

Ocean stratification impedes particulate transport to the plumes of Enceladus

Article

Supplemental Material

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Parameter	Value
Ice shell surface radius (r_s)	252 km
Core radius (r_c)	192 km
Ocean thickness (H_o)	40 km
Ice thickness (H_i)	20 km
Ice density (ρ_i)	925 kg m^{-3}
Gravity at ice shell surface (g)	0.113 m s^{-2}
Ocean mean salinity (S_{ref})	5, 10, 15, 17.5, 20, 22.5 g kg^{-1}
Ocean reference pressure (P_{ref})	$\rho_i \int_{H_i}^0 g(z) dz$
Ocean reference temperature (T_{ref})	Freezing temp at P_{ref} and S_{ref}
Ocean reference density (ρ_0)	Density at T_{ref} , P_{ref} and S_{ref}
Core density (ρ_{core})	2370 kg m^{-3}
Rotation rate (Ω)	$5.307 \times 10^{-5} \text{ s}^{-1}$
Core total heat output (F_{tot})	20 GW
Specific heat capacity (c_p)	$4000 \text{ J kg}^{-1} \text{ K}^{-1}$
Eddy diffusivity (κ_{GM})	$1 \text{ m}^2 \text{ s}^{-1}$
Vertical diffusivity (κ_z)	$10^{-5}, 10^{-4}, 10^{-3} \text{ m}^2 \text{ s}^{-1}$
Prandtl number ($\frac{\nu}{\kappa}$)	10
Ice melting viscosity (η_{melt})	10^{-14} Pa s

Table S1: Key parameters used in default numerical simulations and in computation of boundary forcings.

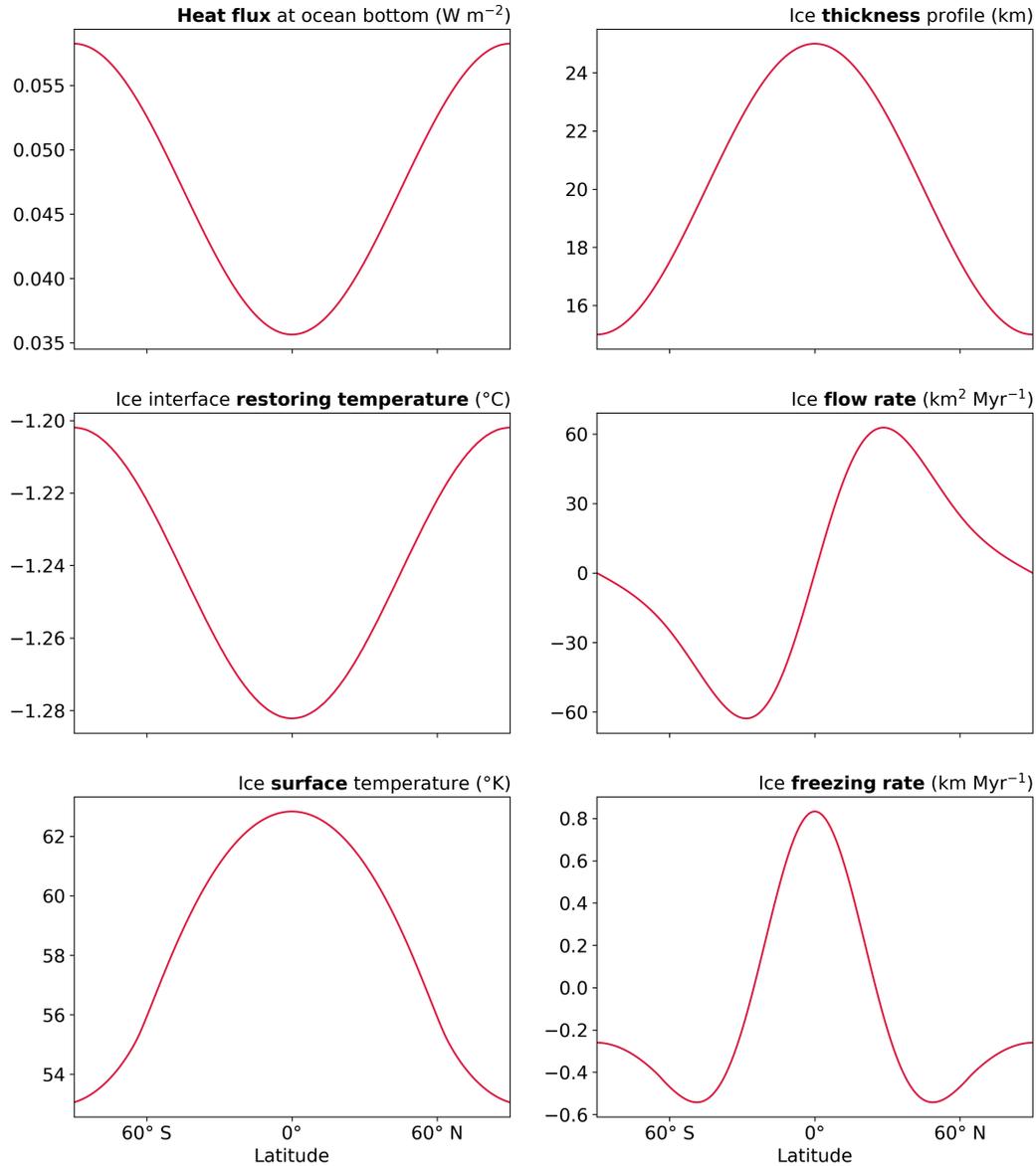


Figure S1: **Top Left:** Bottom heating profile used in simulations (W m^{-2}), assuming a total global core heat output of 20 GW. **Middle left:** An example restoring temperature ($^{\circ}\text{C}$) profile at the ocean top (ice-interface) used in simulations at 20 g kg^{-1} - corresponding to the pressure and salinity dependant freezing temperature. **Bottom left:** Surface temperature ($^{\circ}\text{K}$) of ice used to compute ice flow rate in middle right panel. **Top right:** Idealised ice thickness profile (km) used to compute the ice-interface restoring temperature and ice flow rate (for use in computation of ice freezing rate). Note the topography itself is not used in simulations, which assume a flat ocean top. **Middle right:** Ice flow rate (km^2 per million years) used for computing the freshwater flux at the ocean top, in turn computed from the idealised geometry of the top right panel, assuming a steady state Enceladus ice shell thickness profile. **Bottom right:** Freshwater flux (km per million years) profile applied at the ocean top in simulations, computed using the ice flow rate in the middle right panel.

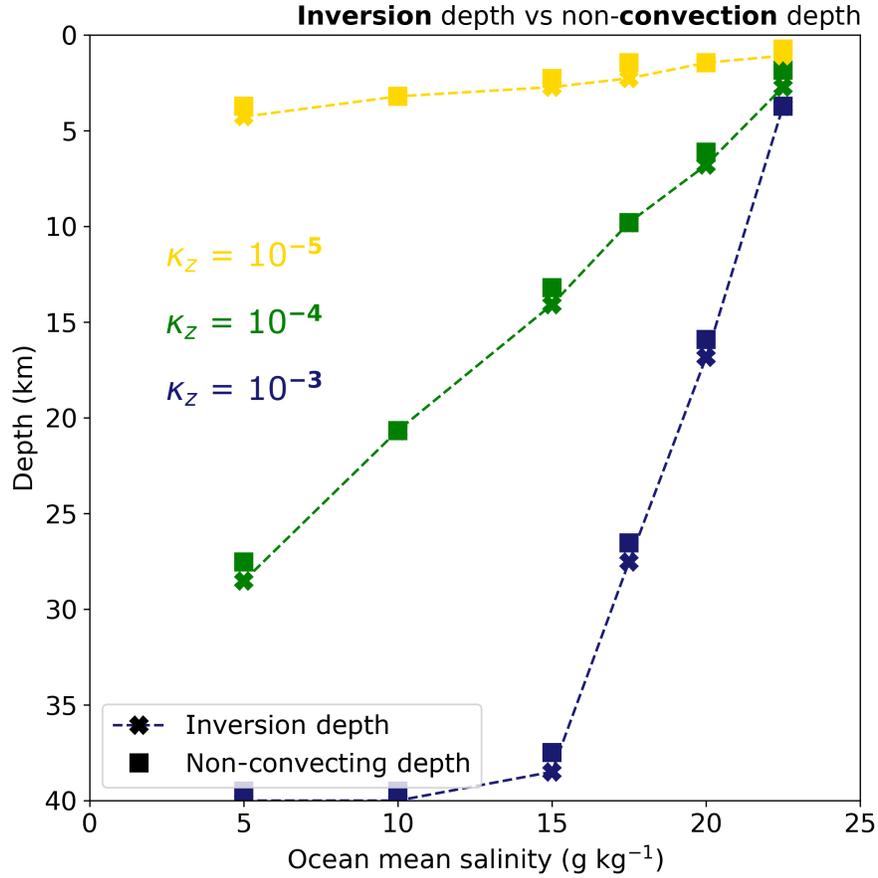


Figure S2: Numerical solution for the inversion depth (km - cross, dashed line - depth at which the thermal expansion coefficient α_T becomes negative) across a range of ocean mean salinity (g kg⁻¹) for three values of κ_z (m² s⁻¹), denoted with different colours, as in Fig. 2 of the main text. Here the inversion depth is plotted alongside the depth at which model convection extending from the ocean bottom stops (km - squares), defined as the depth at which the laterally-averaged convecting time falls below 0.01% of model time. Markers at 40 km depth denote where there is no inversion depth (meaning α_T remained negative across the whole ocean) or non-convecting depth (meaning no convection extended from the ocean bottom) respectively.

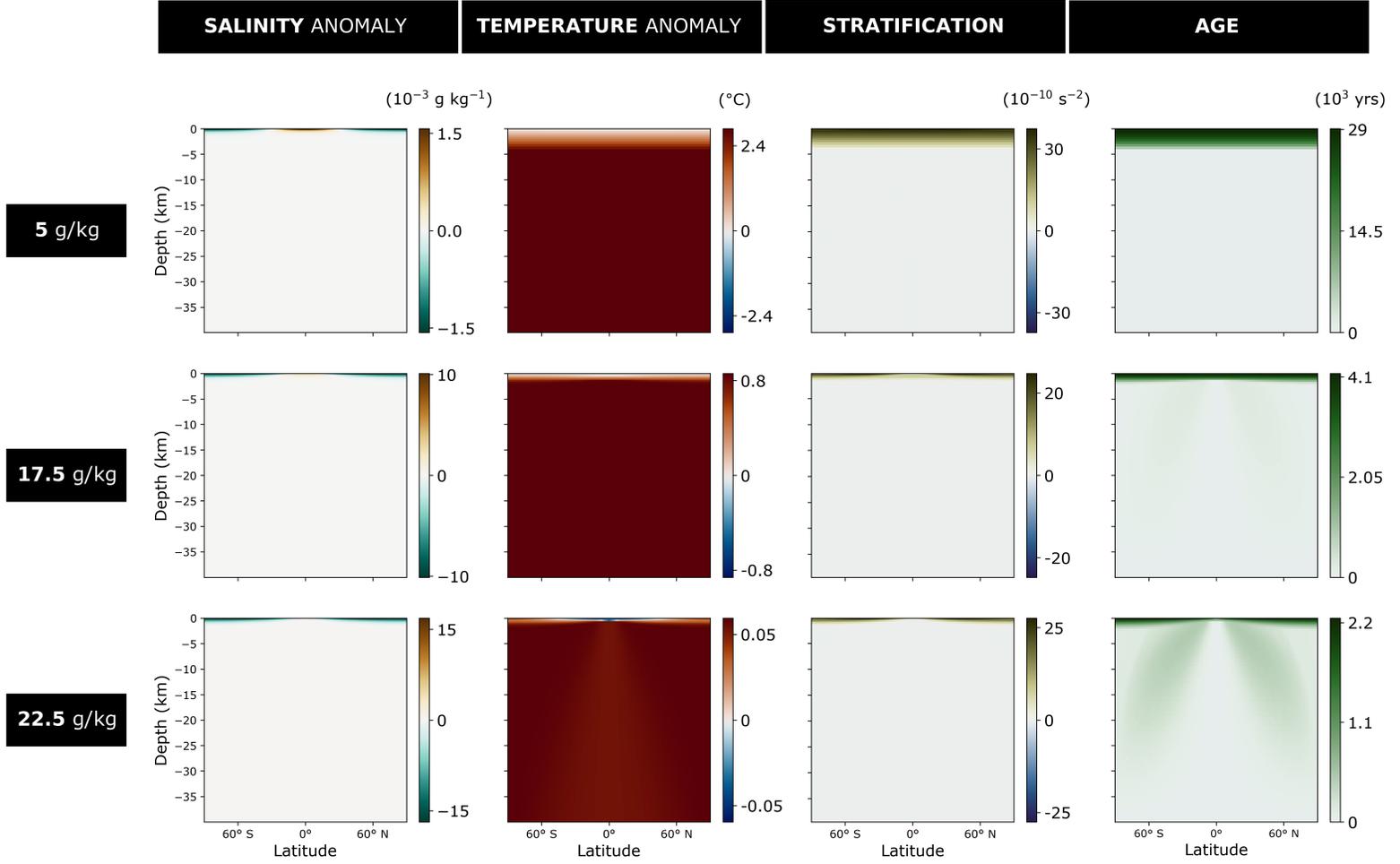


Figure S3: Numerical solutions with effective vertical diffusivity $\kappa_z = 10^{-5} \text{ m}^2 \text{ s}^{-1}$, across three different mean ocean salinities of 5 (top), 17.5 (middle), and 22.5 (bottom) g kg^{-1} , highlighting stratification regimes at a lower κ_z than presented in the main text. Note colour bar scales are saturated and vary throughout. **First column:** Salinity anomaly (g kg^{-1}) taken about the mean salinity. **Second column:** Potential temperature anomaly ($^{\circ}\text{C}$) taken about the simulation reference temperature T_{ref} (freezing temperature computed under 20 km mean ice thickness) of -0.433 (top), -1.106 (middle) and -1.379 (bottom) $^{\circ}\text{C}$. **Third column:** Buoyancy frequency N^2 (s^{-2}) indicating stratification. **Fourth column:** Ideal age of tracers (years), sourced from the ocean bottom.

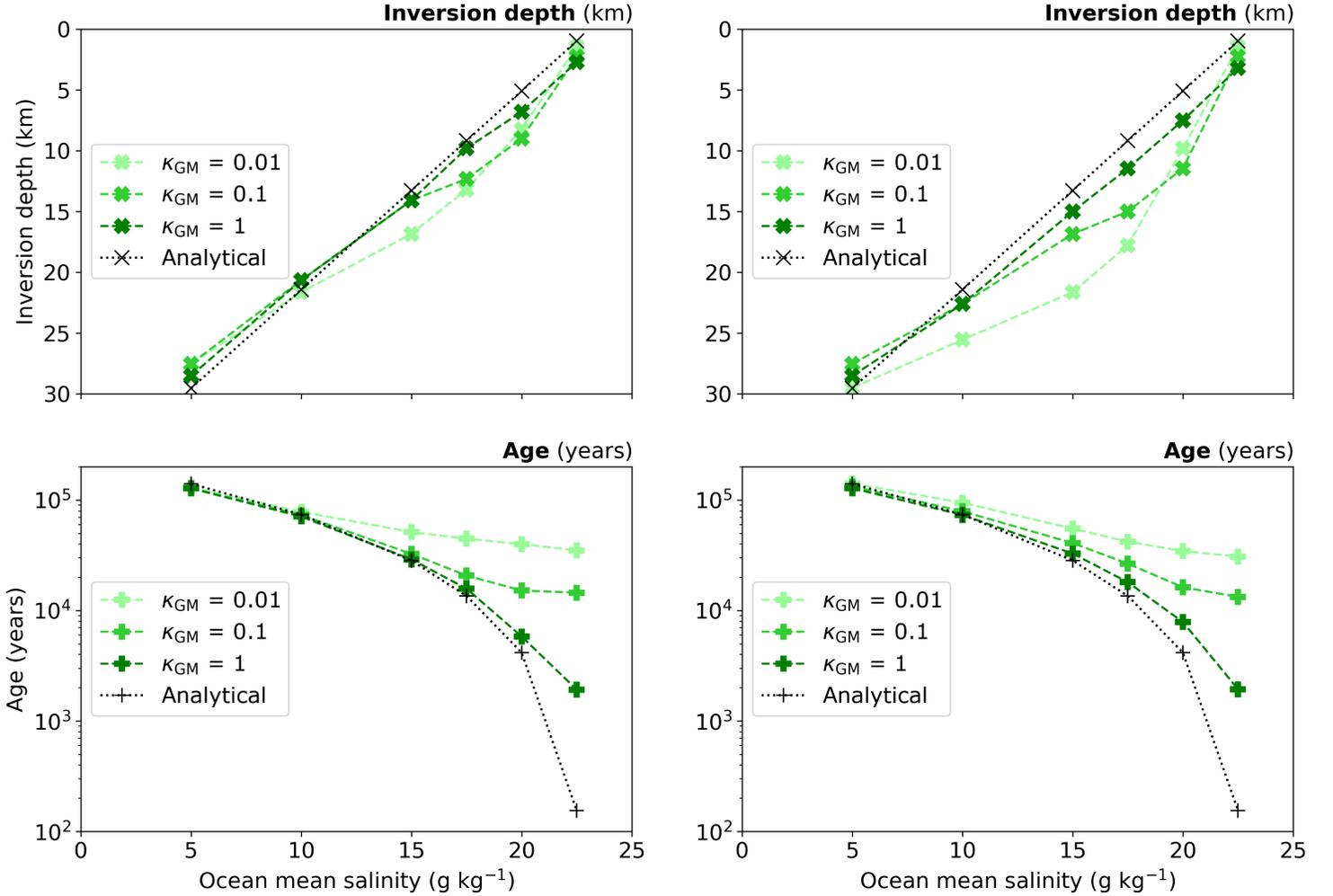


Figure S4: **Top:** Numerical solution for the inversion depth (km - depth at which the thermal expansion coefficient α_T becomes negative), as in Fig. 2 of the main text across a range of ocean mean salinity (g kg^{-1}), but plotted for three different values eddy diffusivity κ_{GM} ($\text{m}^2 \text{s}^{-1}$), indicated with varying shade of green, for constant vertical diffusivity $\kappa_z = 10^{-4} \text{m}^2 \text{s}^{-1}$. The analytical solution (black) using Eq. (5) is plotted alongside for comparison. **Bottom:** Corresponding numerical solution for the ideal age at the south polar ocean-ice interface, plotted alongside the analytical solution (Eq. (6)). **Left:** Solutions at a melting viscosity $\eta_{melt} = 10^{-14} \text{Pa s}$ as used in default simulations of the main paper. **Right:** solutions with $\eta_{melt} = 2 \times 10^{-13} \text{Pa s}$, which creates a 5 times larger magnitude of ice melting and freezing at the ocean top. Note increasing discrepancies relative to analytical solution at higher salinity and lower κ_{GM} , due to the formation of freshwater lenses at the ocean-ice interface at the poles.

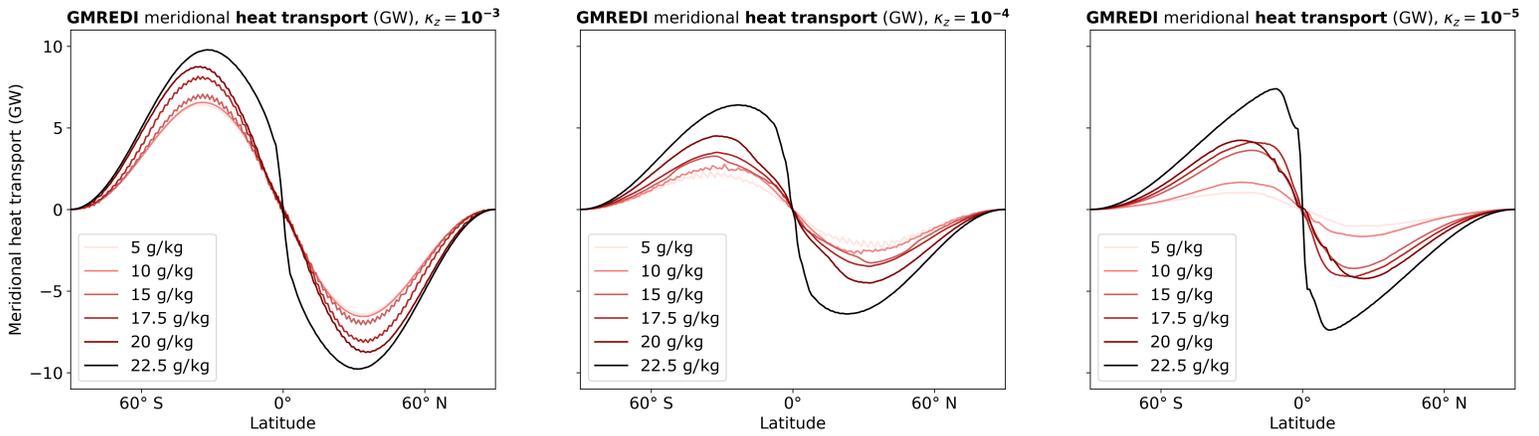


Figure S5: Vertically-integrated global meridional heat transport (GW) achieved by parameterised eddies (GMREDI) in numerical solutions. Plotted across modelled ocean mean salinity at a modelled effective vertical diffusivity κ_z of 10^{-3} (left), 10^{-4} (middle), and 10^{-5} (right) $\text{m}^2 \text{s}^{-1}$. The eddy diffusivity $\kappa_{\text{GM}} = 1 \text{ m}^2 \text{s}^{-1}$ for the presented solutions. Note the heat transports shown here are scaled up by a factor 360 to be representative of the global ocean (as the 2D simulations are based on a 1 degree wide single box in the zonal direction).

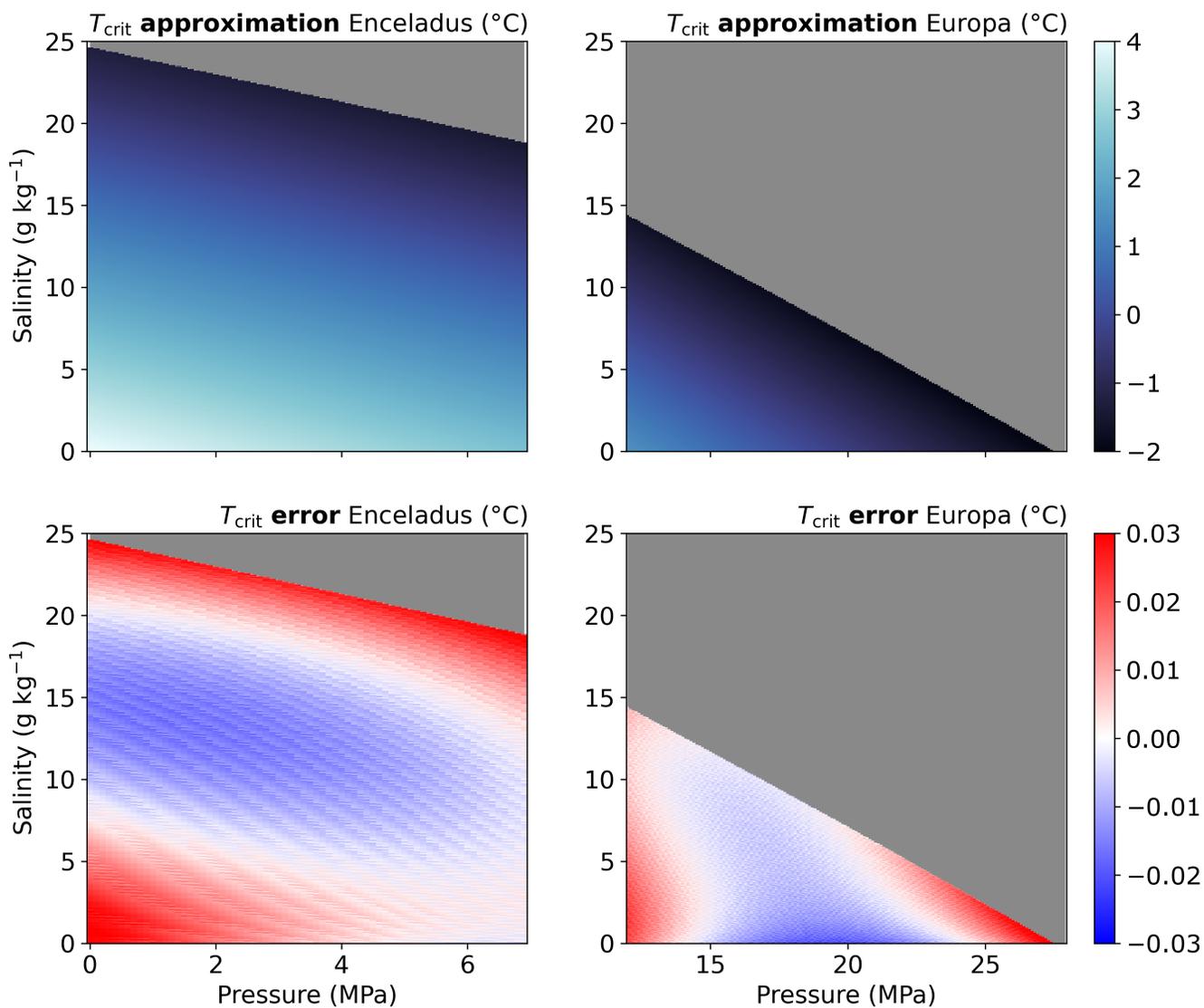


Figure S6: **Top:** Linearly approximated critical temperature (T_{crit} - $^{\circ}\text{C}$) - defining the temperature at which the thermal expansion coefficient (α_T) changes sign - as a function of pressures (MPa) and salinity (g kg^{-1}) plausible for Enceladus (left) and Europa (right) respectively. Grey shading denotes where T_{crit} does not apply because α_T cannot become negative for the given salinity and pressure.

Bottom: Error of linear approximation of critical temperature ($^{\circ}\text{C}$) relative to reference computation described in the main text.

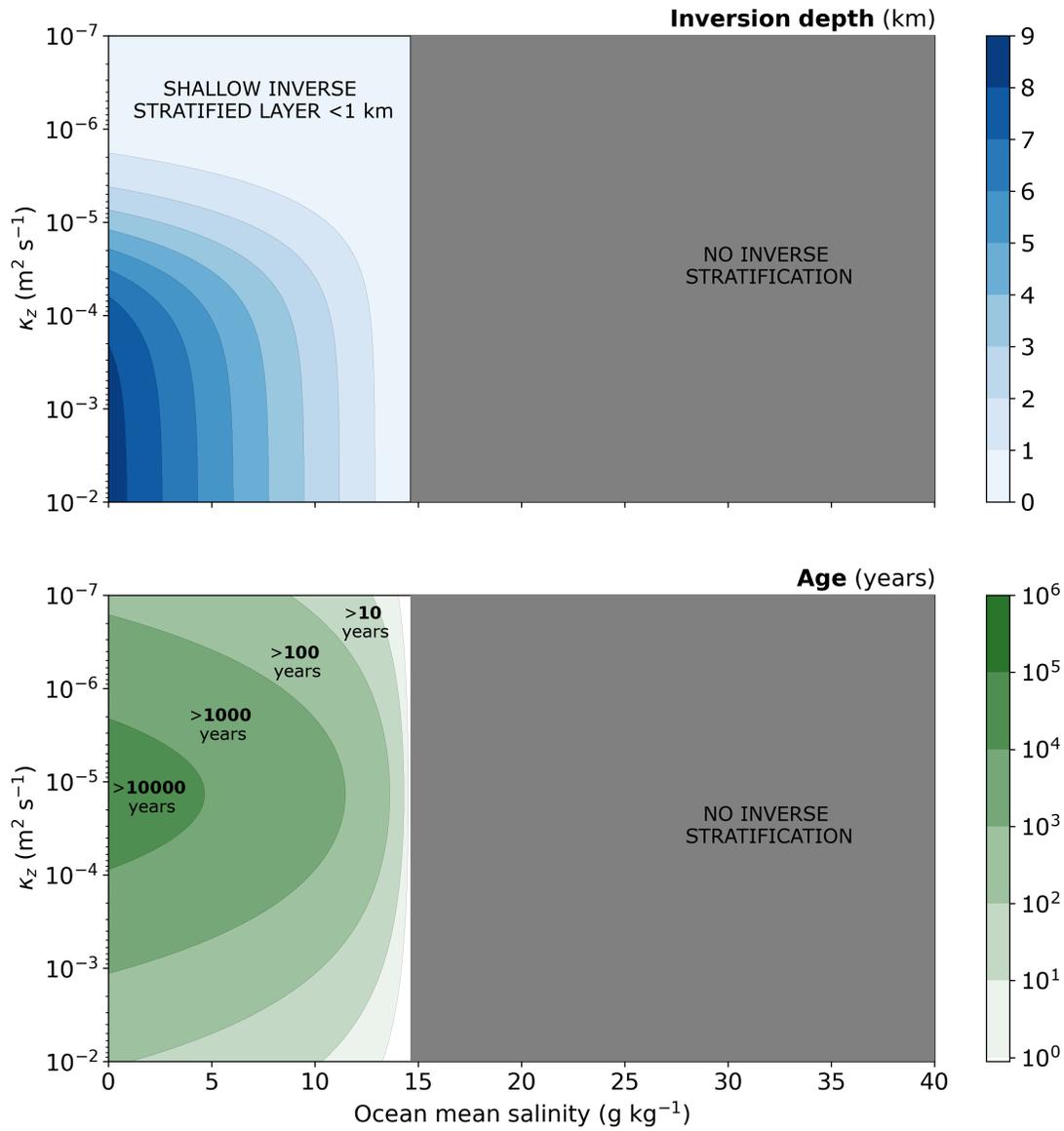


Figure S7: **Top:** Inversion depth H_{strat} (km - depth at which thermal expansion coefficient α_T becomes negative, taken here to define the inverse layer thickness) plotted as a function of ocean mean salinity (g kg^{-1}) and vertical diffusivity κ_z ($\text{m}^2 \text{s}^{-1}$), for Europa - moon of Jupiter. Grey shading denotes where inverse stratification cannot occur, because α_T cannot become negative at the ocean-ice interface pressure (computed under an assumed 10 km ice thickness). **Bottom:** Tracer age (years) at Europa's ocean-ice interface, computed using the theoretical model outlined in the main text. Note that age contours are logarithmic.