance using ERA5 fields at 12 UTC on 11 July 2019, for (rows) bur different quadrature schemes, using (columns) 2, 4 and 6 streams, in units of W m<sup>-2</sup>. The numbers of the form  $B \pm E$  above each panel indicate the global-mean bias (B) and standard deviation of the error (E). Note the different colour scale for each olumn, and also the much wider colour scale for panel a.

not. This means that no single set of  $\mu$  values and corresponding weights (i.e. the quadrature scheme) is optimal for both integrals; in fact, the best performing scheme depends on how much scattering is present.

By far the most widely used quadrature scheme, and indeed he only one available in DISORT, is Gauss-Legendre applied separately in each hemisphere (also known as 'double-Gauss'; Sykes 1951); this is optimal for the scattering integral and so is well suited for shortwave problems. In this paper the accuracy and convergence rate of longwave radiative transfer calculations has been rigorously evaluated for a range of quadrature schemes and numbers of streams, making use of our modified version of DISORT that supports alternative quadratures. Because of the much weaker scattering in the longwave, we find that Gauss-Legendre is the *least* accurate of all quadratures tested for two- and four-stream radiative transfer. In addition to testing several Gaussian quadrature schemes, we have developed three 'Optimized' schemes whereby the angles and weights are chosen to minimize the irradiance and heating-rate errors in a set of clear-sky training profiles. The results are summarized as follows:

• In the two-stream case, the Elsasser (1942) diffusivity of D = 1.66 (corresponding to a zenith angle of  $\theta_1 =$ 52.96°) is close to optimal, although the 'Optimized' value in this paper of  $D = 1.6402 (\theta_1 = 52.43^\circ)$  is a little better for all-sky irradiances, reducing RMSE by around 20% and the magnitude of global-mean biases by 0.3–0.4 W m<sup>-2</sup>. By contrast, the Gauss-Legendre value of D = 2 ( $\theta_1 = 60^\circ$ ) leads to an RMSE six times larger.



Figure 7 As Fig. 6 but for surface downwelling irradiance.

- The Elsasser and Optimized two-stream schemes incur a stratospheric/mesospheric heating-rate bias of 0.25- $0.3 \text{ K} \text{ d}^{-1}$  peaking at the stratopause. This is half the magnitude and of the opposite sign to the bias resulting from D = 2. Unfortunately, selecting a D to minimize the upper-atmosphere heating-rate bias would introduce a significant bias in irradiances, a limitation of the twostream approach.
- For clear-sky calculations in which scattering can be neglected, Optimized quadrature out-performs all others examined in this paper for any number of streams, and is between one and two orders of magnitude more accurate than Gauss-Legendre. 'Gauss-Jacobi-5' quadrature comes a close second.
- For all-sky calculations with 4-8 streams, Gauss-Jacobi-5 quadrature performs best, and is recommended for a weather or climate model looking to increase the number of streams it uses to more than two. For ten or more

streams, Gauss-Legendre quadrature is most accurate, except in the stratosphere and mesosphere.

- Following the idea of Lacis and Oinas (1991), we have proposed quadrature schemes for 4-8 streams that are more efficient via the use of  $\mu$  values that are integer multiples of each other, reducing the number of exponentials that need to be computed. Thus, 'Optimized-IR' is a more efficient but slightly less accurate version of 'Optimized' quadrature and 'Optimized-IRJP' is the same but for Gauss-Jacobi-5.
- We have demonstrated empirically that the quadrature scheme advocated by Li (2000) and which he referred to as Gaussian quadrature with a moment power of infinity (Gauss-Jacobi-∞ in our nomenclature) produces identical angles and weights to Gauss-Laguerre quadrature, although in terms of accuracy this is generally inferior to Gauss-Jacobi-5.

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**igure 8** Errors in clear-sky (top row) top-of-atmosphere upwelling irradiance and (bottom row) surface downwelling irradiance for the ERA5 scene, as a function of column water vapour, for different quadrature schemes with 2, 4 and 6 streams.

The next step will be to implement some of the quadrature choines proposed here in an atmospheric model and test the impact on weather forecasts and on model climate. Two opmizations will be particularly valuable: first, the method of Fu et al. (1997) in which a fast two-stream calculation is used o get an initial estimate of the irradiance profile, followed by ne projection of N radiances through the atmosphere using the two-stream fluxes as the scattering source function, yielding nore accurate irradiances. Second, the number of exponential calculations can be reduced via use of the Optimized-IR and Dptimized-IRJP quadrature schemes whose  $\mu$  values are integer multiples of each other. Preliminary simulations with the ECMWF model suggest that improved longwave quadrature reduces mean stratospheric temperatures by up to 2 K, as well as changing temperature patterns in the troposphere.

This paper has also presented evidence that two-stream schemes may not correctly capture the dependence of surface fluxes on column water vapour, but the magnitude of the effect is reduced by around a factor of 10 when moving to four streams. Further simulations would be required to determine the impact on estimates of the water vapour feedback.

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#### **CONFLICT OF INTEREST**

The author has no conflict of interest to declare.

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