

North Atlantic response to observed North Atlantic oscillation surface heat flux in three climate models

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1 **North Atlantic Response to Observed North Atlantic Oscillation Surface Heat**
2 **flux in Three Climate Models**

3

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10

11 ABSTRACT

12 We investigate how the ocean responds to 10-year persistent surface heat flux forcing over
13 the subpolar North Atlantic (SPNA) associated with the observed winter NAO in three CMIP6-
14 class coupled models. The experiments reveal a broadly consistent ocean response to the
15 imposed NAO forcing. Positive NAO forcing produces anomalously dense water masses in the
16 SPNA, increasing the southward lower (denser) limb of the Atlantic meridional overturning
17 circulation (AMOC) in density coordinates. The southward propagation of the anomalous dense
18 water generates a zonal pressure gradient overlying the models' North Atlantic Current that
19 enhances the upper (lighter) limb of the density-space AMOC, increasing the heat and salt
20 transport into the SPNA. However, the amplitude of the thermohaline process response differs
21 substantially between the models. Intriguingly, the anomalous dense-water formation is not
22 primarily driven directly by the imposed flux anomalies, but rather dominated by changes in
23 isopycnal outcropping area and associated changes in surface water mass transformation (WMT)
24 due to the background surface heat fluxes. The forcing initially alters the outcropping area in
25 dense-water formation regions, but WMT due to the background surface heat fluxes through
26 anomalous outcropping area decisively controls the total dense-water formation response and can
27 explain the inter-model amplitude difference. Our study suggests that coupled models can
28 simulate consistent mechanisms and spatial patterns of decadal SPNA variability when forced
29 with the same anomalous buoyancy fluxes, but the amplitude of the response depends on the
30 background states of the models.

31 **1. Introduction**

32 The subpolar North Atlantic (SPNA) is an intriguing region as various processes induce
33 fluctuations in upper ocean properties on broad timescales with significant implications for
34 climate (Chafik et al. 2016; Zhang et al. 2019). Lying directly below the energetic North Atlantic
35 jet stream, the SPNA exhibits SST variability on a wide spectrum of timescales due to turbulent
36 heat exchanges and Ekman transport associated with the jet's fluctuations (Visbeck et al. 2003;
37 Deser et al. 2010). In the western basin and along its northern periphery, weak stratification and
38 harsh winter conditions allow for deep and intermediate water formation, promoting vertical
39 mixing of heat, salt, and trace gases (e.g., CO₂; Marshall and Schott 1999; Rhein et al. 2017).

40 Observed SPNA SST also exhibits strong variability on decadal to multidecadal (simply decadal
41 hereafter) timescales, as a part of the variability in the wider North Atlantic, often called Atlantic
42 Multidecadal Variability (AMV; Kerr 2000). This decadal temperature variability in the SPNA is
43 widely believed to be primarily generated by anomalous heat transport convergence due to ocean
44 dynamics (Robson et al. 2012; Zhang et al. 2016; Moat et al. 2019; Kim et al. 2020a),
45 particularly by the buoyancy-driven components of the Atlantic meridional overturning
46 circulation (AMOC) and subpolar gyre circulation (Yeager 2020).

47 Mechanisms of decadal upper ocean temperature (UOT) variability in the SPNA are of
48 particular interest for the decadal prediction research community, given high decadal
49 predictability consistently found from initialized hindcasts (Kirtman et al. 2013; Robson et al.
50 2014; Yeager and Robson 2017; Robson et al. 2018; Smith et al. 2019; Borchert et al. 2019;
51 Yeager 2020). Decadal SPNA SST variability is often related to variations in surface temperature
52 and precipitation in the surrounding continents on these timescales (Årthun et al. 2017; Simpson
53 et al. 2019; Kim et al. 2020a), including impactful extreme events (e.g., heat waves; Borchert et
54 al. 2019; Qasmi et al. 2021). It is also ascribed as a source of decadal climate variability in
55 remote regions, including Arctic sea-ice extent in the Atlantic sector (Årthun et al. 2017; Yeager
56 et al. 2015), Sahel precipitation (Dunstone et al. 2011; Kim et al. 2020a), and hurricane activity
57 (Smith et al. 2010; Dunstone et al. 2011; Kim et al. 2020a). Therefore, improving predictions of
58 how the SPNA will evolve (based on solid mechanistic understanding) will have many important
59 implications. However, despite a long history of research, the mechanisms of decadal SPNA SST
60 variability remain an active area of research and debate (Clement et al. 2015; Zhang et al. 2016;
61 O'Reilly et al. 2016; Robson et al. 2016; Kim et al. 2018b; Josey et al. 2018; Oltmanns et al.
62 2020).

63 The leading hypothesis to explain decadal SPNA variability is a mechanistic link between
64 anomalous surface buoyancy fluxes associated with the NAO during boreal winter, anomalous
65 dense-water formation in the SPNA, and subsequent AMOC adjustment (Eden and Willebrand
66 2001; Robson et al. 2012; Yeager and Danabasoglu 2014; Yeager 2020; Kim et al. 2020b). A
67 strong AMOC enhances the northward heat and salt transport into the SPNA from the subtropics,
68 resulting in upper ocean heat and salt content gain in the SPNA, whereas a weaker AMOC does
69 the opposite. The NAO can also directly generate SST variability through changes in the strength

70 of the westerlies on decadal timescales (Deser et al. 2010; Clement et al. 2015; Barrier et al.
71 2015). However, model simulations suggest that this direct atmospheric-driven SST anomaly is
72 overwhelmed by a delayed change in ocean heat transport (Lohmann et al. 2009; Delworth et al.
73 2017). Consistent with this, the decadal SPNA temperature variability lags the NAO by about a
74 decade in observations (W. Kim et al. 2018a; H. Kim et al. 2023), implying that the dominant
75 influence of NAO on decadal SPNA SST variability involves an ocean circulation change.

76 The impact of NAO-related buoyancy forcing on the ocean has been extensively investigated
77 using numerical simulations. When forced with either observation-based surface forcing or
78 idealized NAO forcing, ocean-only model simulations appear to show a robust response of the
79 aforementioned mechanistic link (Eden and Willebrand 2001; Böning et al. 2006; Biastoch et al.
80 2008; Lohmann et al. 2009; Robson et al. 2012; Yeager and Danabasoglu 2014; Polo et al.
81 2014). Systematic comparisons across multiple models forced with the same forcing based on an
82 atmospheric reanalysis reveal a consistent statistical relationship between the NAO and AMOC
83 at subpolar latitudes (Danabasoglu et al. 2016; Xu et al. 2019), suggesting that when the NAO is
84 realistic, the proposed oceanic mechanism seems to be at work in most models.

85 In coupled models, by contrast, the NAO-AMOC link is generally weaker and less robust
86 across models (Ba et al. 2014; Xu et al. 2019; Kim et al. 2023). Xu et al. (2019) compares the
87 statistical relationship between the NAO and AMOC from both forced ocean and coupled
88 configurations that use the same ocean component. They show that in coupled configurations,
89 the relationship is less robust while the same ocean models exhibit a more robust relationship
90 when the models are forced with observed forcing. Many factors can contribute to this weak
91 linkage in coupled models. For example, AMOC variability appears to be more sensitive in some
92 models to freshwater flux from the Arctic Ocean rather than local surface buoyancy fluxes in
93 deep-water formation regions (e.g., Jungclaus et al. 2005; Frankcombe et al. 2010; Lai et al.
94 2022). Different background states in the ocean can also change the efficacy of NAO-related
95 buoyancy forcing for driving AMOC variability. Kim et al. (2023) show using pre-industrial
96 control simulations from CMIP6 that the strength of the NAO-AMOC relationship is
97 significantly correlated with the mean SPNA stratification across the models, which is in turn
98 related to sea-ice extent in the SPNA that can prevent heat loss from the ocean.

99 The diversity of NAO-related surface buoyancy fluxes in coupled models is another factor
100 that could contribute to the wide range of simulated connections between the NAO and decadal
101 SPNA variability. Even if the pattern of the NAO based on SLP is reasonable in coupled models
102 (Wang et al. 2017; Fasullo et al. 2020), the associated buoyancy fluxes may not necessarily be
103 realistic. Turbulent surface heat fluxes, the dominant component of NAO-related surface
104 buoyancy fluxes, in the western SPNA are strongly controlled by air temperature (Kim et al.
105 2016), which is in turn controlled by the strength of westerlies that carry cold air from Canadian
106 Arctic that are enhanced during positive NAO. Therefore, an air temperature bias in these
107 upstream regions or a displacement of the meridional pressure gradient can degrade the realism
108 of simulated heat fluxes in the western SPNA. Also, it has been shown that simulated NAO in
109 coupled models exhibits weaker decadal variability than observed (Wang et al. 2017; Kim et al.
110 2018a; Simpson et al. 2018). Given the importance of persistent SPNA buoyancy forcing in
111 spinning up AMOC (Delworth and Zeng 2016; Kim et al. 2020a; MacGilchrist et al. 2021), the
112 weak decadal NAO variability simulated in models could also contribute to the weak connection.

113 To better understand the diverse NAO-AMOC relationship in coupled models, in this study,
114 we impose surface heat flux forcing associated with the observed winter NAO over the SPNA
115 for 10 years consistently in three CMIP6-class coupled models. By constraining the strength and
116 duration of the NAO-related surface heat flux forcing based on observations, we can remove
117 differences related to the realism of the NAO-related heat flux anomalies and focus on
118 differences in the response to the identical NAO forcing, such as those arising from different
119 background states. These numerical experiments are similar to those performed by previous
120 studies (Delworth and Zeng 2016; Delworth et al. 2016, 2017; Kim et al. 2020b). Delworth and
121 coauthors apply observation-based NAO-related surface heat flux over the North Atlantic in
122 GFDL coupled models in a series of studies and show that the NAO forcing induces expected
123 AMOC and North Atlantic responses with far-reaching climate impacts (Delworth and Zeng
124 2016; Delworth et al. 2016, 2017). Kim et al. (2020) impose the same NAO surface heat flux
125 forcing in the CESM1, but only in the Labrador Sea (LS) to perturb water-mass formation there
126 and examine how the rest of the ocean and climate respond to this perturbation. They find many
127 of the previously reported ocean and climate responses (e.g., changes in AMOC, SPNA UOT,
128 and European surface climate) that are thought to be related to AMV. However, this

129 protocol/approach has not been applied consistently in a multi-model framework to explore the
130 response systematically to the observed NAO surface buoyancy forcing on decadal timescales
131 under different mean background states.

132 To examine the relationship of the imposed forcing to AMOC changes, we adopt the widely
133 used water mass transformation (WMT) analysis framework (e.g., Walin 1982; Speer and
134 Tziperman 1992; Grist et al. 2014; Petit et al. 2021; Yeager et al. 2021) that estimates the volume
135 flux of waters transformed from one density-class to another by surface buoyancy fluxes.
136 Previous studies have shown that surface WMT reasonably captures the mean and decadal
137 variability of AMOC in density coordinates (Grist et al. 2009; Josey et al. 2009).

138 In the next section (Section 2), we briefly introduce the three coupled models used in the
139 present study, describe the experimental design, and explain how WMT is calculated in practice.
140 In Section 3, we present the results of the study. Starting with a brief description of the
141 background state of key variables in each model (Section 3a), we present the responses of
142 surface density fluxes (Section 3b), WMT (Section 3c), and AMOC (Section 3d), followed by
143 wider climate impacts (Section 3e). The salient findings of the study are highlighted in Section 4
144 with some concluding remarks.

145 **2. Experimental design and methods**

146 *a. Models*

147 CESM2 is the latest version of CESM used for CMIP6 simulations (Danabasoglu et al.
148 2020). CESM2 consists of POP2, CAM6, CICE5, and CLM5 for the ocean, atmosphere, sea-ice,
149 and land components, respectively, with a nominal 1° horizontal resolution for all components.
150 Here, we briefly describe a few fundamental features of POP2 since the present study is mostly
151 focused on the ocean. We refer to Danabasoglu et al. (2020) and references therein for a detailed
152 description for each component model that includes updated features from CESM1. POP2 (as
153 well as CICE5) uses a dipole grid with the North Pole displaced over Greenland, allowing for
154 higher horizontal resolution around Greenland (30-50 km). The horizontal resolution also
155 increases to 0.27° near the Equator. It has 60 vertical levels with layer thickness monotonically
156 increasing from 10 m in the upper ocean to 250 m in the deep ocean. POP2 exchanges fluxes

157 with CAM6 and CICE5, calculated using the bulk formulae described in Large and Yeager
158 (2009).

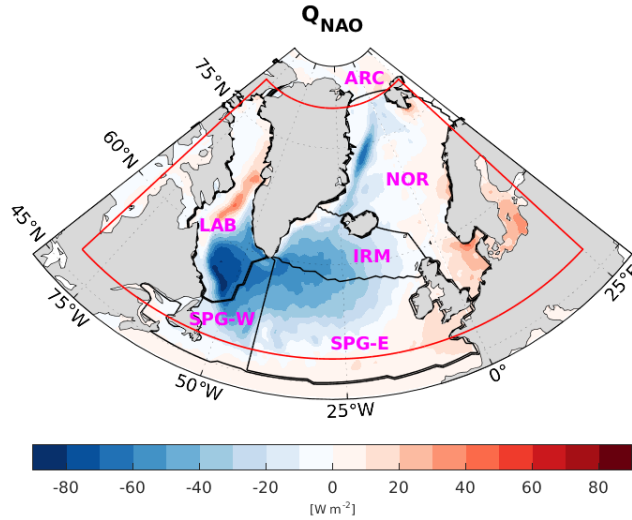
159 EC-Earth3P (Haarsma et al. 2020) is a coupled climate model consisting of an atmospheric
160 component based on the cycle 36r4 of the Integrated Forecast System (IFS) atmosphere-land-
161 wave model of ECMWF coupled to NEMO (v3.6). The H-TESSSEL model is used for the land
162 surface and is an integral part of IFS: for more details see Hazeleger and Bintanja (2012). The
163 atmosphere and ocean/sea ice components are coupled through the OASIS coupler. The ice
164 model, embedded in NEMO, is the Louvain la Neuve sea-ice model version 3 (LIM3), which is a
165 dynamic-thermodynamic sea-ice model with 5 thickness categories. EC-Earth3P uses a reduced
166 Gaussian-grid with 91 vertical levels and a T255 horizontal truncation / N128 grid resolution (~
167 100 km) for the IFS atmosphere. The NEMO ocean has 75 vertical levels and a horizontal
168 resolution of about 1° , reducing to $1/3^\circ$ in the tropics.

169 HadGEM3-GC3.1-LL (GC3.1-LL) is the low-resolution version of the HadGEM3 coupled
170 climate model used for CMIP6 simulations (Kuhlbrodt et al. 2018). The atmospheric component
171 is the Unified Model GA7.1 configuration at N96 horizontal resolution (which equates to
172 ~135 km in the extratropics) with 85 vertical levels up to a model lid at 85 km (35 levels are
173 above 18 km). The ocean component is based on the NEMO ocean model in the GO6.0
174 configuration at 1° resolution with 75 levels. The CICE model is used for sea ice (GSI8.1) and
175 land surface processes are represented using the JULES model (GL7.0). More details about the
176 development of the GC3.1-LL is given in Kuhlbrodt et al. (2018). It is worth noting that the
177 configuration of NEMO used in GC3.1-LL is very close to that used in EC-Earth3P other than a
178 few parameters related to horizontal mixing and turbulent kinetic energy parameterizations.

179 *b. Experimental design*

180 We impose surface heat flux anomalies equivalent to 2 standard deviations of the observed
181 winter (December to March; DJFM) NAO in the ocean component of each model over the
182 SPNA. This amplitude of the NAO heat flux forcing is larger than the decadal averaged
183 amplitude in observations, which is only about 1.2 standard deviation during the decades from
184 the early 1960s to the mid-1990s when there were large multidecadal changes. This amplitude
185 choice was made to obtain clear response signals, particularly in the atmosphere where the

186 signal-to-noise ratio is unrealistically low, possibly due to the low resolution of the models
 187 (Scaife et al. 2019). The surface heat flux anomalies are derived from ERA5 (Hersbach et al.
 188 2020). Specifically, the anomalous forcing is obtained by regressing the anomalous DJFM ERA5
 189 total (turbulent plus radiation) surface heat fluxes onto the station-based DJFM NAO index
 190 (Hurrell 1995), without applying any temporal smoothing, using the data from 1979 through
 191 2018 (Fig. 1).



192

193 Fig. 1. Winter (December through March) NAO-related heat flux anomaly (positive into the ocean)
 194 imposed in the models. The domain outlined by the red line indicate the region where the forcing is applied
 195 with full strength and the outside of the red line (to the outer line) is a transition zone. The black lines indicate
 196 the domains where water-mass transformation is computed with the domain names indicated by pink text
 197 where SPG-W and SPG-E are the western and eastern subpolar gyre, respectively; LAB is the Labrador Sea;
 198 IRM is the Irminger Sea; NOR is the Nordic Seas; ARC is the Arctic Ocean.

199 This anomalous heat flux forcing (Q_{NAO}) is added to the net heat flux passing from the
 200 coupler to the ocean component (Q_c) at each timestep as follows:

$$201 \quad Q_o = Q_c + Q_{NAO}^{eff}, \quad (1)$$

202 where Q_o is the net heat flux into the ocean component and Q_{NAO}^{eff} is the effective NAO heat flux
 203 forcing received by the ocean:

$$204 \quad Q_{NAO}^{eff} = Q_{NAO} \times (1 - a_i) \times W_t(t) \times W_A(x, y), \quad (2)$$

205 , a_i is the sea-ice fraction simulated by the model. $W_t(t)$ is the temporal weight of the forcing,
 206 set to 1 for mid-December through mid-March, with a linear transition from mid-November and

207 to mid-April (zero otherwise). $W_A(x, y)$ is the spatial weight of the forcing, set to 1 within the
 208 SPNA region bounded by 48°-80°N, 80°-25°E (solid red line in Fig. 1), with a 5° linear
 209 transition to zero on each side of the forcing domain (the border of Fig. 1). We stress that only
 210 the ocean component is perturbed by the forcing while other components exchange fluxes
 211 without any constraint. That being said, heat is not conserved within the coupled system when
 212 the forcing is applied.

213 Two parallel coupled ensembles have been conducted, corresponding to positive and
 214 negative NAO forcing. The ensemble size is either 20 (CESM2 and GC3.1-LL) or 25 (EC-
 215 Earth3P). The forcing is applied for the first 10 winters and the simulations continue for another
 216 10 (GC3.1-LL) or 20 (CESM2 and EC-Earth3P) years. All simulations are initialized on January
 217 1 and the forcing is switched on as soon as the run starts. The forcing of the first winter is
 218 therefore only imposed over the months of January to April and because of this reason, we use
 219 January to March (JFM) winter averages in the following analyses. Initial conditions are taken
 220 from the pre-industrial control simulations of each model and external forcing is fixed at 1850
 221 conditions during the experiments. We select a “neutral” set of initial conditions by verifying
 222 that the 20 years following the selected initial conditions in the pre-industrial control simulations
 223 do not exhibit significant decadal anomalies or drift in key variables such as AMOC and SPNA
 224 UOT in the ensemble average.

	CESM2	GC3.1-LL	EC-Earth3P
Hor./Ver. Resolution	1°/60	1°/75	1°/75
Sim. Length (years)	30	20	30
Ensemble Size	20	20	25

225 Table. 1. Summary of the horizontal and vertical resolutions of the ocean models, simulation length, and
 226 ensemble size for each \pm NAO experiment. The vertical resolution shown in the first row is the number of
 227 layers. The simulation length includes the first 10 years with the forcing.

228 *c. Computing the surface Water Mass Transformation and surface Water mass formation*

229 The volume of water being transformed at given density classes to another classes by surface
 230 density fluxes (i.e., Surface WMT) is computed using a widely used method (e.g., Speer and

231 Tziperman 1992; Grist et al. 2014; Petit et al. 2021) based on a pioneering work by Walin
 232 (1982). Specifically, we adopt the details used in Yeager et al. (2021), including the equation of
 233 state (McDougall et al. 2003), the number and range of density layers, and regions where surface
 234 density fluxes (SDF) along the outcropping density layers is integrated to compute WMT.

235 SDF (in units of $\text{kg m}^{-2} \text{s}^{-1}$) is calculated from the monthly net surface heat (Q_o) and
 236 freshwater (F_o) fluxes (positive into the ocean for both fluxes) from the models as follows:

$$237 \quad \text{SDF} = -\frac{\alpha}{C_p} Q_o - \beta \frac{S}{1-S} F_o, \quad (3)$$

238 where α is the thermal expansion coefficient, C_p is the specific heat capacity of seawater, β is the
 239 haline contraction coefficient, and S is the sea surface salinity. α and β are computed using the
 240 non-linear equation of state from McDougall et al. (2003) as mentioned above.

241 To obtain WMT (in $\text{m}^3 \text{s}^{-1} \equiv 10^{-6} \text{ Sv}$) as a function of density, the SDF is integrated along
 242 surface density outcropping areas (A_ρ) north of 45°N in the Atlantic sector including the Arctic
 243 and Subarctic Oceans and within each domain delineated in Fig. 1:

$$244 \quad \text{WMT}(\rho) = \frac{1}{\Delta\rho} \iint \text{SDF} dA_\rho, \quad (4)$$

245 where $\rho = \sigma_2$ (i.e., density referenced to 2000 m after subtracting 1000 kg m^{-3}) in our
 246 application. $\Delta\rho$ is 0.2 kg m^{-3} for $28 \leq \sigma_2 \leq 35 \text{ kg m}^{-3}$, 0.1 kg m^{-3} for $35 < \sigma_2 \leq 36 \text{ kg m}^{-3}$ and 0.05
 247 kg m^{-3} for $36 < \sigma_2 \leq 38 \text{ kg m}^{-3}$ (86 layers in total). Surface water mass formation (WMF in Sv) is
 248 computed as the convergence of WMT:

$$249 \quad \text{WMF}(\rho) = -\frac{d\text{WMT}}{d\rho} \times d\rho. \quad (5)$$

250 Thus, Eq. (5) quantifies the volume of water masses that is formed or destroyed by WMT at
 251 given density classes.

252 Because Q_o can be decomposed into the Q_C and Q_{NAO}^{eff} terms (Eq. 1), Eq. (3) can be
 253 decomposed into:

$$254 \quad \text{SDF} = -\frac{\alpha}{C_p} (Q_C + Q_{NAO}^{eff}) - \beta \frac{S}{1-S} F_o. \quad (6)$$

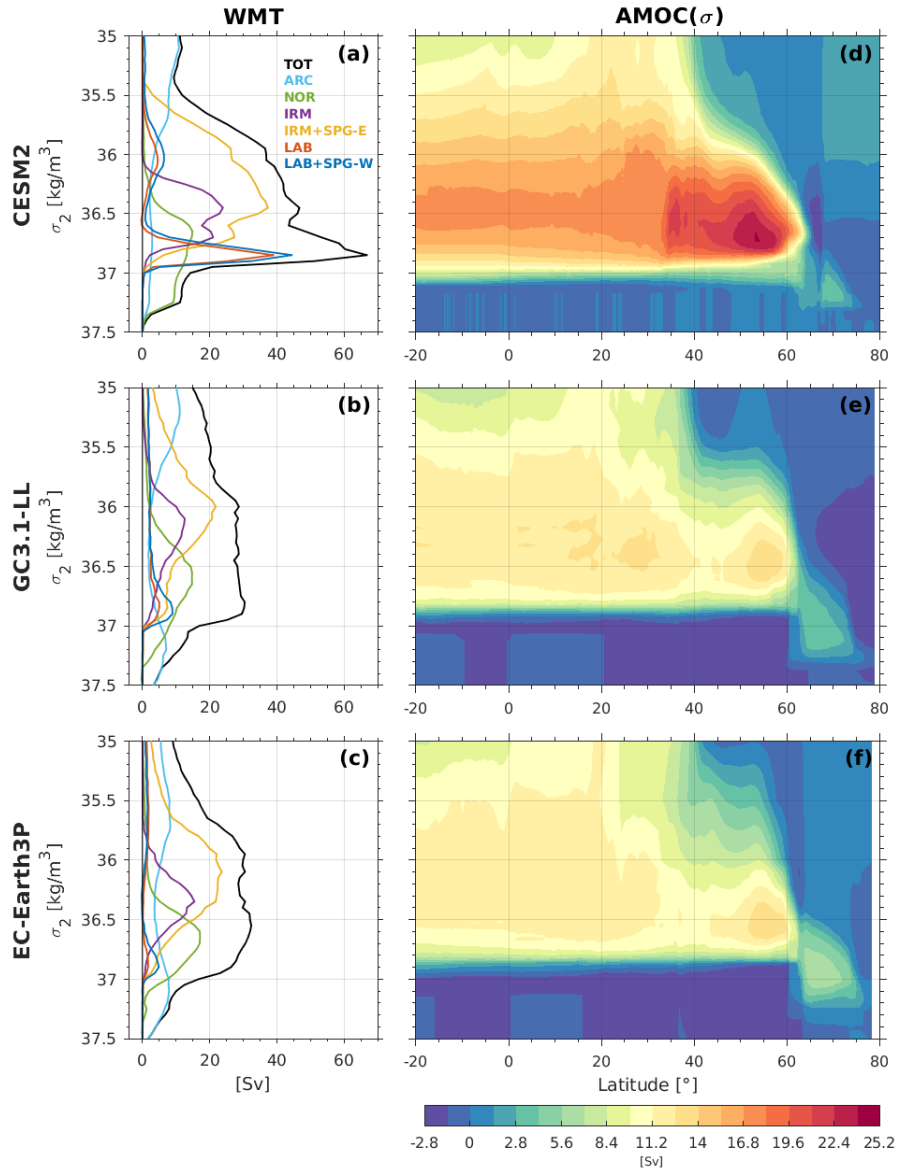
255 We note that all SDF terms can change due to the ocean surface response to imposed NAO
256 forcing, including its effect on α and β , which are function of density. Consequently, the heat
257 flux component of WMT and WMF can also be decomposed into terms related to Q_C , Q_{NAO}^{eff} , and
258 F_o .

259 The ocean components in all three coupled models use depth coordinates, so the overturning
260 streamfunction in density (σ_2) coordinates (hereafter AMOC(σ)) is computed offline in order to
261 be related to WMT. In the following sections, we show annually or seasonally averaged
262 ensemble-mean differences between +NAO and -NAO experiments, which can be interpreted as
263 the linear response to the imposed forcing. The statistical significance of the ensemble-mean
264 difference is assessed at the 95% confidence level using a two-sided Student's t -test with degrees
265 of freedom determined using the Welch-Satterthwaite equation (Welch 1947).

266 **3. Results**

267 *a. Comparison of background states*

268 The present study compares ocean responses to the observed NAO-related heat flux forcing
269 in three climate models focusing on WMT and AMOC, and thus, here we compare essential
270 features of the background state of these variables. We define the background state as a first-year
271 average across both +NAO and -NAO experiments, equivalent to 50-year and 40-year averages
272 for EC-Earth3P, and CESM2 and GC3.1-LL, respectively, because a long pre-industrial control
273 simulation is not available for EC-Earth3P. Although the forcing is active for the first year,
274 responses are generally weak and largely cancel out by averaging across both +NAO and -NAO
275 experiments. As will be shown later, the strength of the WMT response depends on the
276 background state of the surface density and surface heat fluxes. We will show these background
277 states when the WMT response is discussed.

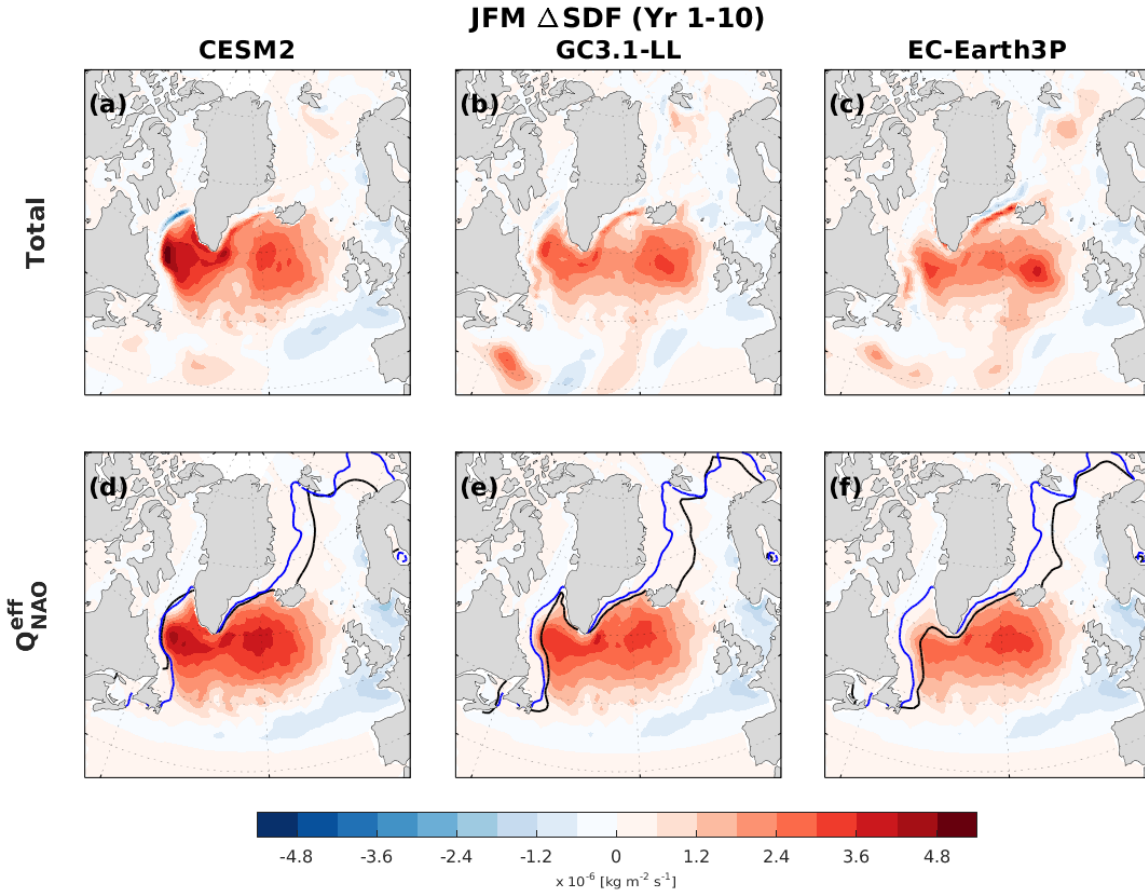


278
 279 Fig. 2. Background state of (a-c) JFM water mass transformation (WMT) and (d-f) annual overturning
 280 streamfunction in density (σ_2) coordinates from CESM2 (top), GC3.1-LL (middle), and EC-Earth3P (bottom).
 281 In (a-c), each line represents WMT in the entire North Atlantic domain north of 45°N (black), ARC (light
 282 blue), NOR (Green), IRM (purple), IRM plus SPG-E (yellow), LAB (red), and LAB plus SPG-W (blue)
 283 domains shown in Fig. 1.

284 Figure 2a-c shows the background state of JFM WMT for the entire Atlantic sector north of
 285 45°N (black line) and subdomains outlined in Fig. 1 (colored lines). All models show enhanced
 286 total WMT at densities (σ_2) roughly between 35.7 and 37 kg m^{-3} , which is most pronounced in
 287 CESM2. In particular, the peak WMT between 36.7 and 36.9 kg m^{-3} is about twice as large as
 288 that of GC3.1-LL and EC-Earth3P. Much of this elevated WMT in CESM2 takes place in the LS

289 in the higher density classes ($36.8\text{-}36.9\text{ kg m}^{-3}$) with a contribution from the Nordic Seas, while
290 WMT in the eastern SPNA (IRM plus SPG-E) contributes much of the total WMT at lower
291 density classes ($36\text{-}36.7\text{ kg m}^{-3}$). Although WMT in the western SPNA (LAB plus SPG-W) is
292 somewhat elevated in GC3.1-LL and EC-Earth3P in the high-density range as in CESM2, its
293 strength is substantially weaker (less than 10 Sv vs. more than 40Sv), especially in EC-Earth3P,
294 likely because of substantially weaker surface heat loss and too extensive sea-ice cover in the LS
295 (see below). As a result, no distinct peak is found for the total WMT in these density classes in
296 these models. Instead, the total WMT in high density classes ($36.5\text{-}36.9\text{ kg m}^{-3}$) is mostly
297 contributed by the Nordic Seas in EC-Earth3P and by multiple regions (both western and eastern
298 SPNA and Nordic Sea) in GC3.1-LL. WMT in the Arctic Ocean is more active in these models
299 in the density classes greater 37 kg m^{-3} , while much of the WMT in these densest classes takes
300 place in the Nordic Seas in CESM2. The annual background WMT shows similar shapes for all
301 models with an amplitude approximately one-third of the JFM mean. The maximum annual
302 surface WMT of $\sim 10\text{ Sv}$ in GC3.1-LL and EC-Earth3P is consistent with observational estimates
303 by Jackson and Petit (2022), while that in CESM2 ($\sim 20\text{ Sv}$) is overestimated.

304 Figure 2d-f shows the background state of AMOC(σ). All three models show a broadly
305 comparable background AMOC(σ) in that most densification of northward flowing waters takes
306 place from the subtropics through the SPNA, which feeds the southward flowing dense-water
307 roughly denser than 36.7 kg m^{-3} in CESM2 and 36.5 kg m^{-3} in GC3.1-LL and EC-Earth3P.
308 However, the maximum overturning strength at subpolar latitudes is substantially stronger in
309 CESM2 ($\sim 25\text{ Sv}$), roughly twice that of other two models ($\sim 13\text{ Sv}$), consistent with the
310 maximum background WMT difference. We also note that a relatively large contribution from
311 Nordic and Arctic Seas seen in WMT in GC3.1-LL and EC-Earth3P is also evident in AMOC(σ)
312 (i.e., the overturning cell north of 60°N at densities greater than 36.5 kg m^{-3}). In comparison to
313 the direct measurements of the AMOC at 26.5°N (RAPID array; Moat et al. 2023), CESM2
314 shows relatively good agreement with the maximum overturning of $\sim 18\text{ Sv}$, compared to $\sim 17\text{ Sv}$
315 in the observations, although the upper (North Atlantic Deep Water) cell is too shallow (Fig. S1).
316 In EC-Earth3P, the upper cell is even shallower, and the maximum overturning strength is too
317 weak ($\sim 14\text{ Sv}$) while GC3.1-LL lies in the middle.

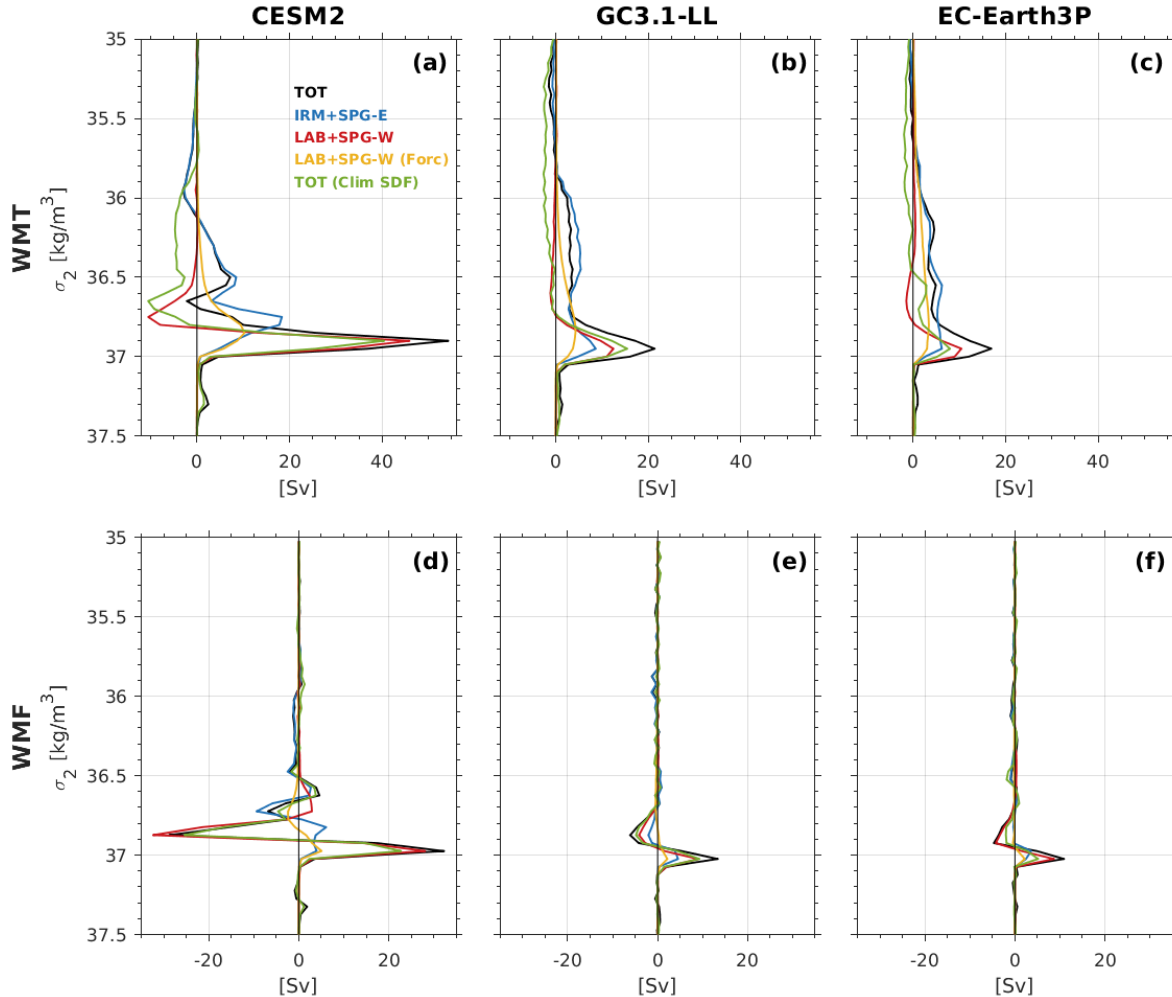


318
 319 Fig. 3. Differences (+NAO minus -NAO experiment) of JFM (January through March) (a-c) total surface
 320 density flux (SDF) and (d-f) SDF associated with the effective NAO forcing (Q_{NAO}^{eff} , the second term on the rhs
 321 of Eq. 6), averaged over the first decade (year 1-10), from CESM2 (left), GC3.1-LL (middle), and EC-Earth3P
 322 (right). The black and blue lines in (e-f) represent the background JFM sea-ice extent (15% ice concentration)
 323 simulated in each model and from satellite-derived estimates (1979-2014 average; Comiso 1999), respectively.

324 *b. Surface density fluxes*

325 Figure 3 shows the differences of total SDF and SDF associated with the imposed heat flux
 326 forcing (i.e., the second term on the rhs of Eq. 6). The total SDF differences (Fig. 3a-c) show a
 327 buoyancy loss over almost entire SPNA (when the +NAO forcing is applied), which is largely
 328 contributed by the imposed forcing (Fig. 3d-f). The contributions coming from the other two
 329 terms on the rhs of Eq. (6), related to Q_C and F_o , whose differences are attributable to feedbacks
 330 in the coupled system in response to the forcing are minor (Fig. S2). In particular, the SDF
 331 differences due to F_o are almost negligible except along the ice edge (Fig. S2d-f). The negative
 332 differences due to Q_C arise because of a cooling driven by the imposed (positive) NAO forcing
 333 and acts to damp moderately the total SDF differences in the central SPNA (Fig. S2a-c).

334 Although the same heat flux forcing is originally used (Q_{NAO}), the ocean component of each
335 model receives slightly different effective heat flux forcing due to different sea-ice conditions
336 (cf. Q_{NAO}^{eff} in Eq. 2), particularly in the LS. The most and least extensive background sea-ice
337 cover in the LS (black lines in Fig. 3d-f) allows for the weakest and strongest heat flux forcing in
338 EC-Earth3P and CESM2, respectively, while GC3.1-LL lies in the middle. In addition, the LS
339 sea-ice cover increases more in GC3.1-LL and EC-Earth3P for the +NAO forcing (while
340 shrinking more for the -NAO forcing), as will be shown later, contributing to a weaker heat flux
341 forcing in these models compared to CESM2. We note that the background sea-ice extent is
342 closest to that of satellite-derived estimates (blue lines) in CESM2, particularly in the LS. The
343 SDF differences associated with Q_{NAO}^{eff} are further affected by background α , which is lowest in
344 EC-Earth3 in the western SPNA, followed by GC3.1-LL and CESM2 (Fig. S3a-c). In addition, α
345 decreases most in GC3.1-LL, followed by EC-Earth3P and CESM2 (Fig. S3d-f), as SST cools in
346 the same order as α in response to the +NAO forcing. Because of these effects of sea-ice extent
347 and α , the total SDF differences in the western SPNA are about 30% larger in CESM2 than in
348 GC3.1-LL, which is in turn about 40% larger than in EC-Earth3P. The differences in the eastern
349 SPNA are relatively small (about 10% larger in CESM2 than GC3.1-LL, which is about 5%
350 larger than in EC-Earth3P).



351
 352 Fig. 4. Differences of (a-c) JFM water mass transformation (WMT) and (d-f) water mass formation
 353 (WMF), averaged over the first decade, from CESM2 (left), GC3.1-LL (middle), and EC-Earth3P (right). The
 354 colored lines represent the differences in the entire North Atlantic domain north of 45°N (black), eastern
 355 SPNA (IRM plus SPG-E; blue), western SPNA (LAB plus SPG-W; red), and due to the imposed forcing in the
 356 western SPNA (yellow). The green lines are same as black lines but computed with the background SDF and
 357 time-varying isopycnal outcropping area.

358 *c. WMT/WMF response*

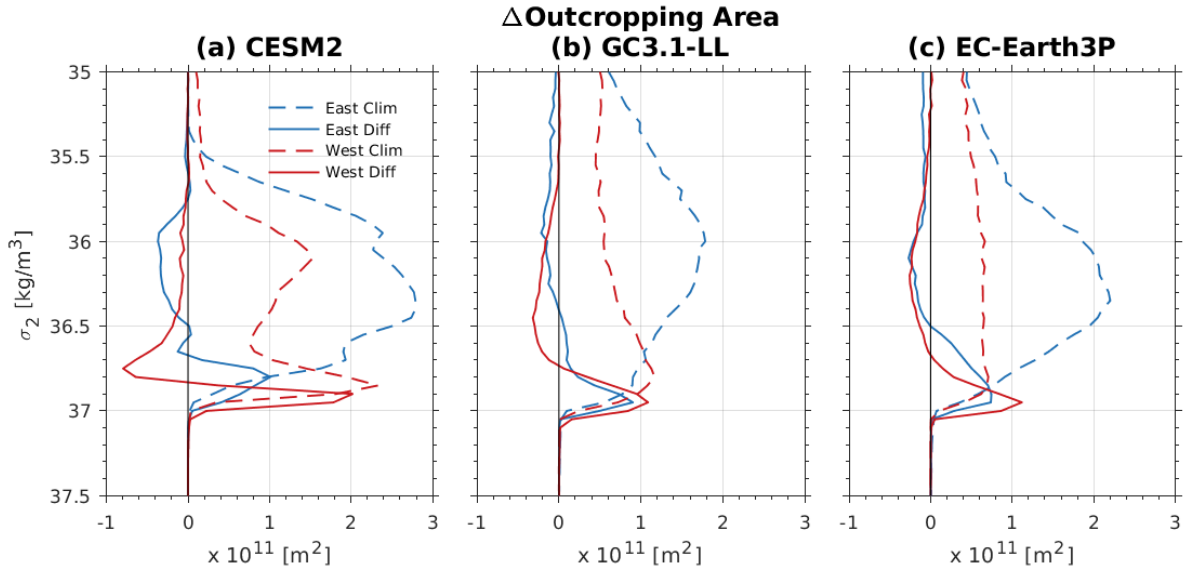
359 Figure 4 shows JFM WMT (a-c) and WMF (d-f) differences, averaged over the first decade
 360 when the forcing is applied, as a function of σ_2 . Integrated over the entire domain north of 45°N,
 361 all models show a strong increase in WMT in the high-density class between 36.8 and 37 kg m⁻³
 362 (black lines in Fig. 4a-c) and weaker (generally positive) WMT change over a broad range of
 363 lighter densities. The density range of the peak response is similar to or denser than that of the
 364 background WMT peak (Fig. 2a-c). As WMF is the convergence of WMT (Eq. 5), the high-

365 density WMT anomaly peak corresponds to a dipole WMF response (Fig. 4d-f). That is, the
366 imposed +NAO forcing produces a densification of high-density water masses that make up the
367 AMOC lower limb. While all models exhibit qualitatively similar WMT/WMF responses, the
368 strength of the peak response is more than twice as large in CESM2 compared to the other
369 models. We will discuss this different response strength later in this subsection.

370 Despite the fact that the primary background high-density WMT takes place in regions other
371 than the western SPNA in GC3.1-LL and EC-Earth3P (Fig. 2), the WMT/WMF peak response is
372 mostly contributed by the western SPNA in all models (red lines in Fig. 4). In GC3.1-LL and
373 EC-Earth3P, there is a relatively large contribution from the eastern SPNA (blue lines), which is
374 the largest contribution for lighter density classes ($\sim 36.0\text{-}36.7 \text{ kg m}^{-3}$) in all models. The
375 importance of the western SPNA is especially clear in CESM2 where the peak response of WMT
376 and the associated dipole WMF response are dominated by the LS (Fig. S4a and d). In contrast,
377 the contributions from the LS, SPG-W, and SPG-E are all comparable in GC3.1-LL and EC-
378 Earth3P (Fig. S4b-c and e-f). Although the maximum peak WMT difference in the western
379 SPNA is roughly three times larger in CESM2 than in the other two models, the total WMT
380 difference integrated over density classes of the respective broad peaks ($> 36.0 \text{ kg m}^{-3}$) is about
381 twice as large in CESM2 compared to GC3.1-LL, which is in turn about 50% larger than in EC-
382 Earth3P. This inter-model difference is much larger than the inter-model SDF difference
383 discussed above can explain. There is also an anomalous WMT/WMF in the Irminger Sea in all
384 models, but it mostly occurs in a lighter σ_2 range ($< 36.8 \text{ kg m}^{-3}$). The WMT/WMF response in
385 the Nordic Sea and the Arctic Ocean is negligible (not shown). The contribution from freshwater
386 flux to the WMT/WMF response is also relatively small, especially in CESM2 (Fig. S4). In all
387 models, but especially in GC3.1-LL and EC-Earth3P, the WMT response due to freshwater flux
388 tends to damp the enhanced WMT in high-density classes ($36.7\text{-}36.9 \text{ kg m}^{-3}$).

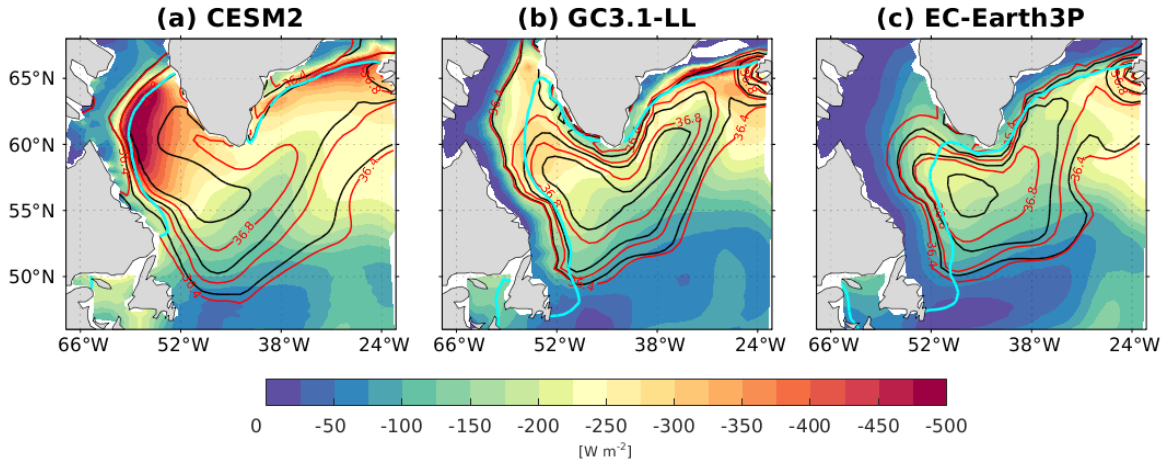
389 Although the WMT (thus WMF) response is ultimately a consequence of the imposed
390 forcing, we find a surprisingly small WMT/WMF response associated with Q_{NAO}^{eff} (yellow lines in
391 Fig. 4), which appears to account for only a small fraction ($< 20\%$) of the peak WMT/WMF
392 response and occurs over lighter density classes than those of the enhanced total WMT/WMF
393 response. This result may appear at odds with the dominance of Q_{NAO}^{eff} in the SDF differences
394 (Fig. 3), but the WMT response also depends on changes in isopycnal outcropping area (A_ρ ; Eq.

395 4). We note that changes in A_ρ are already taken into account in the WMT/WMF response
 396 associated with Q_{NAO}^{eff} . Therefore, much of the WMT/WMF response should result from
 397 interaction between Q_C and A_ρ , although the SDF differences due to Q_C itself are small (Fig. S2).



398
 399 Fig. 5. Differences of JFM isopycnal outcropping area averaged over the first decade (solid) in the western
 400 (red; LAB plus SPG-W) and eastern (blue; IRM plus SPG-E) SPNA. Also, shown are the respective
 401 background state of the isopycnal outcropping area (dashed).

402 Figure 5 shows the JFM A_ρ differences for the first decade in the western (solid red lines)
 403 and eastern (solid blue lines) SPNA along with their background states (dashed lines). All
 404 models reveal the peak response of A_ρ in density classes where the respective peak WMT
 405 response occurs (Fig. 4). A_ρ increase occurs at densities slightly greater than the highest density
 406 of the background outcropping area, at the expense of a decrease at lighter densities, indicating
 407 an expansion of dense A_ρ in the SPNA when the positive NAO heat flux forcing is applied.
 408 Similar to the WMT response, the maximum (positive) A_ρ change in the western SPNA is largest
 409 in CESM2 (roughly twice as large compared to the other two models), but the inter-model
 410 difference is smaller than that of WMT. Therefore, other factors seem to be needed to explain the
 411 inter-model difference of the magnitude of the WMT response. We note that the A_ρ changes take
 412 place over a wider density range in GC3.1-LL and EC-Earth3P, while they are concentrated in
 413 the highest density classes in CESM2, which is also generally consistent with the WMT changes
 414 (Fig. 4a-c).



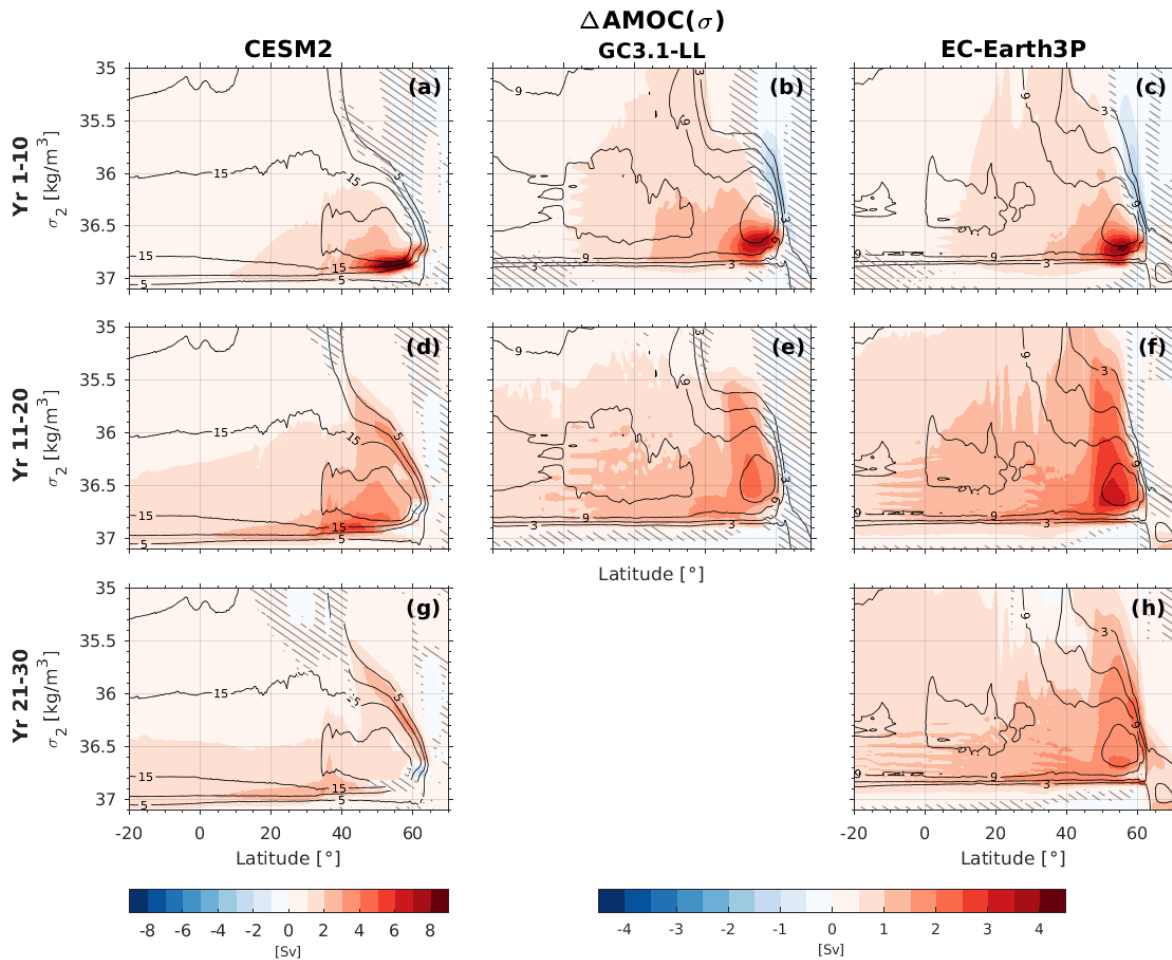
415

416 Fig. 6. JFM surface density (σ_2) from the ensemble mean of +NAO experiments (red contours) and its
 417 background state (black contours), superimposed on the background JFM surface heat fluxes (shading;
 418 negative for heat loss from the ocean), from (a) CESM2, (b) GC3.1-LL, and (c) EC-Earth3P. Also shown is the
 419 JFM sea-ice extent (15%) from the ensemble mean of +NAO experiments (cyan). For σ_2 , only contours greater
 420 than 36.4 kg m^{-3} are shown at intervals of 0.2 kg m^{-3} . Note that the color scale for the surface heat fluxes is
 421 reversed.

422 The expansion of A_ρ towards higher densities implies that these density layers are exposed to
 423 surface heat loss by the background surface heat flux, which is intense in the western SPNA.

424 Figure 6 shows the JFM surface density (σ_2), averaged over the first decade, from the +NAO
 425 experiment (red contours) in the western to central SPNA along with the background surface
 426 density (black contours; densities lower than 36.4 kg m^{-3} are omitted). Also shown are the
 427 background JFM surface heat fluxes (shading) and first-decade average JFM sea-ice edge (15%)
 428 from the +NAO experiment (light blue contours). The expansion of A_ρ of $36.8\text{-}37.0 \text{ kg m}^{-3}$ (note
 429 difference between the black and red contours of 36.8 kg m^{-3}) is most obvious, compared to other
 430 layers, in all models, especially in CESM2, consistent with Fig. 5. This expanded A_ρ coincides
 431 with strong background surface heat fluxes, reaching up to 400 W m^{-2} in the LS around 60°N in
 432 CESM2 but $250\text{-}300 \text{ W m}^{-2}$ in GC3.1-LL and around 200 W m^{-2} in EC-Earth3P. In EC-Earth3P,
 433 moreover, the LS is more covered by sea-ice, thus the background surface heat flux feedback is
 434 less efficient there. Therefore, these results suggest that changes in A_ρ and associated changes in
 435 surface WMT from the background surface heat fluxes are the key processes that determine the
 436 WMT response to the imposed forcing. That is, the WMT response is especially larger in
 437 CESM2 because the expanded A_ρ is exposed to a greater background surface heat loss.

438 To verify the above hypothesis, we repeat the computation of WMT/WMF with background
 439 SDF. That is, A_p only changes in time in this computation. The resultant total WMT/WMF
 440 differences averaged over the first decade are shown as green lines in Fig. 4. The WMT/WMF
 441 with background SDF indeed explains a large portion of the total WMT/WMF: more than 70%
 442 in CESM2 and GC3.1-LL, and about 50% in EC-Earth3P, confirming the key role of changes in
 443 A_p and associated changes in surface WMT from the background surface heat fluxes for the total
 444 WMT/WMF differences. A quantitatively similar conclusion also holds when the western SPNA
 445 is separately considered. We note that the weaker response of WMT/WMF with background
 446 SDF in EC-Earth3P, compared to GC3.1-LL, reflects the weaker background surface heat fluxes
 447 in this model, as changes in outcropping area are comparable between the two models.



448
 449 Fig. 7. Difference of annual overturning streamfunction in density (σ_2) coordinates (AMOC(σ)) averaged
 450 over (a-c) years 1-10, (d-f) years 11-20, and (g-h) years 21-30 from CESM2 (left), GC3.1-LL (middle), and
 451 EC-Earth3P (right). Note that the color scale for GC3.1-LL and EC-Earth3P is half of that for CESM2. The

452 black contours are the climatological AMOC(σ) in each model with contour intervals of 5 (3) Sv for CESM2
453 (GC3.1-LL and EC-Earth3P). The hatched regions indicate that the differences are *not* statistically significant
454 at a 95% confidence level.

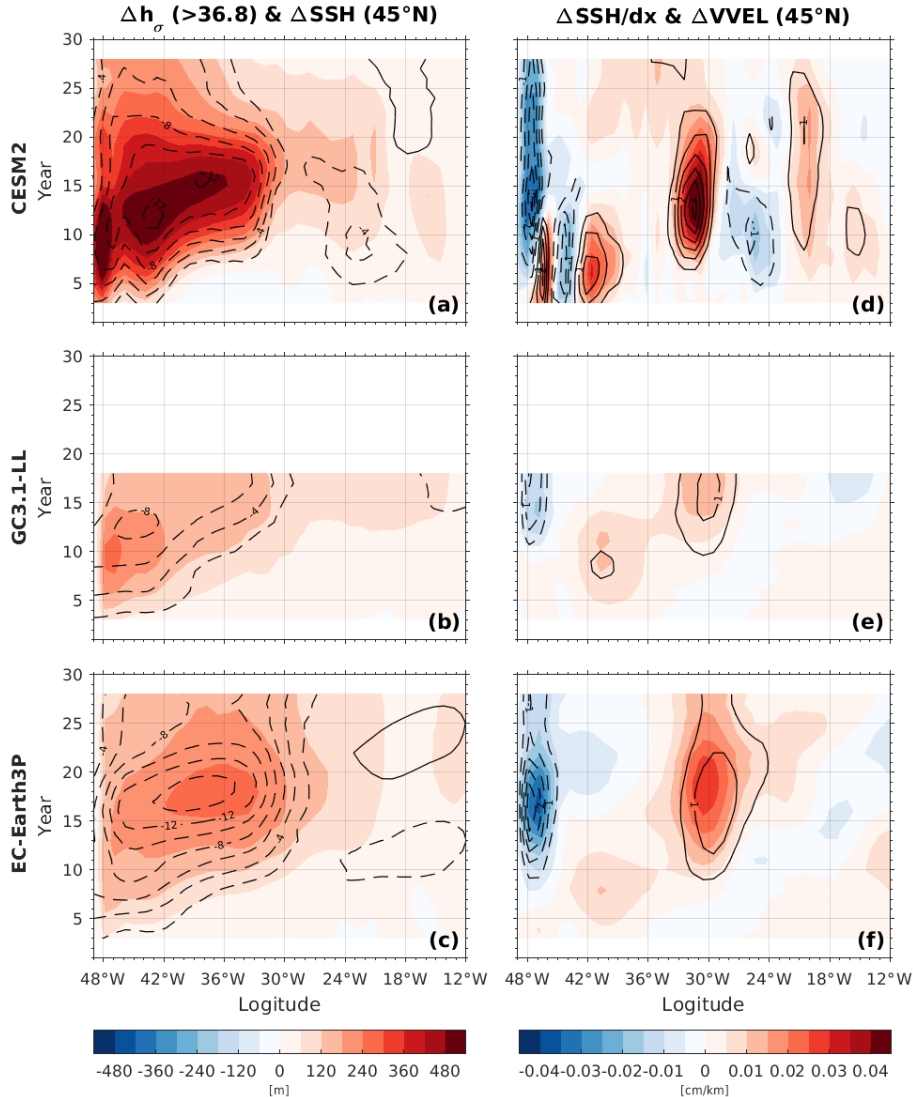
455 *d. AMOC response*

456 Based on the WMT analysis, we would expect the AMOC(σ) response roughly to be twice as
457 large in CESM2 as in GC3.1-LL and EC-Earth3P. Figure 7 shows decadal AMOC(σ) responses
458 superimposed on the background states. In all models, an increase in lower (denser) limb of
459 AMOC(σ) is observed during the first decade at subpolar latitudes (north of 45°N), at densities
460 greater than the density where the background maximum overturning takes place and similar to
461 the density range where anomalous WMT occurs (Fig. 7a-c). The amplitude difference of the
462 maximum lower limb response, which is twice or more as large in CESM2 as in the other two
463 models (~10 Sv in CESM2 vs. ~4 Sv in GC3.1-LL and EC-Earth3P; note the different color
464 scales in Fig. 7), indeed match roughly that of the peak WMT response. In the second decade
465 (years 11-20; Fig. 7d-f), there is an indication of a southward propagation of the overturning
466 anomalies in all three models within the lower limb, seen most prominently in CESM2. In
467 GC3.1-LL and EC-Earth3P, there appears to be a greater communication of the signal to lighter
468 waters during the propagation to the south. Interestingly, the response of the upper (lighter) limb
469 at subpolar latitudes, the northward flow at densities lighter than the density of the background
470 maximum overturning, emerges as the lower limb anomaly propagates to the south in all models
471 (cf. Fig. 7a-c and 7d-f), and persists even after the lower limb anomalies dissipate and move to
472 the subtropics (Fig. 7g-h). Notably, this upper limb response appears as a secondary maximum
473 between densities 36 and 36.5 kg m⁻³ in CESM2 during the second and third decades (Fig. 7d,g)
474 and in EC-Earth3P during the third decade (Fig. 7h).

475 The overturning streamfunction in depth coordinates (AMOC(z)) also shows a consistent
476 pattern of decadal differences across the models (Fig. S5) with an anomalous overturning,
477 centered at around 1000 m (slightly deeper in CESM2), propagating from subpolar latitudes
478 (40°-60°N) during the first decade to subtropical latitudes in the later decades. Consistent with
479 the AMOC(σ) response, the amplitude of the AMOC(z) response is also substantially stronger in
480 CESM2 than in other two models (~3.2 Sv vs. ~1.8 in terms of the maximum overturning
481 anomalies). We note that the delayed response of the upper limb at subpolar latitudes seen in

482 AMOC(σ) is not seen in AMOC(z) in all models, consistent with the idea that the delayed signal
483 is a gyre circulation response that becomes visible when overturning is viewed in density space
484 (Yeager 2020; Yeager et al. 2021).

485 Yeager (2020) put forward a mechanism that clarifies the connection between the upper and
486 lower limbs of AMOC(σ) that involves deep, dense-water flow interacting with bottom
487 topography. The Mid-Atlantic Ridge (MAR) acts as a dam for southeastward flowing dense
488 waters formed in the SPNA and causes these dense waters to accumulate along its western flank
489 near the southern boundary of the SPNA. The accumulation of anomalously dense waters in this
490 region generates a corresponding SSH anomaly through the steric effect. The zonal gradient of
491 the SSH anomaly drives an anomalous meridional geostrophic flow that projects onto the upper
492 limb of AMOC(σ) and brings warm subtropical waters into the SPNA. Away from the influence
493 of surface fluxes and stalled by the MAR, this dense water anomaly persists in time and, hence,
494 provides high predictability of the upper limb of AMOC(σ) and UOT in the SPNA. This
495 mechanism has also been identified in a multi-centennial, high-resolution (eddy-rich) coupled
496 simulation using CESM (Yeager et al. 2021).

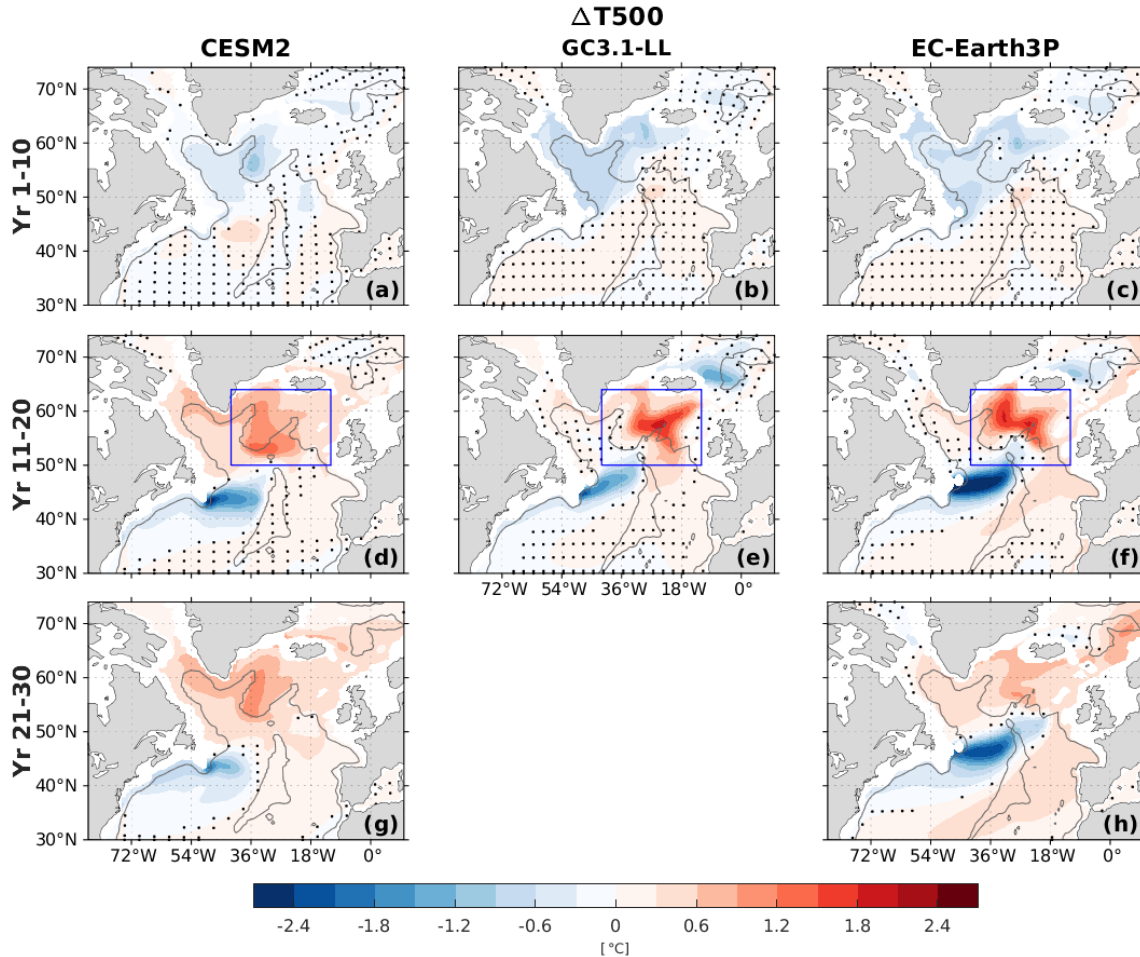


497
 498 Fig. 8. Difference of (a-c) annual dense-water ($\sigma_2 > 36.8 \text{ kg m}^{-3}$) thickness (shading) and SSH (contours),
 499 and (d-f) zonal SSH gradient (shading) and meridional velocity averaged over the upper 700 m (contours), as a
 500 function of longitude (x-axis) and time (y-axis), along the models' zonal grid line closest to 45°N from
 501 CESM2 (top), GC3.1-LL (middle), and EC-Earth3P (bottom). The contour intervals are 2 cm in (a-c) and 0.5
 502 cm s⁻¹ in (d-f) with negative contours dashed and the zero contour omitted. 48°W is roughly where the
 503 continental break is located at the latitude. A 5-year running average is applied to all time series.

504 In line with this proposed mechanism, we find a similar propagation and accumulation of
 505 dense-water thickness anomalies around the MAR in all models, with patterns that are consistent
 506 with concurrent SSH anomalies (Fig. S6). In Fig. 8a-c, we show dense-water thickness ($\sigma_2 >$
 507 36.8 kg m^{-3} ; shading) differences together with SSH differences (contours) along the models'
 508 zonal grid line closest to 45°N (roughly the boundary between the subtropical and subpolar
 509 gyres). We refer to this dense-water thickness as LS Water (LSW) thickness since much of it is

510 generated in the LS in all models. The LSW thickness anomalies gradually increase but remain
511 largely confined west of 30°W (i.e., west of the MAR) in all models, peaking at years between
512 10 and 20 with the largest thickness anomaly in CESM2 (up to 500 m). The patterns of the
513 overlying SSH anomalies essentially mirror those of the LSW thickness anomalies. Figure 8d-f
514 shows anomalous zonal SSH gradient (shading) and meridional velocity averaged over upper
515 700 m (contours). The anomalous SSH gradient is largest around 30°W in all models, which is
516 largest in CESM2 and weakest in GC3.1-LL. Through geostrophy, the positive zonal SSH
517 gradient induces an anomalous northward flow that almost perfectly overlies the zonal SSH
518 gradient anomaly in all models. 30°W is where the major branch of the models' North Atlantic
519 Current (NAC) is located (Fig. S7). Thus, the anomalous northward flow can be interpreted as a
520 strengthened NAC.

521 In CESM2, the NAC anomaly peaks between years 10 and 15 and weakens by years 20-25.
522 However, together with the secondary northward flow anomaly east of the MAR near 20°W, the
523 total northward flow persists through the end of the simulations. In EC-Earth3P, the NAC
524 anomaly develops around year 10, maximizes around year 18, and persists through the end of the
525 simulations. There is also an anomalous meridional flow of opposite sign along the western
526 boundary in all models, which cancels, to a large extent, the NAC anomaly. This cancelation is
527 likely the reason why there is no delayed upper AMOC(z) response (Fig. S5). However, because
528 these two anomalous meridional flows of opposite sign carry different density classes (i.e.,
529 relatively dense subpolar water along the western boundary and relatively light subtropical
530 waters by the NAC; Fig. S8), AMOC(σ) reveals an anomalous overturning in its upper limb in
531 the later years (Fig. 7). In GC3.1-LL, the NAC anomaly develops after year 10, similar to other
532 two models. However, another northward flow anomaly develops around 42°W earlier in the
533 simulations. A similar anomalous flow also presents at a similar location in other models, but the
534 cancelation by the opposite flow near the western boundary is weak in GC3.1-LL. This appears
535 to be the reason for the earlier spinup of the upper AMOC(σ) limb in GC3.1-LL than in the other
536 two models, as will be shown later.

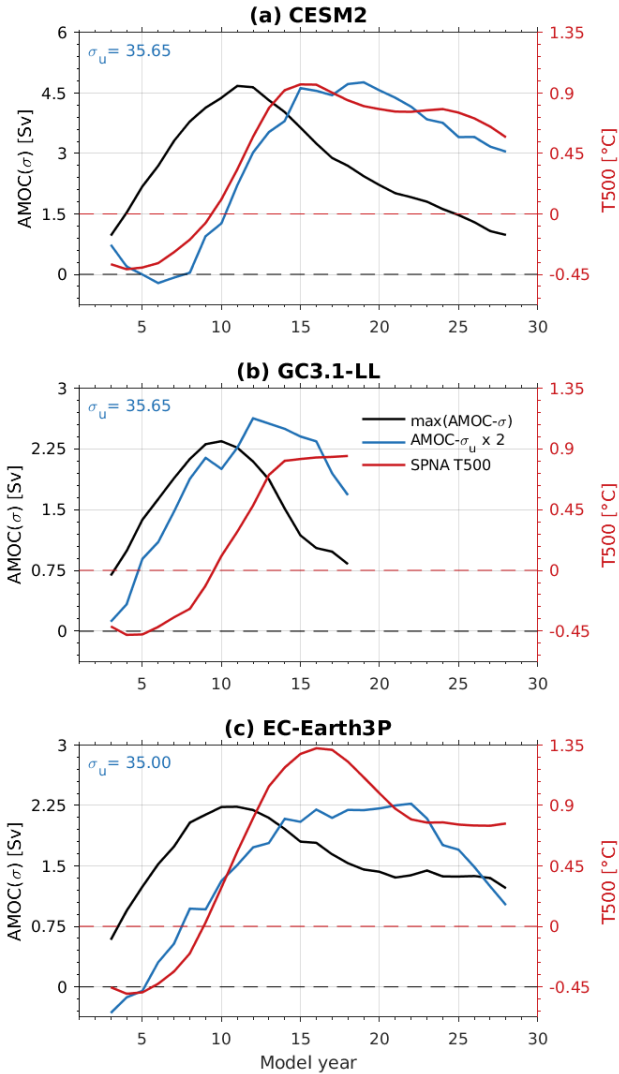


537
 538 Fig. 9. Difference of annual upper 500 m temperature averaged over (a-c) years 1-10, (d-f) years 11-20,
 539 and (g-h) years 21-30 from CESM2 (left), GC3.1-LL (middle), and EC-Earth3P (right). The stippled regions
 540 indicate that the differences are *not* statistically significant at a 95% confidence level with every fourth
 541 stippling shown. The gray contour indicates the 3000 m isobath. The blue box indicates that region where the
 542 SPNA upper 500 m temperature time series in Fig. 4 is averaged.

543 *e. Wider impacts*

544 The anomalously strong NAC at the gyre boundary implies that more warm and salty waters
 545 are advected into the SPNA from the subtropics. Figure 9 shows that the decadal upper ocean
 546 (500 m) temperature anomalies evolve in line with the NAC anomalies. While the forcing
 547 directly generates a cooling in the SPNA (when the +NAO forcing is imposed) for the first
 548 decade (Fig. 9a-c), this cooling is replaced by a much stronger warming in the eastern SPNA
 549 during the second decade (Fig. 9d-f). This warming develops over much of the SPNA and
 550 persists even in the third decade in CESM2 and EC-Earth3P (Fig. 9g-h). The signal-to-noise ratio
 551 of SST exceeds 50% over much of the eastern SPNA during the second decade and even reaches

552 80% in EC-Earth3P at the core of the warming (Fig. S9). The strong heat flux forcing that we
553 impose, equivalent to 2 standard deviations of the NAO, should be also a factor in this large
554 signal-to-noise ratio. Nevertheless, this suggests that a substantial fraction of the decadal SST
555 variability in the eastern SPNA in these models can be explained by the NAO-forced
556 thermohaline processes. Concurrent with the warming in the SPNA, a cold anomaly appears off
557 the Grand Banks west of the MAR, generating a dipole anomaly pattern in all models (Fig. 9d-
558 h). This dipole pattern has been highlighted as a fingerprint of anomalous AMOC strength
559 (Zhang 2008). The anomalous upper ocean salinity pattern closely resembles that of the UOT for
560 the last two decades (Fig. S9). In the absence of forcing that can directly impact upper ocean
561 salinity during the first decade, the upper ocean salinity response in the SPNA is minor. The
562 spatial patterns of both anomalous UOT and salinity are remarkably similar across the models,
563 suggesting that the response of the ocean dynamics to the imposed forcing is consistent across
564 the models regardless of different choices of model numerics and parameterizations. The decadal
565 surface heat flux differences for the second and third decades exhibit a heat release from the
566 ocean in the SPNA (Fig. S10), particularly in the eastern SPNA where the anomalous UOT
567 warms most in all models. This underpins that the SPNA temperature anomalies are driven by
568 the heat convergence associated with the anomalous upper limb AMOC(σ), which is further
569 supported by the paired upper ocean salinity anomalies.

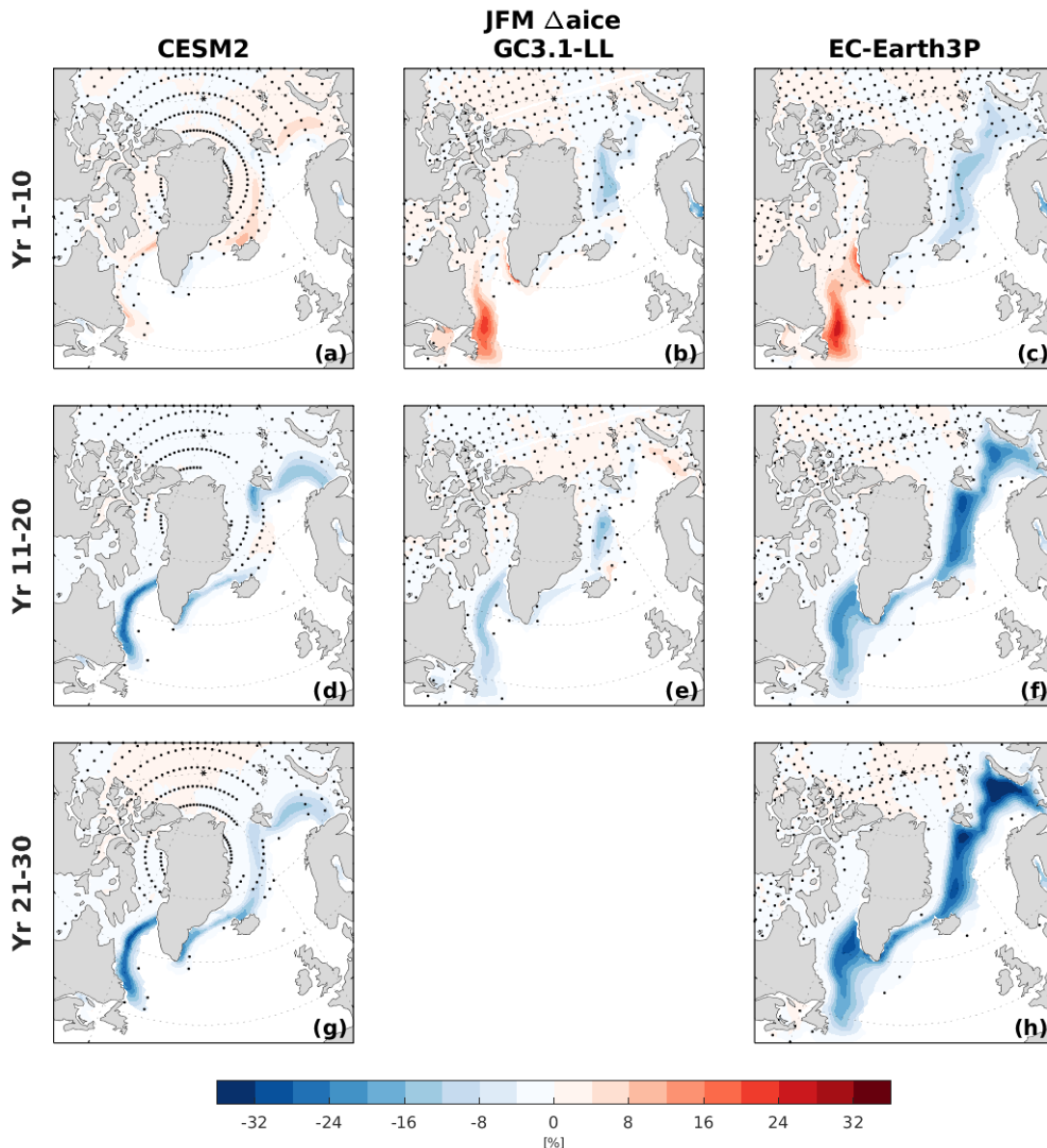


570

571 Fig. 10. Time series of annual maximum AMOC(σ) (black), upper limb of AMOC(σ) (blue), defined at
 572 the density shown in the upper left corner of each panel, and the upper 500 m temperature averaged over the
 573 region in the eastern SPNA (52°-64°N, 15°-40°W) from (a) CESM2, (b) GC3.1-LL, and (c) EC-Earth3P. A 5-
 574 year running average is applied for all time series.

575 Figure 10 shows the time series of the UOT differences averaged over the eastern SPNA
 576 (boxed region in Fig. 9d-f) along with the time series of the maximum and the upper limb of
 577 AMOC(σ) differences at 45°N. The upper AMOC(σ) limb is defined at σ_2 where the anomalous
 578 northward transport is largest in each model (Fig. 7). The UOT, which initially cools under the
 579 forcing, ramps up starting from year 5, reaches a positive maximum around year 15, and stays in
 580 an anomalously warm state through the end of the simulations in all models. This tendency is
 581 generally consistent with the delayed spin-up of the upper AMOC(σ) limb relative to the lower

582 AMOC(σ) limb in all models, supporting the idea that surface northward heat transport
 583 convergence associated with the anomalous upper limb of AMOC is responsible for delayed
 584 UOT changes in the SPNA (Yeager 2020). As discussed earlier, the upper AMOC(σ) limb
 585 increases earlier than the UOT in GC3.1-LL likely because of another anomalous meridional
 586 flow that develops early west of the NAC (Fig. 8).



587
 588 Fig. 11. Difference of JFM sea-ice concentration averaged over (a-c) years 1-10, (d-f) years 11-20, and (g-
 589 h) years 21-30 from CESM2 (left), GC3.1-LL (middle), and EC-Earth3P (right). The stippled regions indicate
 590 that the differences are *not* statistically significant at a 95% confidence level with every fourth stippling shown.

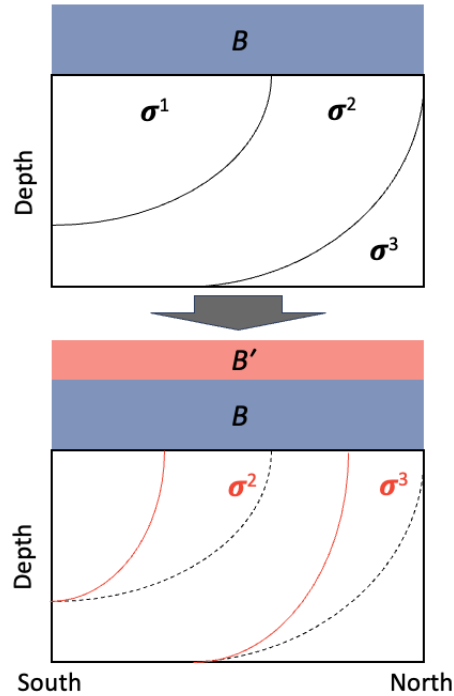
591 Figure 9 also shows that the SPNA temperature anomalies expand to the north into the
592 Nordic Seas and to the west into the LS, particularly evident during the third decade. This
593 suggests that sea ice in the Nordic and Labrador Seas can be affected by the temperature
594 anomalies. The JFM sea-ice concentration differences (Fig. 11) confirm a delayed sea-ice
595 response to NAO forcing. With the arrival of the warm anomaly in the SPNA, all models show a
596 sea-ice concentration decrease along the sea-ice edge, especially in the LS in the second decade
597 (Fig. 11d-f). With the extension of the warm anomaly to the Nordic Seas in CESM2 and EC-
598 Earth3P (Fig. 9d and f), sea-ice concentration also decreases there (Fig. 11d and f), which is
599 further amplified in the third decade (Fig. 11g and h) as the warming further builds up in the
600 Nordic Seas (Fig. 9g and h). The sea-ice concentration decrease is substantially stronger in EC-
601 Earth3P than other two models with more than 30% (20%) decrease in the Barents Sea (the
602 Greenland Sea), consistent with a greater warming in the Nordic Seas than other models. These
603 findings support previous studies (e.g., Mahajan et al. 2011; Yeager et al. 2015) that have
604 demonstrated how AMOC-driven SPNA UOT anomalies can penetrate into the Nordic Seas to
605 drive decadal sea-ice variability there.

606

607 **4. Summary and Discussion**

608 We examine in the present study the response of the North Atlantic to winter NAO-related
609 surface heat flux forcing derived from observational estimates using three CMIP6-class coupled
610 models (CESM2, GC3.1-LL, and EC-Earth3P). The primary focus of the study is to explore the
611 robustness of mechanisms that connect surface-forced WMT/WMF to AMOC strength and
612 associated UOT changes in the SPNA. By focusing on ensemble-mean coupled model response
613 to observation-based NAO forcing, we circumvent biases in each models' own representation of
614 NAO-related surface heat fluxes. This allows for a systematic comparison of the responses to the
615 observed NAO across models that have different background states.

616 The experiments reveal a broadly consistent picture of the ocean responses across the models
617 that support the idea that ocean thermohaline processes play a critical role in NAO-driven
618 changes in UOT and salinity in the SPNA. Although the models produce deep waters at different
619 locations in the background state (western SPNA in CESM2, and eastern SPNA and Nordic Seas
620 in GC3.1-LL and EC-Earth3P), the forcing promotes the largest WMT response consistently in
621 the western SPNA in all models. In the case of +NAO forcing, this WMT response increases the
622 volume of the densest SPNA water masses, leading to an densification and enhancement of the
623 lower (denser) AMOC(σ) limb in the SPNA. The anomalous dense waters generate a zonal
624 pressure (SSH) gradient anomaly through the steric effect around the southern boundary of the
625 SPNA west of the MAR, as they propagate to the south, thus driving an anomalous northward
626 flow in the upper ocean, corresponding to the upper (lighter) limb of AMOC(σ). The anomalous
627 northward flow, equivalent to a strengthening of the NAC, brings more warm and salty
628 subtropical waters into the SPNA, increasing heat and salt content in the SPNA in all models
629 with reverberations on sea ice conditions in the subarctic Atlantic Ocean. The spatial patterns of
630 these responses are strikingly similar between the models, suggesting that the dynamical ocean
631 responses are similar across the models despite the different choices of model numerics and
632 physics. In contrast to the consistent spatial patterns, however, the magnitude of the responses is
633 substantially different across the models. More precisely, CESM2 shows WMT and AMOC
634 responses roughly twice as large as those in GC3.1-LL and EC-Earth3P, and this is largely
635 because more waters are transformed in the western SPNA in CESM2.



- $A(\sigma^2) > A(\sigma^2)$
 $\rightarrow \sigma^2$ exposed to more $B + B'$
- σ^3 exposed to $B + B'$
(WMT=0 before B')

636

637 Fig. 12. Schematic summarizing the processes that lead to the WMT response to the imposed forcing in
638 the western SPNA. The black box represents the idealized western SPNA in depth-latitude plane, and σ^1 , σ^2 ,
639 and σ^3 are isopycnal layers under the climatological buoyancy flux (buoyancy loss) B (black) and when the
640 applied buoyancy forcing B' is applied (red). The black dashed lines in the lower box are the same as the initial
641 isopycnals in the upper box. The outcropping area A is the surface integral of the area where isopycnal layers
642 are in contact with the surface.

643 An intriguing finding of the study is that the WMT response directly related to the imposed
644 heat flux forcing is small (Fig. 4), and largely controlled by the expansion and contraction of A_ρ
645 in response to the forcing and associated exposure of the anomalous A_ρ to background surface
646 heat fluxes. A schematic of this interaction is illustrated in Fig. 12. We may consider
647 climatological isopycnal layers (σ^1 and σ^2) that outcrop in the western SPNA under
648 climatological surface buoyancy forcing B (buoyancy loss; upper box in Fig. 12). When forcing
649 B' (additional buoyancy loss) is imposed, the isopycnals move southward, the area of σ^2
650 becomes wider while the area of σ^1 becomes smaller (lower box in Fig. 12). The wider σ^2 layer
651 is then exposed to B , in addition to B' . In addition, B' can expose an isopycnal layer, σ^3 , to the
652 surface that is not in contact with the surface under B alone. The newly outcropped σ^3 layer is

653 then subjected to surface WMT (which it is not under B alone). The enhanced buoyancy loss in
654 σ^2 and σ^3 layers further expands (i.e., larger A_ρ), thus leading to more exposure to B . Hence,
655 this feedback between outcropping area and background surface buoyancy fluxes enhances
656 WMT much stronger than what B' can directly generate (as $B > B'$). The importance of A_ρ for
657 WMT is pointed out by Petit et al. (2021) for subpolar mode water formation in the Iceland
658 Basin from observational estimates, but here we put forward a more nuanced picture by adding
659 on another important factor, feedback from background surface heat fluxes. The background
660 surface heat loss is stronger in CESM2 than in GC3.1-LL and EC-Earth3P in the western SPNA
661 (Fig. 6), thus so does the interaction with A_ρ . Therefore, the WMT response in the western
662 SPNA is substantially stronger in CESM2.

663 The stronger background surface heat loss in the western SPNA in CESM2 appears to be
664 closely related to larger warm SST biases than other two models (Fig. S12). This suggests that
665 both larger background WMT (Fig. 2) and its response to the NAO heat flux forcing in the
666 western SPNA in CESM2 are likely overestimated. However, while SST biases are relatively
667 small in the western SPNA as well as in the eastern SPNA north of 55°N in GC3.1-LL and EC-
668 Earth3P, sea ice is too extensive relative to satellite observations (Fig.3 d-f). This suggests that
669 the WMT response to the forcing, as well as the background WMT, in these models is possibly
670 underestimated. The importance of surface heat fluxes in WMT highlighted in this study
671 suggests that observational WMT estimates should be sensitive to surface heat flux datasets.
672 Therefore, for a better validation of model performance in WMT, it seems to be important to
673 understand the uncertainty of observational WMT estimates.

674 Despite the larger AMOC response in CESM2, the amplitude of the SPNA temperature (Fig.
675 9) and salinity (Fig. S10) responses is comparable across the models. The northward heat and
676 salinity transport anomaly into the SPNA in our experiment is likely due to, to a great extent, the
677 climatological temperature and salinity carried by anomalous velocity as the surface heat flux
678 forcing is only applied north of 45°N . This suggests that the background temperature and salinity
679 in the subtropics are colder and fresher in CESM2 than in GC3.1-LL and EC-Earth3P. The
680 background upper 500 m temperature and salinity in CESM2 is indeed colder and fresher in the
681 subtropics along the Gulf Stream and the NAC (Fig. S13), suggesting that the effect of the larger
682 anomalous velocity is largely compensated by the colder and fresher conditions relative to

683 GC3.1-LL and EC-Earth3P, thus yielding a comparable SPNA temperature and salinity
684 responses as other two models.

685 A limitation of the present study is the use of low-resolution (1°) ocean models that cannot
686 resolve some important small-scale processes such as mesoscale eddies. Apart from potentially
687 better mean states simulated with a better representation of these processes, high resolution can
688 also have an effect on the response of the ocean to the imposed forcing. For example, eddies
689 shedding from the West Greenland Current have a re-stratifying effect that inhibits the deep-
690 water formation process in the LS (Tagklis et al. 2020). Thus, the WMT response in the LS,
691 which is the largest contribution to the total WMT response to the forcing in all models, may be
692 weakened if such eddies are resolved. Nevertheless, recent studies comparing decadal WMT and
693 AMOC variability between eddy-rich and non-eddy-resolving models show a consistent, central
694 role of WMT in the LS driving decadal AMOC variability at both resolutions (Oldenburg et al.
695 2022; Yeager et al. 2021), despite the fact that the high-resolution mean state is in much better
696 agreement with observations compared to the low-resolution counterpart (i.e., mixed layer depth
697 and overturning strength in the LS; Yeager et al. 2021). While providing insights into the decadal
698 AMOC and SPNA variability in an eddy-rich regime, these studies are all based on CESM1.
699 Thus, it remains to be tested if the same conclusion can be drawn from other high-resolution
700 models.

701 Finally, although the focus of the present study is the ocean response, we conclude with a brief
702 preview of some of the atmospheric responses that will be more fully examined in a forthcoming
703 study (Fig. S14). These responses are based on the multi-model mean of the ensemble means and
704 the second-decade average (year 11-20) when the SPNA response is largest (Fig. 10). The SPNA
705 warming leads to a warming of surface temperature over most of the Northern Hemisphere land
706 (Fig. S14a). However, compared to the response of up to 1°C warming in the eastern SPNA, the
707 signal is generally very weak over land ($< 0.3^\circ\text{C}$). Associated with this summertime warming is a
708 clear northward shift of the tropical rainband in the Atlantic sector that extends further east to
709 Africa and the Indian Ocean (Fig. S14b). In particular, an increase in precipitation in the Sahel
710 region is evident. The experiment also shows a consistent response of SLP in the subtropical
711 North Atlantic centered around the Iberian Peninsula in boreal winter. This subtropical anomaly
712 is accompanied by an anomaly of opposite sign north of 60°N (Fig. S14c) although the sign of

713 the difference does not agree in all models except over Greenland and Iceland. Together with a
714 negative anomaly south of Alaska, the anomaly pattern suggests a negative Northern Annular
715 Mode (NAM)-like response. We note that all these impacts are generally consistent with those
716 associated with AMOC-driven AMV that can be found in literature (e.g., Zhang et al. 2019) and
717 those identified in Kim et al. (2020b) using a similar experiment where NAO heat flux forcing is
718 applied only in the LS.

719

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726

727 *Data Availability Statement.*

728 Because of a large size of data, all NAO heat flux forcing experiments from three models
729 will be made available upon request.

730

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