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climate model*

Article

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1 **The Impact of Two Coupled Cirrus Microphysics-Radiation Parameterizations on**  
2 **the Temperature and Specific Humidity Biases in the Tropical Tropopause Layer in**  
3 **a Climate Model**

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1 ABSTRACT

2 The impact of two different coupled cirrus microphysics-radiation parameterizations  
3 on the zonally averaged temperature and humidity biases in the tropical tropopause layer  
4 (TTL) of a Met Office climate model configuration is assessed. One parameterization is  
5 based on a linear coupling between a model prognostic variable, the ice mass mixing  
6 ratio,  $q_i$ , and the integral optical properties. The second is based on the integral optical  
7 properties being parameterized as functions of  $q_i$  and temperature,  $T_c$ , where the mass  
8 coefficients (i.e. scattering and extinction) are parameterized as nonlinear functions of the  
9 ratio between  $q_i$  and  $T_c$ . The cirrus microphysics parameterization is based on a moment  
10 estimation parameterization of the particle size distribution (PSD), which relates the mass  
11 moment (i.e. second moment if mass is proportional to size raised to the power of 2 ) of  
12 the PSD to all other PSD moments through the magnitude of the second moment and  $T_c$ .  
13 This same microphysics PSD parameterization is applied to calculate the integral optical  
14 properties used in both radiation parameterizations and, thus, ensures PSD and mass  
15 consistency between the cirrus microphysics and radiation schemes.

16 In this paper, the temperature-non-dependent and temperature-dependent  
17 parameterizations are shown to increase and decrease the zonally averaged temperature  
18 biases in the TTL by about 1 K, respectively. The temperature-dependent radiation  
19 parameterization is further demonstrated to have a positive impact on the specific  
20 humidity biases in the TTL, as well as decreasing the shortwave and longwave biases in  
21 the cloudy radiative effect. The temperature-dependent radiation parameterization is  
22 shown to be more consistent with TTL and global radiation observations.

23

1 **1. Introduction**

2 It is well known that cirrus makes an important contribution to the radiative  
3 balance of the tropical tropopause layer (TTL) through its temperature, spatial  
4 distributions, opacity, and composition. Moreover, it influences the water vapor  
5 concentration in the stratosphere (Heymsfield 1986; Sassen et al. 1989; McFarquhar  
6 et al. 2000; Corti et al. 2006; Wang and Dessler 2006; Stubenrauch et al. 2007; Jensen  
7 et al. 2008; Mace et al. 2009; Yang et al. 2010; Schwartz and Mace 2010; Taylor et  
8 al. 2011; Zhou et al. 2014; Hong and Liu 2015; Hardiman et al. 2015, and references  
9 therein).

10 Typical TTL cirrus will efficiently trap outgoing longwave radiation from the  
11 surface and atmosphere, generally leading to a warming of the surface, and absorb  
12 incoming shortwave radiation at near-infrared wavelengths, generally leading to a  
13 local heating of the upper troposphere (Liou 1986; Liou 2005; Edwards et al. 2007;  
14 Baran 2009; Taylor et al. 2011; Baran 2012; Yi et al. 2013; Zhou et al. 2014; Hong  
15 and Liu 2015; Yang et al., 2015; Hardiman et al. 2015). The balance between cirrus  
16 warming or cooling the upper troposphere depends on its visible optical depth, as  
17 shown by Hong and Liu (2015), who demonstrated that cirrus with visible optical  
18 depths less than unity leads to a net heating of the upper troposphere, whilst optically  
19 thicker cirrus results in a net cooling of the upper troposphere. Moreover, subvisual  
20 cirrus occurs in the TTL, as found by Lawson et al. (2008), and Lee et al. (2009)  
21 showed that these clouds may exert a net radiative effect on the order of about 1.1  
22  $\text{Wm}^{-2}$ .

23 In general, the role of cirrus in either heating or cooling the TTL depends not only

1 on the visible optical depth but also on the microphysics and the scattering and  
2 absorption properties of atmospheric ice (Baran et al. 2014a; Yang et al. 2015, and  
3 references therein). There have been aircraft campaigns which have examined the ice  
4 microphysics composition of the TTL. In the study by Heymsfield (1986), the  
5 common occurrence of trigonal particles was reported but more recent studies by  
6 Lawson et al. (2008) have found few occurrences of trigonal ice particles. In the study  
7 by McFarquhar et al. (2000), they found habit mixtures comprising 50% of hexagonal  
8 columns and trigonal ice crystals in the subvisual TTL cirrus they studied.  
9 Meanwhile, Lawson et al. (2008) found the occurrence of quasi-spherical ice particles  
10 to be the most common particle type for all crystal sizes. However, it is uncertain as  
11 to whether these particles are actually quasi-spherical due to the limiting resolving  
12 power of the microphysics instrumentation used at the time; therefore, the appearance  
13 of quasi-sphericity could be due to diffractive and optical effects, as noted by Cotton  
14 et al. (2010) and references therein. On the other hand, these particles could be quasi-  
15 spherical, but instruments are required that can adequately resolve these ice crystals  
16 of an uncertain shape, such as the small ice detector described in Ulanowski et al.  
17 (2006). For crystal sizes greater than 65  $\mu\text{m}$ , Lawson et al. (2008) found habit  
18 mixtures comprising mostly hexagonal plates and irregular ice crystals. In the study  
19 by Jensen et al. (2008) they noted the occurrence of hexagonal plate aspect ratios of  
20 6:1 of nearly 100  $\mu\text{m}$  in size near the TTL, which were reported to be unassociated  
21 with deep tropical convection. The more recent Airborne Tropical Tropopause  
22 Experiment (ATTREX) reported by Jensen et al. (2015) found bullet rosettes to be  
23 commonly occurring with little evidence of aggregated ice crystals in the cirrus cases

1 they sampled. The two particle size distributions shown in Jensen et al. (2015) extend  
2 to maximum dimensions of about 50  $\mu\text{m}$  and almost 200  $\mu\text{m}$ , at temperatures of <  
3 195 K and between about 195 and less than 210 K, respectively. Meanwhile, Lawson  
4 et al. (2008) report particle distributions extending to 165  $\mu\text{m}$  in the subvisual TTL  
5 cirrus they sampled, which contrasts with the study of McFarquhar et al. (2000), who  
6 found that particle maximum dimensions did not extend to beyond 50  $\mu\text{m}$ .

7 Determining the habit mixtures and particle distributions in TTL cirrus is important,  
8 as the choice of assumed particle distributions in climate models will influence their  
9 predicted radiative effect of cirrus. For instance, in the study of Mitchell et al. (2008)  
10 they show that by assuming two different parameterizations of the small ice mode,  
11 the shortwave radiative effect differences in the annual zonally averaged fluxes at  
12 top-of-atmosphere in the tropics can be up to about  $-25 \text{ Wm}^{-2}$ . They also show that  
13 the longwave differences in the tropics can be up to  $20 \text{ Wm}^{-2}$ . Clearly, from the  
14 studies cited above, there is uncertainty as to the most general ice microphysics  
15 composition in the TTL to assume in models, in turn, this will affect climate model  
16 simulations of the TTL, as differing climate models assume dissimilar microphysics  
17 properties, which may lead to an excess in TTL warming or cooling (Hardiman et al.  
18 2015). Further campaigns in the TTL are necessary if such uncertainties reported by  
19 Hardiman et al. (2015) are to be reduced in climate models. However, recent TTL  
20 process studies by Hardiman et al. (2015) have shown that physically improved  
21 climate model parameterizations do not necessarily improve simulations of the TTL.  
22 For example, recent parameterizations of cirrus microphysics and bulk optical  
23 properties by Furtado et al. (2015) and Baran et al. (2014a), respectively, were shown



1 to increase the temperature and specific humidity biases in the TTL in recent global  
2 configurations of the Met Office Unified Model (MetUM). In this paper, the physical  
3 reasons are explored as to why the Baran et al. (2014b) parameterizations, hereinafter  
4 referred to as B014b, increased the TTL temperature bias in the climate model. This  
5 paper also presents a new bulk optical property parameterization of cirrus that is  
6 demonstrated to provide a more accurate simulation of the TTL in the MetUM as well  
7 as an improved representation of MetUM simulations of the cloudy shortwave and  
8 longwave radiative effects. The paper is organised as follows: Section 2 describes the  
9 bulk optical property parameterization and its accuracy, and includes comparisons  
10 between the new and the B014b parameterizations. Section 3 briefly describes a  
11 baseline MetUM configuration and compares the impact of the new and B014b  
12 parameterizations on this configuration against both observations and analyses.  
13 Section 4 discusses the results.

## 14 **2. The parameterization**

15 The parameterization presented in this paper is based on the ensemble model of cirrus  
16 ice crystals fully described in Baran and Labonnote (2007), whereby the bulk optical  
17 properties are derived by averaging habit-dependent, single-scattering properties over an  
18 assumed PSD; a figure of the model is shown in Fig. 2 in B014b. A brief description of  
19 the model is given here. The ensemble model consists of six shapes, and these are a  
20 hexagonal ice column of an assumed aspect ratio (AR) of unity, the six-branched bullet  
21 rosette, and then hexagonal monomers are randomly attached to build four ice aggregate  
22 models, which consist of three-, five-, eight- and, finally, ten-branched hexagonal ice  
23 aggregates. The latter aggregate models can cover the largest ice crystal sizes found in the

1 PSD, whilst the former members can populate the smaller end of the ice crystal PSD.  
2 Alternatively, weights can be assigned to each member of the ensemble model at each  
3 PSD bin size to compute the bulk optical properties. Here, use is also made of the same  
4 bulk ice optical property database that was used to develop the B014b parameterization.  
5 The database is composed of the following bulk optical properties: the extinction and  
6 scattering coefficients, the single-scattering albedo,  $\omega_0$ , and the asymmetry parameter,  $g$ .  
7 These bulk optical properties are defined by Eqs. (1–3) in B014b, and the same  
8 definitions are used in this paper. A full description of that database can be found in  
9 B014b and in Baran et al. (2014a), hereinafter referred to as B014a. However, brief  
10 descriptions of the bulk optical property database and its experimental validation are  
11 given here. The database described in B014a consists of 20662 values of ice water  
12 content (IWC) and  $T_c$ , which were compiled from a number of cirrus field campaigns  
13 described in Baran et al. (2011a), and these in-situ campaigns were predominantly  
14 located between northern Europe and the tropics. However, most of the in-situ IWC  
15 values that were compiled by Baran et al. (2011a) were not obtained at temperatures  
16 colder than  $-60^\circ\text{C}$ . Due to this lack of in-situ ice microphysics measurements down to  
17 temperatures as low as  $-80^\circ\text{C}$ , in B014a, IWC values were randomly selected from  
18 temperatures warmer than  $-60^\circ\text{C}$  and placed at temperatures between  $-60$  and  $-80^\circ\text{C}$ . In  
19 this way, the IWC and  $T_c$  space was built up to 20662 values, which ranged between  
20 about  $3.0 \times 10^{-3}$  and about  $10^{-9} \text{ kg m}^{-3}$ , and between about  $-80$  and  $0^\circ\text{C}$ , respectively. This  
21 range in IWC and  $T_c$  is sufficient to cover the range in IWC and  $T_c$  values that are likely  
22 to be predicted in the MetUM.

23 In B014a and B014b, the IWC and  $T_c$  values were related to the PSD through a

1 moment estimation parameterization of the PSD due to Field et al. (2007), and this  
2 parameterization is further described below. The IWC– $T_c$  generated PSDs were  
3 previously applied to the ensemble model single-scattering properties described in B014a  
4 to obtain the cirrus bulk optical properties used in B014b at 145 wavelengths between  
5 0.20 and 120  $\mu\text{m}$ . The same bulk ice optical properties utilized in B014b are used in this  
6 paper.

7 The coupled IWC– $T_c$  bulk ice optical properties used in this paper have been validated  
8 using a variety of aircraft and space-based solar and infrared radiometric measurements.  
9 For instance, in B014a the ensemble model bulk ice optical properties were applied to a  
10 fast radiative transfer model to simulate very high-resolution solar (between about 3.4  
11 and 4.1  $\mu\text{m}$ ) and infrared (between about 8.0 and 12.0  $\mu\text{m}$ ) aircraft-based brightness  
12 temperature measurements obtained from directly above midlatitude cirrus of visible  
13 optical depth varying between about 0.03 and 0.06. It is shown in the paper that the solar  
14 and infrared brightness temperature measurements were mostly simulated to within  $\pm 2$   
15 and  $\pm 1$  K, respectively. Note also that the lower visible optical depth of 0.03 is generally  
16 taken as the upper limit for “subvisual” cirrus, as defined by Sassen and Cho (1992).  
17 Moreover, B014a also showed that the same ensemble microphysical model could  
18 forward model deep frontal cirrus effective radar reflectivity at 35 GHz and microwave  
19 brightness temperatures at 190 GHz to generally within  $\pm 2$  dBZ<sub>e</sub> and  $\pm 2$ K, respectively.  
20 Thus, B014a showed that in the case of the solar and infrared measurements, the same  
21 microphysical model can be consistently applied across the spectrum using a full set of  
22 optical properties (i.e. the scattering phase function and integral optical properties) to  
23 simulate those observations. Meanwhile, Sourdeval et al. (2015) demonstrated that the

1 same ensemble model bulk ice optical properties used in B014a could also replicate  
2 Cloud-Aerosol Lidar with Orthogonal Polarization (CALIOP) and in-situ polar  
3 nephelometer measurements of the volume extinction coefficient of midlatitude cirrus to  
4 generally within the standard deviation of both CALIOP and polar nephelometer  
5 measurements. The in-situ and CALIOP volume extinction coefficients varied between  
6 less than about 0.2 and 1.4 km<sup>-1</sup>, at the wavelengths of 0.532 and 0.80 μm, respectively.  
7 The ensemble model's predicted bulk optical properties have also been globally validated  
8 by Vidot et al. (2015) and Letu et al. (2015).

9 In Vidot et al. (2015) it is shown that an equivalent weighting of 0.30, 0.30, 0.10, 0.20,  
10 and 0.10 applied to each of the ensemble habit extinction and scattering coefficients at  
11 each bin size of the Field et al. (2007) PSD parameterization best simulated cirrus  
12 infrared brightness temperature measurements from the Imaging Infrared Radiometer  
13 (IIR) instrument at wavelengths of 8.65, 10.60 and 12.05 μm. The combined mean  
14 brightness temperature bias using the weighted model optical properties was found to be  
15 only 0.43 K with a standard deviation of 6.85 K for visible optical depths between 0.03  
16 and 4.0. The IIR brightness temperature simulations were based on more than 26000  
17 profiles of IWC from the 2C-ICE and DARDAR products (Deng et al. 2010; Delanöe and  
18 Hogan 2010). The IWC product profiles and atmospheric profiles (the latter profiles came  
19 from the European Centre for Medium Range Weather Forecasts and were colocated with  
20 CALIOP) were located between the latitudes of about ±60° and at altitudes between about  
21 440 and 50 hPa, thus covering the region of the TTL, and most of the profiles were  
22 located in the tropics. Meanwhile, Letu et al. (2015) used global observations from the  
23 Polarization and Directionality of the Earth's Reflectances-3 (POLDER-3) measurements

1 on board the Polarization and Anisotropy of Reflectances for Atmospheric Sciences  
2 coupled with Observations from a Lidar (PARASOL) to show that the ensemble model's  
3 predicted ice cloud solar optical depths at a wavelength of 0.865  $\mu\text{m}$  best minimized  
4 differences between observations, with all ice crystal models considered. The POLDER-3  
5 analysis of Letu et al. (2015) is based on 589246 pixels, with each pixel size being  
6 approximately 6 km  $\times$  6 km. In most pixels the scattering angle range covered is between  
7 60° and 160°, and in some pixels the scattering angle range covered is between 160° and  
8 180°. The pixels were located between the latitudes of about  $\pm 90^\circ$ , but the vast majority  
9 of pixels were located at latitudes between about  $\pm 60^\circ$ . The data covers the period  
10 between the 20<sup>th</sup> to 22<sup>nd</sup> of March, June, September and December 2008. Therefore, all  
11 meteorological seasons are covered in the time period studied by Letu et al. (2015), and  
12 the study is, thus far, the most comprehensive and consistent analysis of POLDER-3  
13 global data. Given the preceding evidence, the ensemble model bulk ice optical properties  
14 have been sufficiently validated globally to apply to a parameterization of cirrus optical  
15 properties in a climate model. Such validation exercises described above are necessary to  
16 carry out to show that the full set of optical properties are consistent with measurements  
17 from across the spectrum in order to provide confidence in their applicability to climate  
18 models. As a further example of this validation approach, see for example, the paper by  
19 Holz et al. (2015).

20 The PSDs applied to the ensemble model single-scattering properties in the  
21 preceding literature assumed the Field et al. (2007) moment estimation  
22 parameterization of the PSD, hereinafter referred to as F07. A full derivation of the  
23 PSD parameterization is given in F07, but a brief description is given here. The Field

1 et al. (2007) parameterization is based on 10000 in-situ measurements of the PSD and  
2 IWC, which were measured between the temperatures of 0°C and -60°C during a  
3 number of cirrus field campaigns located in the midlatitudes and tropics. The  
4 parameterization is based on the moments of the PSD, given by the integral product  
5 of  $D^n$  and  $f(D)$ , where  $D$  is the maximum dimension of the ice crystal and  $f(D)$  is the  
6 PSD, which gives the particle number concentration at each  $D$ , and  $n \geq 0$ . Therefore,  
7 the zeroth moment (i.e.  $n=0$ ) is the total number concentration of particles per unit  
8 volume of cloudy air. The PSD parameterization is related to the total IWC through  
9 some assumed relationship between mass and  $D$ , where in the case of aggregating ice  
10 crystals the mass of ice  $\propto D^2$  (Westbrook et al. 2004; Cotton et al. 2013, and  
11 references therein). The other moments of the PSD are related to the 2<sup>nd</sup> moment ( $M_2$ )  
12 through a relationship of the form  $M_n = \sigma_n M_2^{\beta_n}$ , where  $\sigma_n$  and  $\beta_n$  are functions of  $T_c$ .  
13 Therefore, given all  $M_n$ , the full PSD can be estimated from the IWC and  $T_c$  values  
14 via an assumed mass–dimensional relationship. It should be noted here that to reduce  
15 the impact of ice crystal shattering on the PSD parameterization, the in-situ PSDs in  
16 F07 were filtered by using the measured inter-arrival times of the ice crystals as  
17 described in F07 and in Field et al. (2006), and ice crystals with  $D < 100 \mu\text{m}$  were  
18 ignored. However, the PSD parameterization does not ignore the shape of the ice  
19 crystal PSD at  $D < 100 \mu\text{m}$ , but instead assumes an exponential PSD (Field and  
20 Heymsfield 2003) that is added to a modified gamma distribution at an ice crystal size  
21 of about  $100 \mu\text{m}$ . The degree to which in-situ PSDs are skewed due to ice crystal  
22 shattering has more recently been studied by Korolev et al. (2013). In that paper, they  
23 found that if filtering alone is applied without any modifications to the in-situ

1 microphysics probes on board aircraft, such as the use of anti-shatter tips, the PSD  
2 becomes significantly skewed from the best measured PSD at  $D \leq 200\text{--}175 \mu\text{m}$ ,  
3 where the best measured PSDs were determined using both anti-shatter tips and  
4 filtering. Korolev et al. (2013) state that the precise ice crystal size at which  
5 significant divergence from the best-estimated PSD occurs cannot as yet be  
6 determined due to insufficient statistical sampling of ice clouds. In other words, there  
7 are too few cirrus field campaigns on which to base firm conclusions. Therefore, it is  
8 currently unknown as to the degree to which the F07 parameterization has been  
9 affected by the shattering of ice. However, this parameterization is more  
10 representative of cirrus PSDs than the Houze et al. (1979) PSD parameterization,  
11 which is the current assumption in the MetUM operational model. The Houze et al.  
12 (1979) parameterization is based on 37 in-situ PSDs, 90% of which were measured at  
13 temperatures warmer than  $-30^\circ\text{C}$ . Currently, within the operational MetUM, the  
14 Houze et al. (1979) estimated PSD is kept constant at temperatures colder than  $-35^\circ\text{C}$ .  
15 This assumption means that at much colder temperatures, such as those that occur  
16 within the TTL region, there will be orders of magnitude more frequently occurring  
17 large ice crystals than there should be. This has clear implications for the assumed fall  
18 speed of ice crystals within the MetUM model, which must be artificially increased to  
19 several  $\text{ms}^{-1}$  to accommodate space-based radiometric shortwave and outgoing  
20 longwave measurements of flux at top-of-atmosphere (TOA).

21         The F07 parameterization has been experimentally validated by Baran et al.  
22 (2011b) and Furtado et al. (2015). In Baran et al. (2011b) the moment estimated PSD  
23 was found to fit well to several cases of averaged in-situ measured tropical PSDs. For

1 several cases of midlatitude cirrus Furtado et al. (2015) compared the F07 moments to  
2 in-situ measured moments, whilst ignoring particles less than 100  $\mu\text{m}$  in size, due to  
3 the current uncertainties in measuring the size of small ice at  $D < 100 \mu\text{m}$ , and found  
4 good correlations between the parameterization and measurements for all cases  
5 considered. However, poor correlations emerged in that study when the moments  
6 predicted by the Houze et al. (1979) parameterization were compared against the in-  
7 situ derived moments. The study of Furtado et al. (2015) emphasizes that the F07  
8 parameterization is a better representation of the ice PSD to apply to climate models  
9 in general than the Houze et al. (1979) parameterization. This last statement is  
10 especially true in the TTL, given the above description of the Houze et al. (1979)  
11 parameterization.

12         Given that the F07 PSDs are related to a model prognostic variable (i.e. the mass  
13 carrying moment  $M_2$ ) and this varies as a function of  $T_c$ , given that the ensemble model  
14 single-scattering properties have been integrated over these PSDs, the bulk ice optical  
15 properties can also be directly related to the climate model prognostic variable IWC (i.e.  
16  $m(D)=\text{constant} \times D^2$ , where  $m$  is the mass in units of kg) and  $T_c$  or just IWC (as is the  
17 case in B014b). The behavior of the 20662 bulk ice optical properties,  $\omega_0$  and  $g$ , in  $M_2$ - $T_c$   
18 space, is shown in Figs. 1a and 1b, respectively, at a wavelength of 1.575  $\mu\text{m}$ . This  
19 wavelength is chosen as an example to show the variation of the bulk ice optical  
20 properties in the full space used for the later parameterizations. In Fig. 1a, it is shown that  
21  $\omega_0$  varies between the values of 0.7 and 1.0, at the coldest temperature values, down to  
22 about  $-80^\circ\text{C}$  and  $M_2 \sim 10^{-5} \text{ m}^{-1}$ , and  $\omega_0$  is near unity. At these values of  $M_2$  and  $T_c$ , the PSD  
23 is very narrow and will be largely composed of small ice crystals with much fewer



1 occurring large ice crystals; therefore at this wavelength the smaller ice crystals will  
 2 efficiently scatter incident radiation. At much larger values of  $M_2$  and at much warmer  
 3 temperatures, at  $0.01 \text{ m}^{-1}$  and  $-15^\circ\text{C}$ , respectively, the value of  $\omega_0$  decreases to a value of  
 4 about 0.8. At such  $M_2$  and  $T_c$  values, the PSD becomes much broader than previously,  
 5 with much more frequently occurring large ice crystals, which at this wavelength will  
 6 absorb incident radiation, thereby decreasing  $\omega_0$ . The converse behavior is shown, at the  
 7 same values of  $M_2$  and  $T_c$ , for  $g$  in Fig. 1b, where  $g$  has the corresponding values of about  
 8 0.78 and 0.88. The reasons for these values are the same as those given for the behavior  
 9 of  $\omega_0$ . Figures 1a–1b show that the spatial distribution of  $\omega_0$  and  $g$  is physically to  
 10 expectation and would be a similar distribution if the optical properties were plotted as a  
 11 function of some characteristic size. Here, we demonstrate this using the characteristic  
 12 size usually assumed in radiation schemes in climate models, which is the mean effective  
 13 dimension,  $D_{\text{eff}}$ , defined after Foot (1988) as:

$$14 \quad D_{\text{eff}} = \frac{\text{mass}_t}{\rho A_t} \quad (1)$$

15 where in Eq. (1)  $\text{mass}_t$  and  $A_t$  are the total mass and orientation-averaged projected area  
 16 of the PSD, respectively, and  $\rho$  is the density of solid ice. Here, we assume that  $\rho=1000$   
 17  $\text{kg m}^{-3}$  to keep units in Eq. (1) consistent, hence the units of  $\text{mass}_t$  and  $A_t$  being in SI.  
 18 Firstly, we show that the moment  $M_2$  can easily be related to  $D_{\text{eff}}$ , as shown in Fig. 2,  
 19 where in Fig. 2, as an example, we assume that the  $\text{mass}_t$  and  $A_t$  are computed from the  
 20 first member of the ensemble model, which is the hexagonal ice column of  $\text{AR}=1.0$ . As  
 21 can be seen from Fig. 2, the 2<sup>nd</sup> moment or, equivalently, IWC (by assuming some mass–  
 22  $D$  relationship) could easily be related to  $D_{\text{eff}}$  as a function of  $T_c$  or just by relating IWC

1 directly to  $D_{\text{eff}}$  by using some fitting procedure. Indeed, this latter approach was used by  
2 McFarquhar et al. (2003), who parameterized IWC explicitly as a function of an effective  
3 radius ( $D_{\text{eff}}$  can be related to the effective radius,  $r_{\text{eff}}$ , by  $r_{\text{eff}}=(3/4)\times D_{\text{eff}}$ ). Therefore, the  
4 bulk ice optical properties can also be easily parameterized as a function of  $M_2$  or IWC  
5 and  $T_c$ , as is done in B014b (i.e. IWC alone) and in this paper (i.e. IWC and  $T_c$ ). We  
6 note that, for aggregating ice crystals both mass and diameter are proportional to  $D^2$ , so  
7  $D_{\text{eff}}$  becomes independent of IWC. A further characteristic size could also be defined,  
8 such as the ratio between the 3<sup>rd</sup> and 2<sup>nd</sup> moments of the PSD. If the 2<sup>nd</sup> moment is the  
9 mass carrying moment  $M_2$ , this characteristic size is the mean mass-weighted size,  $D_{\text{mmw}}$ .  
10 The relationship between  $D_{\text{eff}}$  and  $D_{\text{mmw}}$  is characterized by Figs. 3a–3b, respectively,  
11 which show the variation of  $\omega_0$  at a wavelength of 1.575  $\mu\text{m}$  as a function of the  
12 characteristic sizes and  $T_c$  for all 20662 F07 PSDs. In the case of  $D_{\text{eff}}$ , in the figures, the  
13 same definitions are used as above. The figure shows that  $D_{\text{eff}}$  is simply some multiple of  
14  $D_{\text{mmw}}$  and the spatial distribution of  $\omega_0$  in the spaces of characteristic size— $T_c$  is exactly  
15 the same. Furthermore, the distribution of  $\omega_0$  shown in Figs. 3a–3b is relatable to Fig.  
16 1a. That is, at the same  $T_c$  values of  $-80^\circ\text{C}$  and  $-15^\circ\text{C}$ , small and large characteristic  
17 sizes can be chosen to give the same values of  $\omega_0$  for the same values of  $M_2$  assumed  
18 above. The above figures illustrate that there is a choice to be made as to how the bulk  
19 ice optical properties are parameterized in climate models.

20 As is done in B014b and in this paper, we avoid relating the bulk ice optical  
21 properties to  $D_{\text{eff}}$  as is traditionally done. Instead, our choice is to relate the optical  
22 properties directly to climate model prognostic variables such as  $M_2/\text{IWC}$ . This is because  
23 in climate models,  $D_{\text{eff}}$  is generally diagnosed in the radiation scheme and the  $D_{\text{eff}}$  in the

1 radiation scheme is different from the  $D_{\text{eff}}$  in the cloud physics scheme of a climate model  
2 because different PSDs are usually assumed. In the approach adopted by B014b and in  
3 this paper, we ensure that the PSDs generated through the F07 parameterization in the  
4 cloud physics scheme in the climate model are the *same* as those used in the radiation  
5 scheme, thus providing internal physical consistency within the climate model. In the  
6 traditional approach, this physical internal consistency is generally currently lost. Internal  
7 physical consistency could be satisfied using the traditional approach if  $D_{\text{eff}}$  were  
8 determined from the PSDs within the cloud physics scheme of the climate model and  
9 these values passed to the radiation scheme. In this way, the effective size then becomes  
10 consistent between the two schemes. However, this recovery of physical consistency  
11 requires an extra step within the climate model. Here, we prefer to avoid such an  
12 unnecessary step. It should also be noted that from Eq. (1), the integral over the mass  
13 requires some mass–D relationship to be applied. However, this mass–D relationship will  
14 obviously change depending on the choice of habit models; as a consequence, this has the  
15 potential to change the spatial distribution of  $\omega_0$  shown in Fig. 3a. However, in the  
16 approach adopted here, the mass of ice predicted by the climate model cannot change  
17 and, thus, will always be the same mass of ice between the cloud physics and radiation  
18 schemes.

19 In this paper, to generate the PSDs, the F07 moments are generated while assuming  
20 the recommended Furtado et al. (2015) mass–dimensional pre-factor and exponent values  
21 of 0.0257 and 2.0, respectively, which were derived by Cotton et al. (2013), and by using  
22 the 20662 IWC estimates and in-cloud temperature measurements compiled by B014a.  
23 Note that this mass–dimensional relationship is the same as that assumed by Hardiman et

1 al. (2015), and was used by Vidot et al. (2015) to find the optimal ensemble model optical  
2 property parameterization that best simulated globally measured infrared radiometric  
3 brightness temperatures. Furthermore, in this paper, the weights assigned to each of the  
4 ensemble members at each F07 PSD bin size are the same as those assumed in  
5 experiment 4 of B014b (i.e. Eq. (9) in that paper), which were 0.50, 0.20 and 0.30,  
6 respectively. This assignment of weights to the ensemble model is different from that  
7 found by Vidot et al. (2015) to achieve radiometric equivalence between the model and  
8 measurements. However, experiment 4 of B014b gave the best comparisons between the  
9 climate model runs and space-based shortwave and longwave radiation measurements. In  
10 contrast, the Vidot et al. (2015) analysis is based on measurements obtained at three  
11 wavelengths in the infrared. It is yet to be seen whether the weightings found in that  
12 paper would remain the same if more solar and infrared wavelengths were to be utilized.  
13 This will be the subject of a future paper; thus, we assume the same weightings as those  
14 used in experiment 4 of B014b.

15 In the MetUM configuration of models, atmospheric particulates such as aerosol, ice  
16 and water clouds are represented by vertical profiles of their mixing ratios with respect to  
17 air. In this paper and in B014b, the IWC becomes the ice mass mixing ratio,  $q_i$ , that is, the  
18 ratio between ice mass per unit volume and the mass of cloudy air per unit volume, and is  
19 in units of  $\text{kg kg}^{-1}$ . Likewise, the bulk optical properties, the volume extinction, and  
20 scattering coefficients become the mass extinction,  $K_{\text{ext}}$ , and mass scattering coefficients,  
21  $K_{\text{sca}}$ , and both are in units of  $\text{m}^2 \text{kg}^{-1}$  (i.e. the extinction and scattering coefficients per  
22 unit mass of cloudy air).

23 In this paper, the bulk optical property database is first divided into the six shortwave

1 and nine longwave Edwards–Slingo (1996, hereafter referred to as ES96) bands; these  
2 bands are defined in Table 1. In each of the ES96 bands, a total of about 10000  $q_i$  and  $T_c$   
3 values were randomly selected from the bulk optical property database of B014a.  
4 Selected values of  $q_i$  and  $T_c$  were between  $1.0 \times 10^{-8}$  and  $0.004 \text{ kg kg}^{-1}$  and  $-80^\circ\text{C}$  and  
5  $0^\circ\text{C}$ , respectively. This randomly selected range in  $q_i$ – $T_c$  space is sufficient to capture the  
6 full range that might be generated in an atmospheric model. In each of the ES96 bands,  
7 the bulk optical properties were parameterized as nonlinear and linear functions,  
8 dependent on  $q_i$  and  $T_c$ , by an iterative minimization procedure. That is, the forms of the  
9 parameterizations were first assumed and then the coefficients, for each of the  
10 parameterizations, were estimated by iterating through possible coefficient values. This  
11 was done until differences between the parameterizations and randomly selected 10000  
12 bulk optical properties were within acceptable experimental limits (i.e. estimated  $K_{\text{ext}}$  and  
13  $K_{\text{sca}}$  values must be within  $\pm 50\%$  of their true values for  $> 90\%$  of all possible  $q_i$ – $T_c$   
14 values). The error of  $\pm 50\%$  in the mass coefficients is based on likely in-situ errors in  
15 extinction estimates previously discussed in Baran et al. (2009). The errors resulting from  
16 this fitting procedure are discussed further below. The above fitting procedure resulted in  
17 the following best-fit bulk optical property parameterizations of cirrus:

$$18 \quad K_{\text{ext}}(\lambda, q_i, T_c) = a_\lambda (q_i / T_c^4) \quad (2)$$

$$19 \quad \omega_0(\lambda, q_i, T_c) = b_\lambda + c_\lambda q_i T_c \quad (3)$$

$$20 \quad g(\lambda, q_i, T_c) = d_\lambda + e_\lambda q_i T_c \quad (4)$$

21 where the temperature,  $T_c$ , is in units of Kelvin. The ES96 estimated band-dependent  
22 coefficients  $a_\lambda$ ,  $b_\lambda$ ,  $c_\lambda$ ,  $d_\lambda$  and  $e_\lambda$  are listed in Table 1. It should be noted that to obtain  
23 the correct asymptotic behavior for  $\omega_0(\lambda, q_i, T_c)$  and  $g(\lambda, q_i, T_c)$ , for the cases where  $q_i >$

1  $10^{-3} \text{ kg kg}^{-1}$ , the values of  $\omega_0$  and  $g$  should assume the same values as  $\omega_0$  and  $g$  when  
2  $q_i = 10^{-3} \text{ kg kg}^{-1}$ . If this condition is applied,  $\omega_0$  and  $g$  should never attain unphysical  
3 values. In B014b the above bulk optical properties are expressed as functions of  $q_i$   
4 only (i.e. Eqs. 4–7 in that paper). This is possible to do, as shown by Fig. 2, and is  
5 equivalent to the parameterization proposed by McFarquhar et al. (2003). In that  
6 paper, as previously stated, the IWC is explicitly derived as a function of  $r_{\text{eff}}$ ; in turn,  
7 the bulk optical properties are derived as a function of  $r_{\text{eff}}$ . Thus,  $r_{\text{eff}}$  can be eliminated  
8 between relationships to leave the bulk optical properties expressed as a function of  
9 IWC only, which is equivalent to B014b.

10 The difference between the above parameterization and that of B014b is  
11 highlighted by the following example. If we assume that  $T_c = 190 \text{ K}$  and  $q_i = 1.0 \times 10^{-3}$   
12  $\text{ kg kg}^{-1}$ , Eq. (3) at ES96 shortwave band 5 (1.19–2.38  $\mu\text{m}$ ) gives  $\omega_0 = 0.9015$ , whilst  
13 Eq. (6) from B014b gives  $\omega_0 = 0.8663$ . The former calculation is about 5% greater  
14 than the latter calculation, which means that in the former case the cirrus is about  
15 27% less absorbing (i.e.,  $\sim 1 - \omega_0$ ) than in the latter case. With such a difference in  
16 shortwave absorption between the two parameterizations, the B014b parameterization  
17 is clearly more likely to result in an increase in TTL temperature than Eq. (3).

18 In the introduction to this paper, it was noted that trigonal ice crystals were  
19 observed by a number of studies in the TTL. Single-scattering calculations based on  
20 trigonal ice crystals using the anomalous diffraction approximation (ADT) have been  
21 reported by Murray et al. (2015). These calculations show that  $\omega_0$ , computed  
22 assuming trigonal columns can be as much as 20% higher than their equivalent  
23 hexagonal column counterparts for a range of crystal sizes they considered at an

1 absorbing wavelength. If the TTL were composed mostly of trigonal columns, then  
2 this would imply less solar absorption and, as a consequence, less in-cloud heating,  
3 due to generally larger  $\omega_0$  values, relative to Eq. (3). However, ADT is a soft particle  
4 approximation (i.e. assumes real refractive indices near unity), originally due to van  
5 de Hulst (1957), and so by using this approximation,  $\omega_0$  values presented in Murray et  
6 al. (2015) may be overestimated (due to the neglect of reflection; refraction and  
7 particle edge effects, which all tend to increase absorption, see for instance, Mitchell  
8 et al. 2006). The latter limitations of ADT were noted by Murray et al. (2015), who  
9 also called for more accurate computations of the single-scattering properties of  
10 trigonal particles. On the other hand, if the TTL were composed of quasi-spherical  
11 particles as reported in some studies cited in the introduction to this paper, the  
12 calculated  $\omega_0$  values of these particles would tend to be lower (due to geometrical  
13 considerations and edge effects, see Baran and Havemann 1999; Mitchell et al. 2006;  
14 Bi and Yang, 2014) than that calculated for the equivalent trigonal particles. This  
15 implies more absorption, and so, greater in-cloud heating in the TTL, assuming equal  
16 microphysics assumptions. This range in potential TTL in-cloud heating is indicative  
17 of the current uncertainties in the microphysics composition within the TTL.  
18 Reducing uncertainty in calculated  $\omega_0$  values requires observations of PSDs, ice  
19 particle shapes, and application of electromagnetic methods and not gross  
20 approximations such as ADT to observed TTL microphysics measurements.

21 The accuracy of the new parameterizations was tested by selecting an independent  
22 set of about 10000  $q_i$  and  $T_c$  values taken from the ice optical property database of  
23 B014a. As before, the bulk optical properties, as a function of  $q_i$  and  $T_c$ , are grouped

1 into their ES96 bands. The test of accuracy of the new parameterizations is based on  
2 the relative percent error,  $\varepsilon(\lambda)$ , which given by:

$$3 \quad \varepsilon_{\lambda} = \frac{\text{true}_{\lambda} - \text{estimated}_{\lambda}}{\text{true}_{\lambda}} \times 100\% \quad (5)$$

4 where in Eq. (5)  $\text{true}_{\lambda}$  and  $\text{estimated}_{\lambda}$  are the actual bulk optical properties in the  
5 B014b database consisting of  $2.986 \times 10^6$  values (i.e. 20662  $q_i$ - $T_c$  values  $\times$  145  
6 wavelengths – 10000 randomly selected values) and the estimated values using Eqs.  
7 (2–4), respectively. Here, the accuracy of the parameterizations is illustrated using  
8 only ES96 shortwave band 5 (1.19–2.38  $\mu\text{m}$ ), as all other bands have similar  
9 accuracies, and this band is important for the shortwave heating of cirrus in the TTL.  
10 The calculated normalized PDFs of  $\varepsilon(\lambda)$  are shown for  $K_{\text{ext}}(\lambda, q_i, T_c)$ ,  $K_{\text{sca}}(\lambda, q_i, T_c)$  and  
11  $g$  in Figs. 4a–c, respectively. Figs. 4a and b show that  $\varepsilon(\lambda)$  in the mass  
12 parameterizations is within  $\pm 50\%$  for about 90% and 94% of the independently  
13 selected database, respectively. Typically, the in-situ measurement error in  
14  $K_{\text{ext}}(\lambda, q_i, T_c)$  is usually  $\pm 50\%$  (Baran et al. 2009), so the relative error distribution  
15 shown in Figs. 4a–b is acceptable. Fig. 4c shows  $\varepsilon(\lambda)$  for  $g$  and, as can be seen from  
16 the figure,  $\varepsilon(\lambda)$  is within  $\pm 2.5\%$  for about 83% of the database, which is also  
17 acceptable. Theoretical and in-situ uncertainty in the asymmetry parameter value is  
18 far greater than the error in the  $g$  parameterization (Ulanowski et al. 2006; Fu 2007;  
19 Garrett 2008; Baran 2012; van Diedenhoven et al. 2014). We compare our  
20 parameterization of  $g$  to the  $g$  parameterization developed by Fu (2007) by assuming  
21  $q_i$  and  $T_c$  values of  $1.0 \times 10^{-4} \text{ kg kg}^{-1}$  and 190 K, respectively. In Fu (2007),  $g$  is  
22 parameterized in the shortwave by assuming hexagonal columns and hexagonal plates



1 as a function of an effective AR by using either rough or smooth collections of  
2 particles. The range in the effective aspect ratio given by Fu (2007) represents the  
3 aspect ratios of monomer ice crystals that make up more complex ice crystals. Indeed,  
4 the effective AR of the ensemble model members assumed in this paper are 1.0, 0.42  
5 and 0.42 (the first being the hexagonal ice column of AR=1.0, and the latter two  
6 being the AR and averaged AR values found for the six-branched bullet rosette and  
7 the three monomer hexagonal ice aggregate, respectively), and the following weights  
8 of 0.50, 0.20 and 0.30 are assigned to the three ensemble members, respectively. The  
9 comparison is made by using the coefficients  $d_\lambda$  and  $e_\lambda$  tabulated in Table 1 at ES96  
10 shortwave band 1 (i.e. 0.20–0.32  $\mu\text{m}$ ) and at ES96 shortwave band 3 (0.32–0.69  $\mu\text{m}$ ).  
11 These two shortwave bands cover the wavelength range given in Fu (2007), for  
12 shortwave band 1 (i.e. 0.25–0.70  $\mu\text{m}$ ) given in that paper. Although the comparison is  
13 not exact in terms of the wavelength range and assumed ice microphysics, it is  
14 considered sufficiently accurate for the purposes of this paper. Applying the above  
15 three effective AR ratios for each of the ensemble models to Eqs. (3.2 and 2.2) given  
16 in Fu (2007), and weighting the calculated asymmetry parameter values by the three  
17 weights given above, we find a weighted value of 0.765 for  $g$  (i.e.  
18  $0.5 \times 0.73 + 0.20 \times 0.80 + 0.3 \times 0.80$ ). This Fu (2007) asymmetry parameter value  
19 compares to the values of 0.757 and 0.787 found for ES96 shortwave bands 1 and 3,  
20 respectively, and the average of these two values is 0.772. The averaged value is  
21 within about 1% of the Fu (2007) calculation, and the values calculated at the two  
22 ES96 shortwave bands are within about  $\pm 2\%$  of Fu (2007). If we take  $g=0.765$  and  
23  $g=0.787$  (i.e. the most extreme difference) and assume that the cirrus is located over a

1 dark ocean and conservative scattering, then the backward solar reflection is about  
2 23.5 and 21.3%, respectively, where backward reflection  $\sim (1.-g)$ . The difference in  
3 reflected shortwave flux is at most about  $7 \text{ Wm}^{-2}$ , assuming an area-averaged incident  
4 solar flux of  $330 \text{ Wm}^{-2}$ , at least for the case considered here. These differences found  
5 for the  $g$  parameterizations are far less than the range in  $g$  found by the following  
6 studies (Ulanowski et al. 2006; Fu 2007; Garrett 2008; Baran 2012; van Dierenhoven  
7 et al. 2014, Yang et al. 2015). Indeed, in the case of Ulanowski et al. (2006),  
8 experimentally derived  $g$  values were found to vary between  $0.80 \pm 0.04$  and  $0.63 \pm 0.05$   
9 for smooth and rough ice analog rosettes, respectively, and this difference results in a  
10 shortwave flux uncertainty of about  $-56 \text{ Wm}^{-2}$ . Furthermore, the calculated  
11 asymmetry parameter values using the two parameterizations compare well against  
12 radiometrically derived asymmetry parameter values using POLDER observations  
13 from van Dierenhoven et al. (2014). The observations from POLDER were located  
14 over the north coast of Australia. In that paper, at a  $T_c$  value of about  $-85^\circ\text{C}$ , the  
15 asymmetry parameter values are shown to vary between about  $0.74 \pm 0.02$  and  
16  $0.79 \pm 0.04$ , which encompass the values of  $g$ , derived using the two completely  
17 different parameterizations. These experimental results obtained at  $0.865 \mu\text{m}$ , were  
18 roughly coincident with the three assumed shortwave bands used in the comparisons,  
19 and the imaginary index of ice is very weakly absorbing at all of the chosen  
20 shortwave bands as shown in B014b. At least at TTL temperatures, the parameterized  
21  $g$  values derived in this paper appear representative of other independent derivations  
22 of  $g$ . However,  $g$  is one of the bulk ice optical properties that is important to constrain  
23 in climate models, as noted by Stephens et al. (1990). In this paper, we are concerned

1 about TTL heating and specific humidity errors in a climate model, and these errors  
2 are more associated with the parameterizations found for  $K_{\text{ext}}(\lambda, q_i, T_c)$  and  $\omega_0(\lambda, q_i, T_c)$ .

3 Eqs. (2) and (3) are now compared against the equivalent parameterizations  
4 developed by B014b. The parameterizations of  $g$  are not compared here as the results are  
5 not sufficiently different. In comparing the parameterizations a constant value of  $q_i$  is  
6 assumed, with its value being  $1.0 \times 10^{-5} \text{ kg kg}^{-1}$ , whilst the temperature is allowed to vary  
7 between  $-80^\circ\text{C}$  and  $0^\circ\text{C}$ . The comparisons are shown in Figs. 5a–c at ES96 shortwave  
8 band 5 for  $K_{\text{ext}}(\lambda, q_i, T_c)$ , the co-albedo (i.e.,  $1 - \omega_0(\lambda, q_i, T_c)$ ), and (c)  $K_{\text{abs}}(\lambda, q_i, T_c)$  (i.e.,  
9  $K_{\text{abs}}(\lambda, q_i, T_c) = K_{\text{ext}}(\lambda, q_i, T_c) - K_{\text{sca}}(\lambda, q_i, T_c)$ ), respectively. Other bands show generally  
10 similar results to Figs. 5a–c and are not shown here for reasons of brevity. Fig. 5a shows  
11 that at  $T_c$  values of about  $-80^\circ\text{C}$ , the nonlinear parameterization of  $K_{\text{ext}}(\lambda, q_i, T_c)$  is about a  
12 factor of 2.5 greater than that predicted by B014b. This means that at typical TTL  
13 temperatures, the new parameterization, relative to B014b, will transmit less longwave  
14 terrestrial radiation to space through the cirrus, which will result in less longwave  
15 absorption by trace gases above the TTL cirrus, consequently lowering TTL temperatures  
16 through emission at cold temperatures. Also critical to the TTL temperature is the co-  
17 albedo. The co-albedo comparisons are shown in Fig. 5b; again, the figure shows that at  
18  $T_c = -80^\circ\text{C}$ , the new parameterization co-albedo is less than B014b by almost a factor of  
19 7. This change in absorption between the parameterizations is shown clearly in Fig. 5c,  
20 which shows  $K_{\text{abs}}(\lambda, q_i, T_c)$ , and at  $T_c = -80^\circ\text{C}$ , the new parameterization is almost a factor  
21 3 times less absorbing than B014b. The comparisons show that the temperature-  
22 dependent parameterization in the TTL should result in less cirrus heating relative to  
23 B014b. The next section presents the impact of these different heating profiles on the

1 simulation of the TTL in an atmosphere-only climate integration using the MetUM.

### 2 **3. The impact of the parameterizations on the simulation of the TTL**

3 To assess the fidelity of the TTL simulation in models using the B014b parametrization  
4 and the parametrization described in section 2, we perform a pair of 20-year atmosphere-  
5 only climate simulations using each parametrisation. Apart from the formulation of the  
6 ice cloud optical properties, these simulations each use an identical baseline of the Global  
7 Atmosphere 6.0 (GA6.0) configuration of the MetUM. GA6.0 is described by Walters et  
8 al. (2015), so a description is not repeated here. The simulations are performed at a  
9 horizontal resolution of N96 (~135 km in the mid-latitudes) and use a vertical level set  
10 with spacings of between 500 and 700 m in the TTL region with a model “lid” at 85 km  
11 from the surface. The simulations run from December 1988 to November 2008 and use  
12 prescribed (but time-varying) sea-surface temperatures, greenhouse gas concentrations,  
13 and aerosol emissions while roughly following the protocol of the Atmospheric Model  
14 Intercomparison Project (AMIP, Taylor et al. 2012). Results are presented for the  
15 meteorological season December–January–February (DJF), as this season represents the  
16 period during which cirrus occurrence in the tropics is at a maximum (Sassen et al. 2008).  
17 The climate model predictions are compared against the ERA-Interim reanalysis  
18 temperature product (Dee et al. 2011) and the Modern-ERA Retrospective analysis for  
19 Research and Applications (MERRA) specific humidity product (Rienecker et al. 2011).  
20 The impacts of the parameterizations on the 20-year averaged DJF cloudy shortwave and  
21 longwave radiative effects at TOA are compared against the Loeb et al. (2009) reanalysis  
22 of the Clouds and the Earth’s Radiant Energy System (CERES) product.

1 Firstly, the DJF zonally averaged temperature differences between the B014b  
2 parameterization and the ERA-Interim reanalysis are shown in Fig. 6a, whilst the same  
3 differences are shown in Fig. 6b but for the new parameterizations. It is clear from Fig.  
4 6b, relative to Fig. 6a, that the new parameterizations have reduced the TTL temperature  
5 by about 1 K, and generally throughout the tropics, which is consistent with the  
6 discussion surrounding Figs. 5a and b. In addition, the new parameterization has also  
7 removed the warming in the southern sub-tropical troposphere region by about 1–2 K,  
8 and reduced cooling and warming over the South and North Poles, respectively. On the  
9 negative side, the new parameterizations lead to a cooling of the tropical troposphere by  
10 about 1 K, and to a similar cooling at altitudes between about 15 and 30 hPa, which  
11 occurs near the equator.

12 The impact of the new parameterizations on the model bias of zonally averaged  
13  $\log_{10}$  (specific humidity), relative to MERRA, is shown in Fig. 6d. The figures show that  
14 the impact of the new parameterizations on the  $\log_{10}$  (specific humidity) model bias in the  
15 TTL is to reduce it, and its distribution throughout the TTL, especially around the  
16 equatorial region. Relative to the B014b parameterization, shown in Fig. 6c, the new  
17 parameterization reduces the  $\log_{10}$  (specific humidity) model bias in the TTL by about 0.1  
18  $\log_{10}$  ( $\text{kg kg}^{-1}$ ) and generally reduces the extent of this bias in the sub-tropics. For each  
19 panel shown in Figs. 6a–d, the grid point root mean square errors (RMS) were calculated  
20 over the region between 20°S and 20°N, and between 150 and 70 hPa. The resulting RMS  
21 errors were found to be 2.52, 1.90 K, and 0.13 and 0.12  $\log_{10}$  ( $\text{kg kg}^{-1}$ ), respectively.  
22 Therefore, in the TTL region, the new parameterization reduces the zonally averaged  
23 temperature and specific humidity biases in the model.

1 Finally, to show that the new parameterizations do not result in any detriment to the  
2 climate model, in terms of the cloud radiative effect at TOA and global means. Results  
3 are presented for the model's predicted radiative effects in Figs. 7a–d and global means  
4 in Table 2. The figures show that the new parameterization improves the shortwave and  
5 longwave cloud radiative effects in the climate model, relative to B014b. The extent of  
6 the shortwave bias in the model is significantly reduced, as shown by the reduction in the  
7 area-weighted RMS error by  $1.08 \text{ Wm}^{-2}$ , and this improvement is especially evident in the  
8 tropics. Note also the reduction in bias brought about by the new parameterization  
9 throughout the Southern Ocean. The longwave biases in the climate model are also  
10 reduced by the new parameterizations as shown in Figs. 7c and 7d, where it can be seen  
11 that the area-averaged RMS error is reduced by  $0.41 \text{ Wm}^{-2}$ . The significant longwave  
12 negative biases around the warm pool shown in Fig. 7c have been reduced by the new  
13 parameterizations by up to about  $20 \text{ Wm}^{-2}$ , as demonstrated by Fig. 7d.

14 Table 2 shows a comparison between the Stephens et al. (2012) estimated CERES  
15 global means for seven variables together with their estimated standard deviation  $\pm\sigma$ , and  
16 the B014b and temperature-dependent parameterizations. The table shows that the  
17 temperature-dependent parameterization is within the estimated uncertainty for six out of  
18 seven of the global mean variables shown in table 2. This improves on the B014b  
19 parameterization, which has four out of seven variables within the estimated uncertainty.  
20 In particular, the B014b parameterization significantly underestimates the reflected  
21 shortwave flux at TOA. As a result of this underestimation, the predicted shortwave  
22 radiative effect is too low relative to the CERES estimated mean. The temperature-  
23 dependent parameterization does predict adequate reflected shortwave flux at TOA, but

1 the prediction of the shortwave radiative effect is just over  $1\sigma$  from the CERES estimated  
2 mean. This compares to about  $1\frac{3}{4}\sigma$  from the estimated mean found for the B014b  
3 parameterization. The reason for this discrepancy between the parameterizations and the  
4 CERES observations for this variable is probably due to, in the case of B014b; the mass  
5 extinction coefficient being underestimated at about temperatures  $< -40^{\circ}\text{C}$ , as can be seen  
6 from Fig. 5a, relative to the temperature-dependent parameterization. In the case of the  
7 temperature-dependent parameterization, the physical reasons for this underestimation  
8 could be due to: (i) There is too little ice mass being predicted by the model cloud scheme  
9 and/or (ii) the model asymmetry parameter is too large. It is yet to be seen which of these  
10 reasons might account for the discrepancy in the shortwave radiative effect. However, in  
11 general, the new parameterization presented within this paper improves the model  
12 performance, relative to B014b, in terms of the shortwave and longwave cloudy radiative  
13 effect as can be seen from Figs. 7b–d, but also the predictions of model global mean  
14 fluxes.

#### 15 **4. Conclusions**

16 A cirrus bulk optical property parameterization has been presented, and it has been  
17 demonstrated that the parameterization reduces zonally averaged temperature biases in  
18 the TTL of GA 6.0 by about 1–2 K relative to the B014b parameterizations. The  
19 parameterization of the asymmetry parameter was compared against an independent  
20 parameterization and was shown to be within about 1% and  $\pm 2\%$  of the latter  
21 parameterization at ES96 short wavebands 1 and 2, respectively. Moreover, the  
22 asymmetry parameters predicted by both parameterizations were shown to encompass  
23 space-based estimations of  $g$  from observations from around the north coast of Australia

1 at TTL cirrus temperatures of  $-85^{\circ}\text{C}$ . The reason for this improvement in climate model  
2 performance is through coupling the bulk optical properties to  $q_i$  and  $T_c$ . At the low  
3 temperatures in the TTL, the B014b parameterization will under-predict the single-  
4 scattering albedo, resulting in TTL warming due, in part, to the absorption of shortwave  
5 radiation, highlighting the importance of wavelengths in the near-infrared. The new  
6 temperature-dependent parameterization also improves, relative to B014b, on  
7 representing the model's prediction of the shortwave and longwave cloudy radiative  
8 effect as well as global flux means. To improve climate model representations of the  
9 TTL, more careful consideration should be given to the calculation of ice optical  
10 properties, especially the mass extinction coefficients and single-scattering albedo, which  
11 are the bulk ice optical properties that principally determine the temperature and specific  
12 humidity distributions in our simulations of TTL cirrus. Further observations of cirrus  
13 PSDs, IWCs, habits (in addition to humidity and temperature measurements), and solar  
14 and infrared radiative properties in the TTL should be considered a priority.

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1 Fig. 1. The bulk optical properties (a)  $\omega_0$  and (b)  $g$  as a function of the mass carrying  
2 moment  $M(2)$  and temperature at the wavelength of  $1.575 \mu\text{m}$  for all 20662  
3 values. The calculated values for  $\omega_0$  and  $g$  are shown as the color bar on the right-  
4 hand side of the figures.

5 Fig. 2. The mass carrying moment  $M(2)$  as a function of the mean effective dimension,  
6  $D_{\text{eff}}$ , and temperature,  $T_c$ , for all 20662 values. The key on the right-hand side of  
7 the figure is  $T_c$  in units of  $^\circ\text{C}$ .

8 Fig. 3. The 20662 calculated values of  $\omega_0$  as a function of (a)  $D_{\text{eff}}$  and (b)  $M3/M2$ , the  
9 mean mass-weighted size,  $D_{\text{mmw}}$ , of the PSD. The key on the right-hand side of  
10 the figures is the calculated values of  $\omega_0$  at the wavelength of  $1.575 \mu\text{m}$ .

11 Fig. 4. The normalized PDFs of the relative percent error in the parameterization of (a)  
12  $K_{\text{ext}}(\lambda, q_i, T_c)$ , (b)  $K_{\text{sca}}(\lambda, q_i, T_c)$ , and (c)  $g(\lambda, q_i, T_c)$ . Relative percent error results  
13 are shown for Edwards and Slingo (1996) shortwave band 5 ( $1.19\text{--}2.38 \mu\text{m}$ ).

14 Fig. 5. Comparing the new parameterization (full line) to the B014b parameterization  
15 (dashed line) at Edwards and Slingo (1996) shortwave band 5 ( $1.19\text{--}2.38 \mu\text{m}$ ).  
16 Comparisons are shown for (a)  $K_{\text{ext}}(\lambda, q_i, T_c)$ , (b) the co-albedo, and (c)  $K_{\text{abs}}(\lambda, q_i,$   
17  $T_c)$ .

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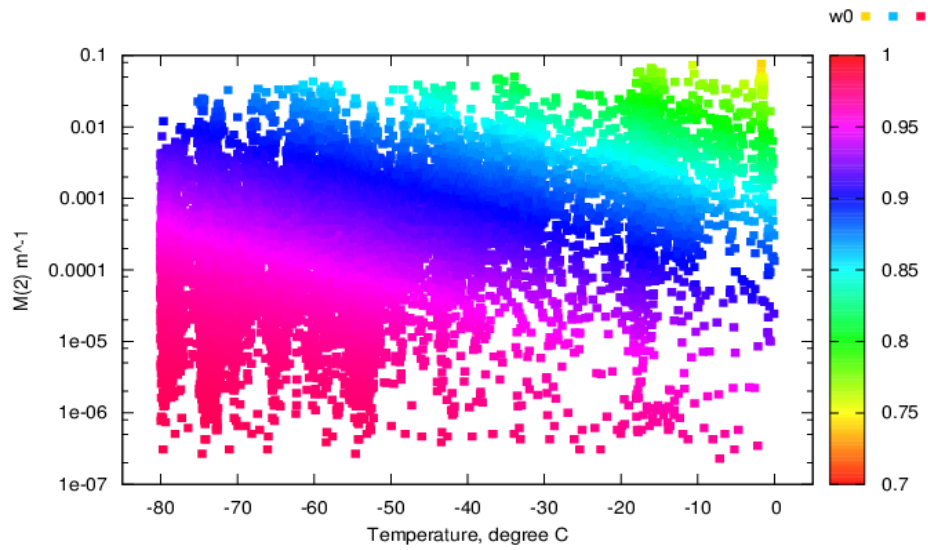
1 Fig. 6. The 20-year DJF zonally averaged temperature and  $\log_{10}$  (specific humidity)  
2 differences, shown in the top and bottom panels, respectively. The differences are  
3 between the MetUM configuration 6.0 run, the ERA-Interim and MERRA  
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8 respectively.

9 Fig. 7. The annual 20-year TOA cloud radiative effect differences between MetUM  
10 configuration 6.0 and the CERES products. Results are shown for the shortwave  
11 while assuming the (a) B014b parameterization and (b) the new parameterization.  
12 The longwave results are shown while assuming the (c) B014b parameterization  
13 and (d) the new parameterization. The area-averaged root mean square difference  
14 is shown in each of the panels, and differences are in units of  $\text{Wm}^{-2}$ .

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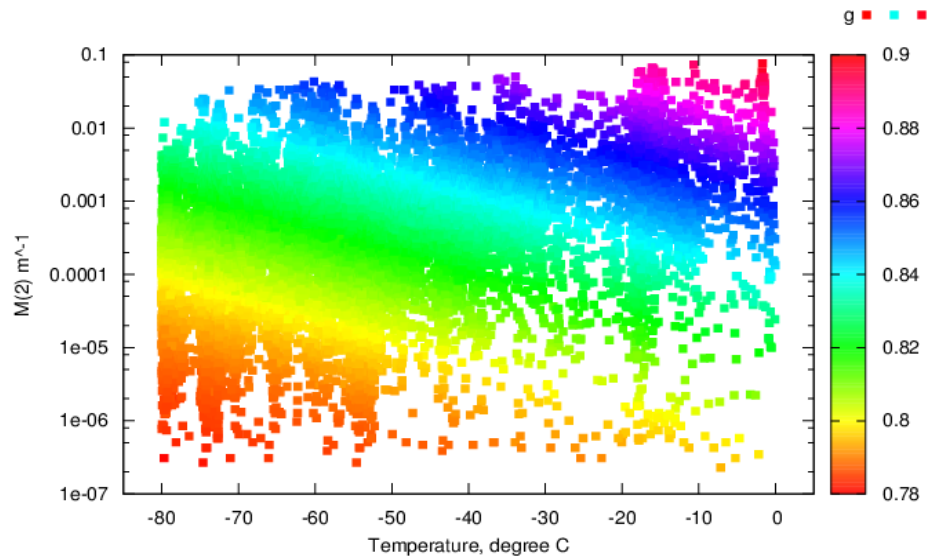
(a)



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(b)



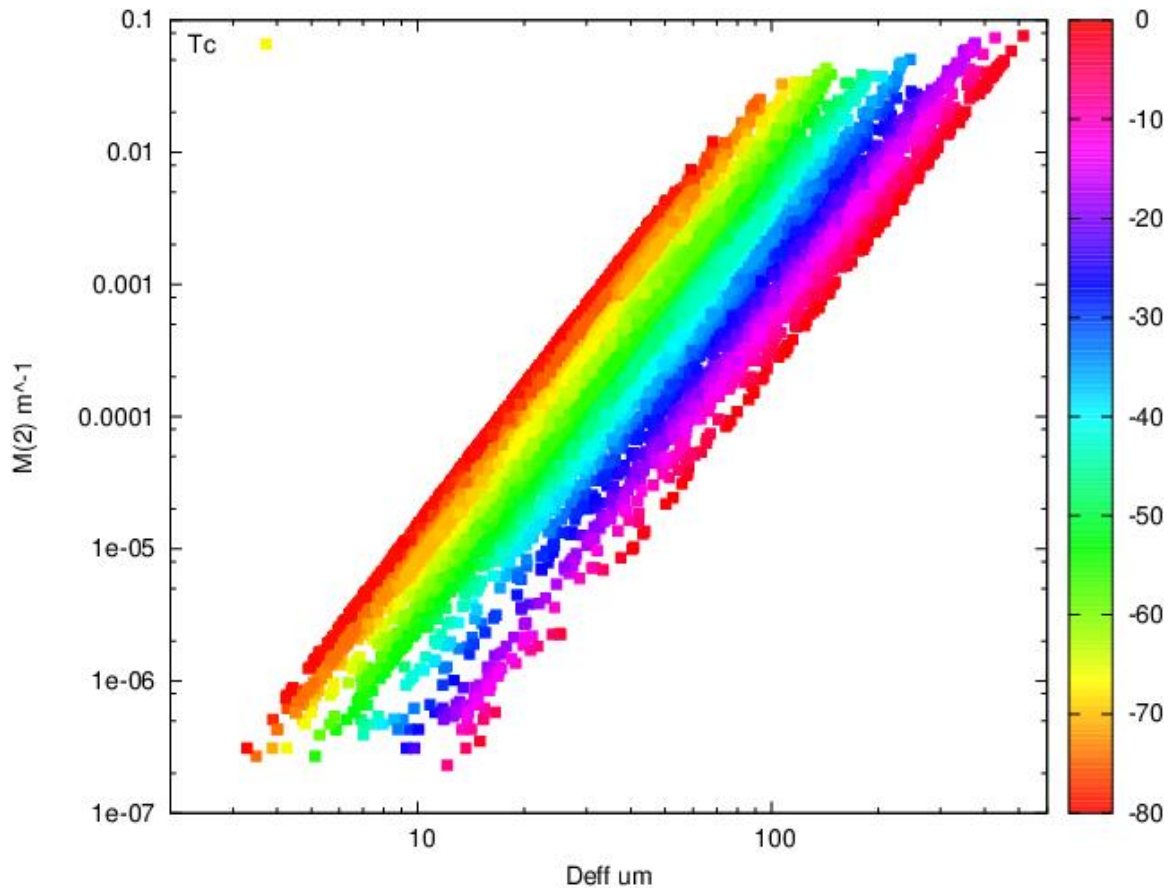
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5 Fig. 1. The bulk optical properties (a)  $w_0$  and (b)  $g$  as a function of the mass carrying  
6 moment  $M(2)$  and temperature at the wavelength of  $1.575 \mu\text{m}$  for all 20662  
7 values. The calculated values for  $w_0$  and  $g$  are shown as the color bar on the right-  
8 hand side of the figures.

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3 Fig. 2. The mass carrying moment  $M(2)$  as a function of the mean effective dimension,  
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5 the figure is  $T_c$  in units of  $^{\circ}\text{C}$ .

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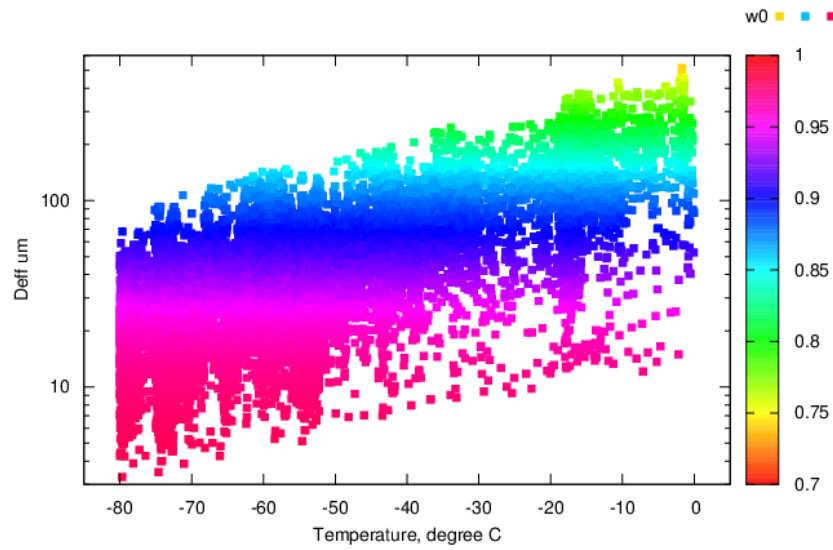
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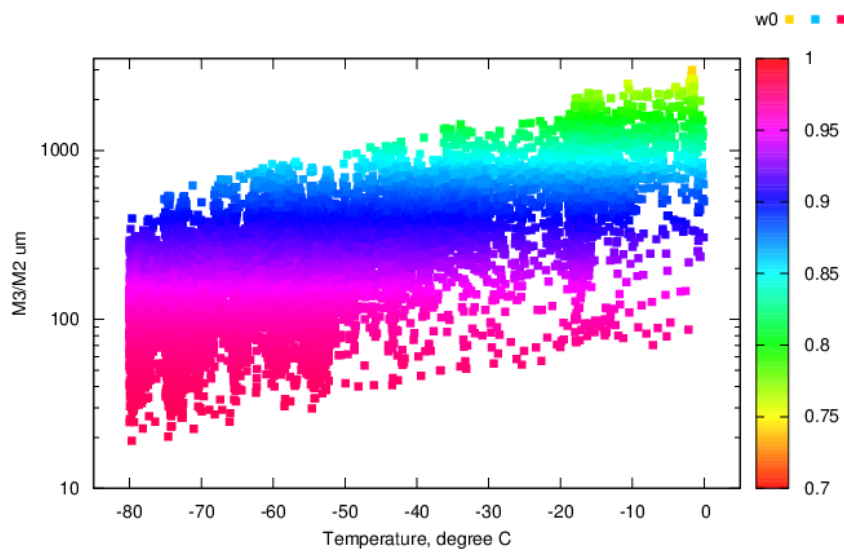
(a)



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(b)



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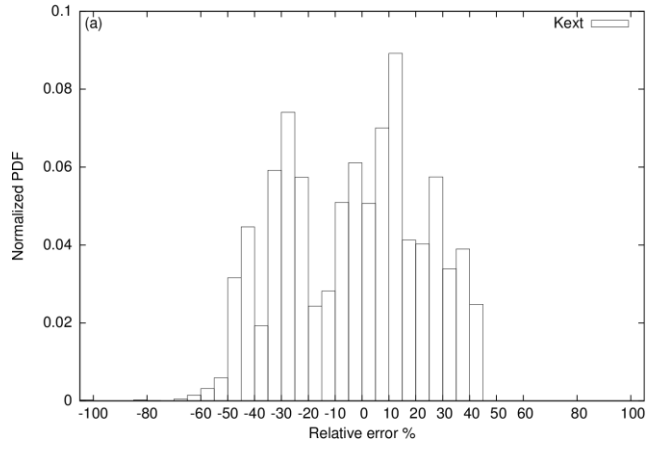
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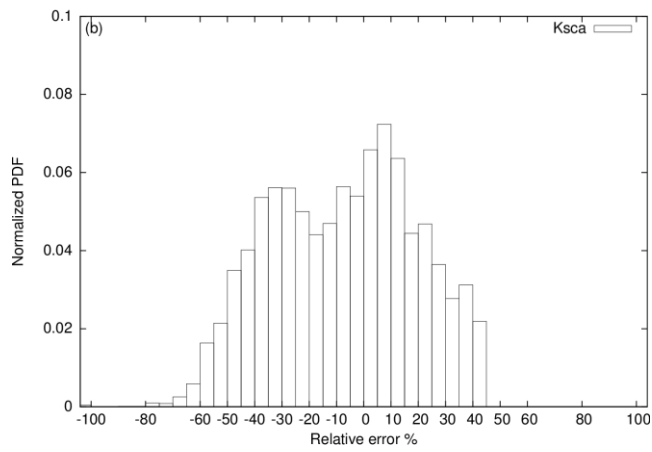
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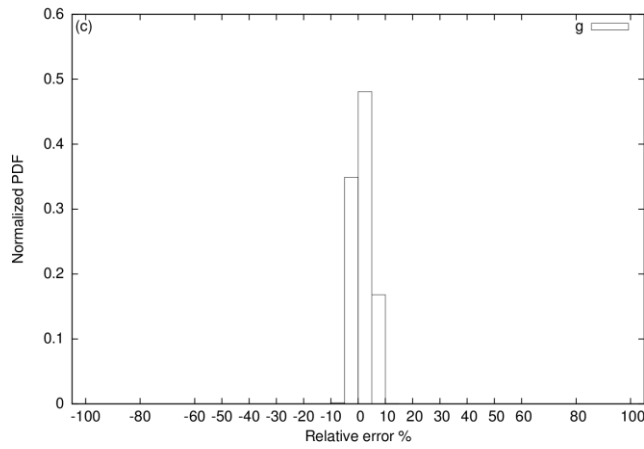
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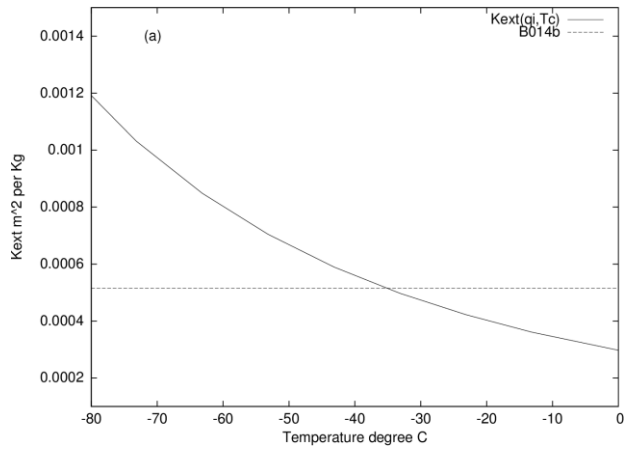
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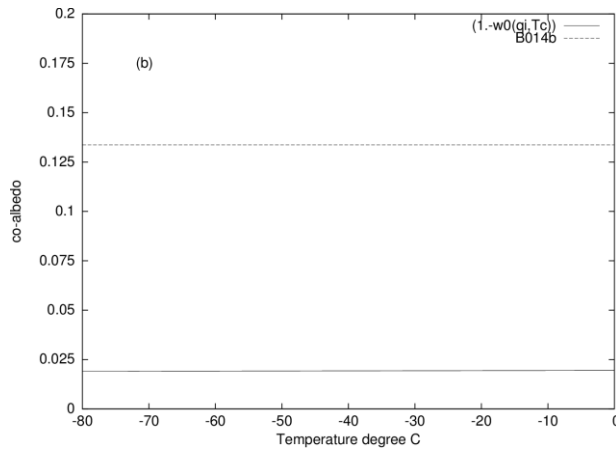
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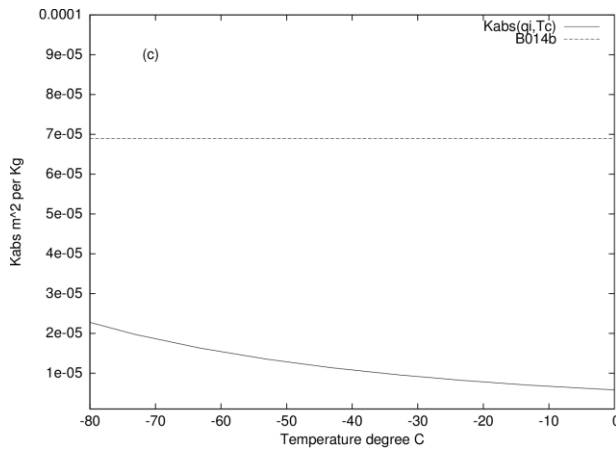
4 Fig. 4. The normalized PDFs of the relative percent error in the parameterization of (a)  
5  $K_{ext}(\lambda, q_i, T_c)$ , (b)  $K_{sca}(\lambda, q_i, T_c)$ , and (c)  $g(\lambda, q_i, T_c)$ . Relative percent error results  
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4 Fig. 5. Comparing the new parameterization (full line) to the B014b parameterization

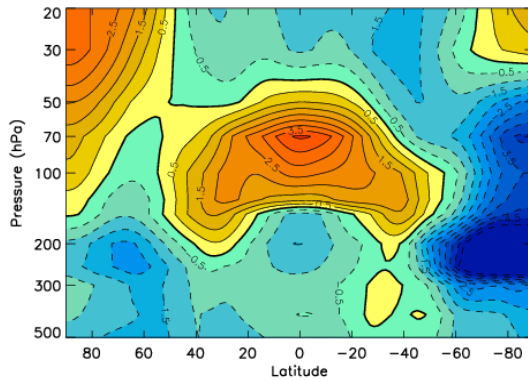
5 (dashed line) at Edwards and Slingo (1996) shortwave band 5 (1.19–2.38  $\mu\text{m}$ ).

6 Comparisons are shown for (a)  $K_{\text{ext}}(\lambda, q_i, T_c)$ , (b) the co-albedo,  $(1.-\omega_0(\lambda, q_i, T_c))$ ,

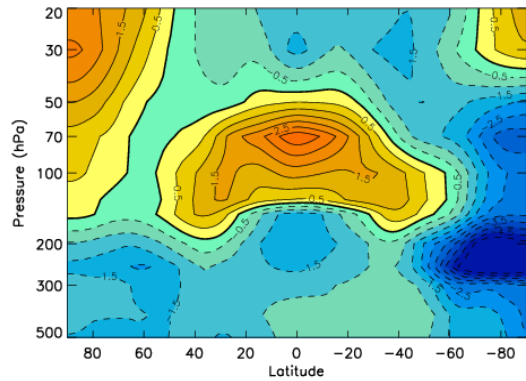
7 and (c)  $K_{\text{abs}}(\lambda, q_i, T_c)$ .

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(a)



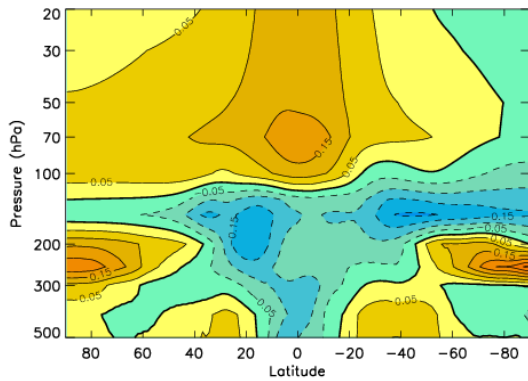
(b)



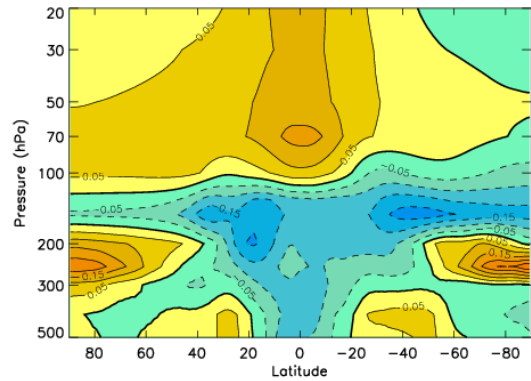
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(c)



(d)



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5 Fig. 6. The 20-year DJF zonally averaged temperature and log<sub>10</sub> (specific humidity)

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differences, shown in the top and bottom panels, respectively. The differences are

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between the MetUM configuration 6.0 run, the ERA-Interim and MERRA

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while assuming the B014b parameterization. In the right column, panels (b) and

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(d), results are shown while assuming the new parameterization. The units of

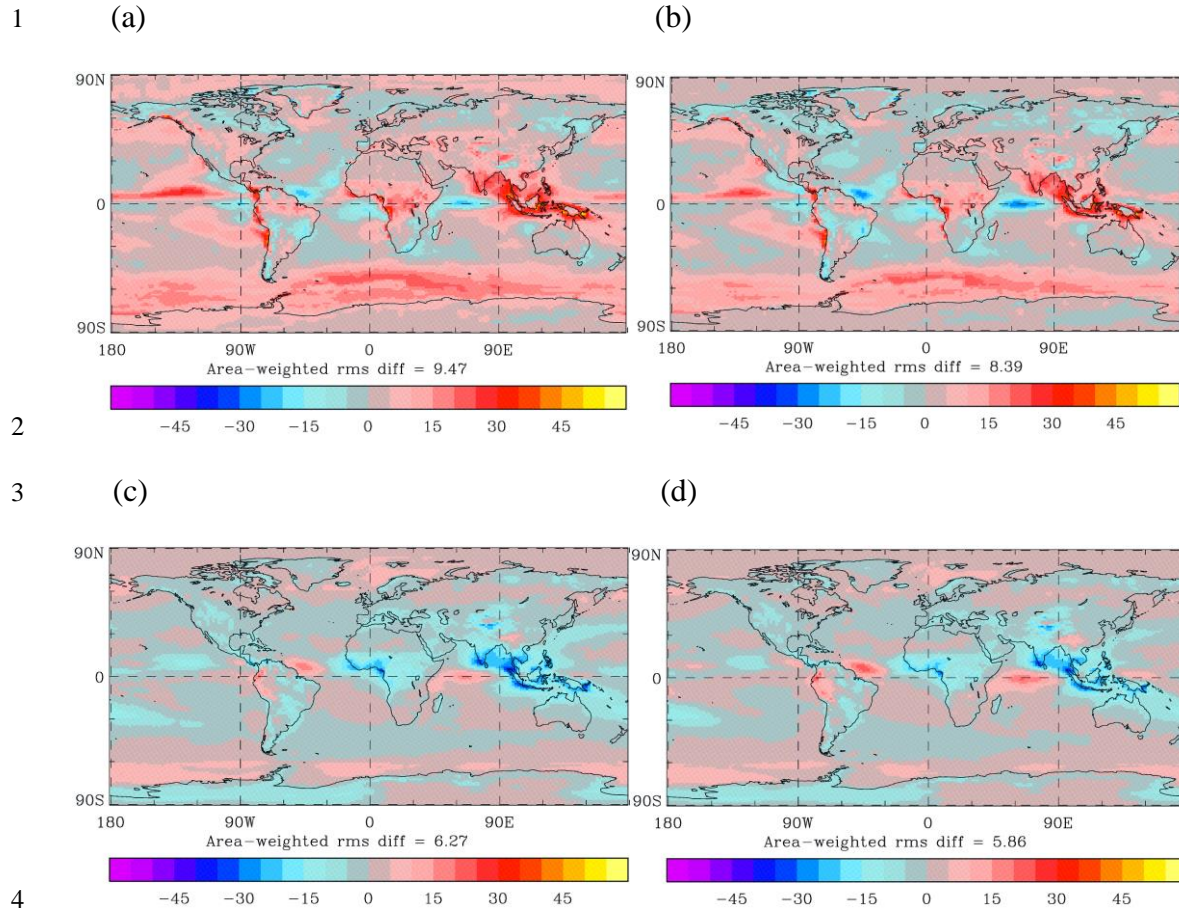
11

temperature difference and log (specific humidity) are K and log<sub>10</sub> (kg kg<sup>-1</sup>),

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respectively.

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5 Fig. 7. The annual 20-year TOA cloud radiative effect differences between MetUM  
6 configuration 6.0 and the CERES products. Results are shown for the shortwave  
7 while assuming the (a) B014b parameterization and (b) the new parameterization.  
8 The longwave results are shown while assuming the (c) B014b parameterization  
9 and (d) the new parameterization. The area-averaged root mean square difference  
10 is shown in each of the panels, and differences are in units of  $\text{Wm}^{-2}$ .

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1 Table 1. The derived values of the band-dependent coefficients for each of the  
 2 Edwards and Slingo (1996) shortwave and longwave bands. The first six entries in the  
 3 table are the shortwave bands and the following nine entries are the longwave bands.

Band $\mu\text{m}$	$a_\lambda$	$b_\lambda$	$c_\lambda$	$d_\lambda$	$e_\lambda$
0.20–0.32	$1.64 \times 10^{11}$	1.0000	$6.5 \times 10^{-15}$	0.7560	0.0378
0.32–0.69	$1.66 \times 10^{11}$	1.0000	$-1.4 \times 10^{-6}$	0.7804	0.0419
0.32–0.69	$1.64 \times 10^{11}$	0.9999	$-6.4 \times 10^{-5}$	0.7860	0.0434
0.69–1.19	$1.64 \times 10^{11}$	0.9996	-0.0031	0.7897	0.0468
1.19–2.38	$1.65 \times 10^{11}$	0.9817	-0.4218	0.8208	0.1351
2.38–10.0	$1.60 \times 10^{11}$	0.7500	-0.2353	0.9130	0.1240
25.0– $10^4$	$1.50 \times 10^{11}$	0.6300	-0.2500	0.7129	0.6568
18.18–25.0	$1.75 \times 10^{11}$	0.7700	-0.2866	0.8356	0.3085
12.50– 18.18	$1.73 \times 10^{11}$	0.5480	0.0108	0.8843	0.1833
13.33– 16.95	$1.74 \times 10^{11}$	0.5469	0.0146	0.8845	0.1730
8.33–12.50	$1.44 \times 10^{11}$	0.5346	0.0106	0.9338	0.1013
8.93–10.10	$1.54 \times 10^{11}$	0.6500	-0.1420	0.9401	0.1144
7.52–8.33	$1.70 \times 10^{11}$	0.6000	-0.1781	0.9264	0.1513
6.67–7.52	$1.72 \times 10^{11}$	0.6000	-0.1781	0.9331	0.1287
3.34–6.67	$1.71 \times 10^{11}$	0.6009	-0.1304	0.9080	0.1762

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1 Table 2. CERES means obtained from Stephens et al. (2012) compared against the  
 2 MetUM configuration 6.0 means predicted using the B014b parameterization and the  
 3 temperature-dependent parameterization. All flux units (SW and LW) are in  $\text{Wm}^{-2}$ .  
 4 Asterisks denote that the predicted variable is within the current CERES measurement  
 5 uncertainty.

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Variable	Observation	B014b	Temperature-dependent
Outgoing SW	100±2	97.58	99.66*
Outgoing LW	239.7±3.3	240.53*	239.52*
Absorbed SW	240.2±2	242.80	240.73*
Albedo TOA	29.41±1	28.67*	29.28*
SW CF	-47.5±3	-42.18	-44.23
LW CF	26.7±4	23.88*	24.74*
Net surf SW	165±6	168.64*	170.92*

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