

The role of the cloud radiative effect in the sensitivity of the Intertropical Convergence Zone to convective mixing

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1 **The role of the cloud radiative effect in the sensitivity of the Intertropical**
2 **Convergence Zone to convective mixing.**

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ABSTRACT

12 Studies have shown that the location and structure of the simulated Intertrop-
13 ical Convergence Zone (ITCZ) is sensitive to the treatment of sub-gridscale
14 convection and cloud-radiation interactions. This sensitivity remains in ide-
15 alised aquaplanet experiments with fixed surface temperatures. However,
16 studies have not considered the role of cloud-radiative effects (CRE, atmo-
17 spheric heating due to cloud-radiation interactions) in the sensitivity of the
18 ITCZ to the treatment of convection. We use an atmospheric energy input
19 (AEI) framework to explore how the CRE modulates the sensitivity of the
20 ITCZ to convective mixing in aquaplanet simulations. Simulations show a
21 sensitivity of the ITCZ to convective mixing, with stronger convective mixing
22 favoring a single ITCZ. For simulations with a single ITCZ, the CRE main-
23 tains the positive, equatorial AEI. To explore the role of the CRE further, we
24 prescribe the CRE as either zero or a meridionally and diurnally varying cli-
25 matology. Removing the CRE is associated with a reduced equatorial AEI
26 and an increase in the range of convective mixing rates that produce a double
27 ITCZ. Prescribing the CRE reduces the sensitivity of the ITCZ to convective
28 mixing by 50%. In prescribed-CRE simulations, other AEI components, in
29 particular the surface latent heat flux, modulate the sensitivity of the AEI to
30 convective mixing. Analysis of the meridional moist static energy transport
31 shows that a shallower Hadley circulation can produce an equatorward energy
32 transport at low latitudes even with equatorial ascent.

33 **1. Introduction**

34 Tropical rainfall is often associated with a discontinuous zonal precipitation band commonly
35 known as the Intertropical Convergence Zone (ITCZ). The ITCZ migrates between the Northern
36 and Southern Hemispheres with the seasonal cycle, with a zonal-, time-mean position of approx-
37 imately 6°N (Schneider et al. 2014). The ITCZ is co-located with the ascending branch of the
38 Hadley circulation, where strong moist convection leads to high rainfall. The upper branches of
39 the Hadley circulation typically transport energy poleward, away from the ITCZ. Recent studies
40 have associated characteristics of the ITCZ with the energy transport by the Hadley circulation
41 (Frierson and Hwang 2012; Donohoe et al. 2013; Adam et al. 2016; Bischoff and Schneider
42 2016).

43 A double ITCZ bias is prominent in current and previous generations of coupled general
44 circulation models (GCMs; Li and Xie 2014; Oueslati and Bellon 2015). The ITCZ is too
45 intense in the Southern Hemisphere (Lin 2007), resulting in two annual-, zonal-mean tropical
46 precipitation maxima, one in each hemisphere. A bias remains in atmosphere-only simulations
47 with prescribed sea surface temperatures (SSTs) (Li and Xie 2014). Aquaplanet simulations
48 provide an idealised modelling environment in which some complex boundary conditions in
49 tropical circulation such as land/sea contrasts and orography are removed. However aquaplanet
50 configurations of GCMs coupled to a slab ocean produce a broad range of tropical precipitation
51 mean states (Voigt et al. 2016); even prescribing zonally uniform SSTs does not resolve the
52 inter-model variability (Blackburn et al. 2013).

53

54 *a. Modelling studies*

55 Characteristics of the simulated ITCZ are sensitive to the representation of cloud-radiation inter-
56 actions (Fermepin and Bony 2014; Li et al. 2015; Harrop and Hartmann 2016). In the deep tropics
57 the cloud radiative effect (CRE) warms the atmosphere (Allan 2011), with important effects on
58 tropical circulation (Slingo and Slingo 1988; Crueger and Stevens 2015). The CRE is associated
59 with a more prominent single ITCZ (Crueger and Stevens 2015; Harrop and Hartmann 2016; Popp
60 and Silvers 2017). Both Harrop and Hartmann (2016) and Popp and Silvers (2017) investigated
61 the association between the Hadley circulation and CRE in a range of aquaplanet simulations with
62 and without the CRE. In all GCMs used, the CRE is associated with increased equatorial rainfall,
63 an equatorward contraction of the ITCZ, and a strengthening of the mean meridional circulation.
64 The authors emphasise different mechanisms by which the CRE promotes a single ITCZ. Harrop
65 and Hartmann (2016) propose that the CRE warms the upper tropical troposphere, which reduces
66 the convective available potential energy and restricts deep convection to the region of warmest
67 SSTs, whilst Popp and Silvers (2017) argue that the CRE strengthens the Hadley circulation and
68 moves the ITCZ equatorward, associated with increased moist static energy (MSE) advection by
69 the lower branches of the Hadley circulation. The strengthening of the mean circulation is asso-
70 ciated with the CRE meridional gradient, as the CRE is positive in the tropics and negative in the
71 extra-tropics ($\geq \pm 45^\circ$ latitude; Allan 2011). However, it should be noted that the CRE reduces
72 total tropical-mean ($\leq \pm 30^\circ$ latitude) precipitation due to reduced radiative cooling (Harrop and
73 Hartmann 2016).

74 Across a hierarchy of models it has been shown that the simulation of tropical precipitation
75 is sensitive to the representation of convection (Terray 1998; Frierson 2007; Wang et al. 2007;
76 Chikira 2010; Mobis and Stevens 2012; Oueslati and Bellon 2013; Bush et al. 2015; Nolan et al.

77 2016). For example, variations in lateral entrainment and detrainment rates, which alter the repre-
78 sentation of deep convection, affect the diurnal cycle of precipitation over the Maritime Continent
79 (Wang et al. 2007) and South Asian monsoon precipitation rates (Bush et al. 2015). Increasing
80 convective mixing strengthens deep convection in convergence zones, associated with an increased
81 moisture flux from subsidence regions (Terray 1998; Oueslati and Bellon 2013).

82 In full GCMs, complex surface characteristics and boundary conditions including land-sea con-
83 trasts, orography and SST gradients, make it challenging to understand the sensitivity of tropical
84 precipitation to the representation of convection (Oueslati and Bellon 2013; Bush et al. 2015).
85 Even in the absence of complex surface topography, aquaplanet studies have also shown that
86 characteristics of tropical precipitation, in particular the location and intensity of the ITCZ, are
87 sensitive to the sub-gridscale treatment of convection (Hess et al. 1993; Numaguti 1995; Chao
88 and Chen 2004; Liu et al. 2010; Mobis and Stevens 2012). Mobis and Stevens (2012) studied the
89 sensitivity of the ITCZ location to the choice of convective parameterisation scheme in an aqua-
90 planet configuration of the ECHAM GCM by comparing the Nordeng (1994) and Tiedtke (1989)
91 schemes, which vary in their formulations of entrainment, detrainment and cloud base mass flux
92 for deep convection. The Nordeng scheme, with a higher lateral entrainment rate, produced a
93 single ITCZ, whilst the Tiedtke scheme produced a double ITCZ. The authors associate the loca-
94 tion of maximum boundary layer MSE with the ITCZ location; they argue that mechanisms that
95 control the boundary layer MSE are important to the sensitivity of the ITCZ to the representation
96 of convection. The boundary layer MSE distribution is predominantly controlled by the surface
97 winds, which are influenced by convective heating, allowing variations in convective heating to
98 influence the ITCZ structure. The importance of the surface winds is further emphasised by simu-
99 lations with prescribed surface winds in the computation of the surface fluxes (Mobis and Stevens
100 2012). These simulations lead to the conclusion that there is a strong association between surface

101 turbulent fluxes and the ITCZ.

102 While the ITCZ has been shown to be sensitive to the CRE and the convective parameterisation
103 scheme, no study has separated these effects. This paper will analyse the sensitivity of the ITCZ to
104 convective mixing in aquaplanet simulations using the Met Office Unified Model (MetUM), and
105 the role of the CRE in this sensitivity.

106 *b. Atmospheric Energy framework*

107 Literature based on a hierarchy of models, as well as reanalysis data and observations, concludes
108 that the northward displacement of the ITCZ from the equator is anti-correlated with the north-
109 ward cross-equatorial atmospheric energy transport (Kang et al. 2008; Frierson and Hwang 2012;
110 Donohoe et al. 2013). Bischoff and Schneider (2014) developed a diagnostic framework to relate
111 the location of the ITCZ to this energy transport.

112 The zonal-mean atmospheric MSE budget is (Neelin and Held 1987):

$$[AEI] = \partial_t[\hat{h}_e] + \partial_y[\hat{v}h] \quad (1)$$

113 where AEI is the atmospheric energy input (AEI); vh is the meridional MSE flux, (v is meridional
114 wind; h is MSE); h_e is the moist enthalpy; $[\]$ denotes a zonal- and time-mean; $\hat{\ }^$ represents a mass
115 weighted vertical integral; ∂_y is the meridional derivative; and ∂_t is the time derivative. Local
116 Cartesian coordinates are printed with $y = a\phi$, (where a is Earth's radius and ϕ is latitude,) but
117 all calculations are performed in spherical coordinates. Bischoff and Schneider (2014) assume
118 a statistically steady state ($\partial_t[\hat{h}_e]=0$) and that $[\hat{v}h]$ in the tropics is dominated by the zonal-mean
119 circulation and therefore $[\hat{v}h]$ equals zero at the ITCZ. Through performing a first-order Taylor
120 expansion of the equatorial $[\hat{v}h]$, Bischoff and Schneider (2014) derive the dependence of the
121 ITCZ location on the equatorial MSE flux and equatorial AEI:

$$\delta \approx -\frac{1}{a} \frac{[\widehat{vh}]_0}{[AEI]_0} \quad (2)$$

122 with the AEI defined as:

$$[AEI] = [S] - [L] - [O] \quad (3)$$

123 where subscript 0 denotes the equatorial value, S is the net incoming shortwave radiation at the
 124 top of the atmosphere (TOA), L is the outgoing longwave radiation at the TOA, and O is the net
 125 downward flux at the surface. Bischoff and Schneider (2016) retain higher order terms in the
 126 Taylor expansion to derive a framework for negative $[AEI]_0$. A negative $[AEI]_0$ is associated with
 127 a double ITCZ as $[\widehat{vh}]$ no longer increases with latitude; energy is transported equatorward at low
 128 latitudes to achieve equilibrium. A double ITCZ is associated with two off-equatorial energy flux
 129 equators, where the total meridional energy flux equals zero. Bischoff and Schneider (2016) derive
 130 an expression for the locations of a double ITCZ:

$$\delta \approx \pm \frac{1}{a} \left\{ -\frac{6([AEI]_0)}{\partial_{yy}([AEI]_0)} \right\}^{\frac{1}{2}} + \frac{[\widehat{vh}]_0}{2a([AEI]_0)} \quad (4)$$

131 Note equation 4 is from a corrigendum for the original paper.

132 Bischoff and Schneider (2014) explore the relationship derived in (2) using an idealised slab-
 133 ocean GCM with a prescribed oceanic heat transport. They investigate the effects of the $[AEI]_0$ and
 134 the $[\widehat{vh}]_0$ through varying the imposed equatorial ocean heat flux and the atmospheric longwave
 135 absorption. Changes in both $[AEI]_0$ and $[\widehat{vh}]_0$ affect the latitude of the ITCZ; this theoretical rela-
 136 tionship is supported in observations and reanalyses (Adam et al. 2016). Bischoff and Schneider
 137 (2016) examine the double ITCZ framework (4) using a slab-ocean GCM and varying the tropical
 138 and extra-tropical components of the imposed ocean energy flux divergence. An increased tropical
 139 ocean energy flux divergence decreases the $[AEI]_0$. For double ITCZ scenarios and when $[\widehat{vh}]_0$ is
 140 negligible, decreasing the $[AEI]_0$ shifts the energy flux equator poleward. The diagnosed energy

141 flux equators from (2) and (4) are close to the simulated precipitation maxima, highlighting the
142 association between the AEI and ITCZ.

143 However, Bischoff and Schneider (2014)’s definition of the $[AEI]$ (3) is chosen as their simu-
144 lations prescribe O , which allows only the TOA energy budget ($S - L$) to vary. This constrains
145 the AEI response to model perturbations, as surface radiation and turbulent fluxes are constrained
146 at equilibrium, which could reduce the impact of surface-flux feedbacks on the ITCZ. We use
147 atmosphere-only simulations with prescribed SSTs, allowing variations in the components of O .
148 As our experiments do not have a closed surface energy balance and we are interested in cloudy-
149 sky radiation AEI components, we choose to write the AEI as:

$$[AEI] = [SW] + [LW] + [H] \quad (5)$$

150 where SW and LW represent the net atmospheric heating from shortwave and longwave radiation,
151 respectively, and H denotes the atmospheric heating from surface sensible and latent heat fluxes.
152 Both fixed SST and prescribed O frameworks misrepresent the real climate system by restricting
153 air-sea coupled feedbacks (discussed further in section 4). From an AEI perspective, Mobis and
154 Stevens (2012) severely constrain H in a subset of experiments by prescribing the surface winds
155 when computing the surface fluxes. This reduces the sensitivity of the ITCZ to the convective
156 parameterisation scheme.

157 Previous research on the response of the simulated ITCZ to variations in the sub-gridscale rep-
158 resentation of convection have not considered the role of the CRE or used an energy budget frame-
159 work like that proposed by Bischoff and Schneider (2014). We hypothesise that the sensitivity of
160 the ITCZ to these factors can be linked to variations in AEI and $[\widehat{vh}]$.

161 2. Methodology

162 We use variations of an N96 (1.25° latitude × 1.875° longitude) aquaplanet configuration of the
163 Met Office Unified Model (MetUM) Global Atmosphere 6.0 (GA6.0) configuration (Walters et al.
164 2017). The deep convective parameterisation scheme is an altered form of the mass flux scheme
165 in Gregory and Rowntree (1990), including a convective available potential energy closure based
166 on Fritsch and Chappell (1980) and a mixing detrainment rate dependent on the relative humidity
167 (Derbyshire et al. 2004). Unless noted, all simulations are run for three years with a “Qobs” SST
168 profile (Neale and Hoskins 2001), with the first sixty days discarded as spin-up.

169 *a. Simulations performed*

170 To explore the sensitivity of the simulated ITCZ to convective mixing, we perform five simu-
171 lations varying the lateral entrainment (ε) and detrainment (d_m) rates for deep-level convection
172 (Table 1). In GA6.0 these rates are:

$$\varepsilon = 4.5 f_{dp} \frac{p(z)\rho(z)g}{p_*^2} \quad (6)$$

$$d_m = 3.0(1 - RH)\varepsilon \quad (7)$$

174 Both ε and d_m are given as a fractional mixing rate per unit length (m^{-1}). In (6) and (7), p and
175 p_* are pressure and surface pressure (Pa); ρ is density ($kg\ m^{-3}$); g is gravitational acceleration
176 ($m\ s^{-2}$); f_{dp} is a constant with the default value of 1.13; RH is relative humidity. We control ε
177 and d_m by scaling f_{dp} to five values between 0.25 and $1.5 \times$ the default value: 0.28 (F0.28), 0.57
178 (F0.57), 0.85 (F0.85), 1.13 (F1.13) or 1.70 (F1.70).

179 To explore the influence of the CRE on the sensitivity of the ITCZ to convective mixing we
180 perform a companion set of experiments with cloud-radiation interactions removed: F0.28NC,

181 F0.57NC, F0.85NC, F1.13NC and F1.70NC (Table 1). Cloud-radiation interactions are removed
182 by setting the cloud liquid and cloud ice to zero in the radiation scheme.

183 Finally, a third set of simulations use a prescribed CRE (Table 2) to investigate the relative
184 importance of f_{dp} and the CRE to characteristics of the ITCZ. The four simulations have a pre-
185 scribed, diurnally varying CRE vertical profile computed from a single-year simulation with f_{dp}
186 equal to 0.57 or 1.13 (PC0.57 and PC1.13, respectively). The CRE is prescribed using cloudy-sky
187 upward and downward fluxes at each model level at every model timestep. The diurnally varying
188 CRE profile is computed as a hemispherically symmetric and zonally uniform composite of the
189 climatological diurnal cycle at each grid point, referenced to local solar time. Two of the four
190 simulations prescribe a CRE at a different f_{dp} constant from that in the simulation (F1.13PC0.57,
191 F0.57PC1.13), whilst the other two simulations use a CRE from the same f_{dp} value to assess the
192 sensitivity to prescribing cloud-radiation interactions (F1.13PC1.13, F0.57PC0.57).

193 3. Results

194 a. Sensitivity of the ITCZ to the convective mixing.

195 Figure 1a shows the sensitivity of the ITCZ to f_{dp} with a single ITCZ at higher values (F1.13,
196 F1.70). Reducing f_{dp} promotes a double ITCZ, with peak precipitation further away from the
197 equator (F0.28, F0.57). F0.85 has a marginal double ITCZ with no substantial difference between
198 equatorial and off-equatorial precipitation. Decreasing f_{dp} is associated with a weaker horizontal
199 gradient of the mass meridional streamfunction (Figure 2). F0.28 is the only simulation to
200 show a reversed Hadley circulation in the deep tropics (Figure 2e), associated with upper-level
201 zonal-mean equatorial subsidence, typical of a double ITCZ. F0.57 meanwhile has a typical
202 double ITCZ structure in precipitation but not in the mass meridional streamfunction (Figure 1a

203 and 2d), which we refer to as a “split ITCZ”: two off-equatorial precipitation maxima and two
204 ascending branches of the Hadley circulation, without any substantial zonal-mean subsidence
205 equatorward of the precipitation maxima.

206 Convective mixing reduces the difference in MSE between a convective plume, determined by
207 the boundary layer MSE, and the free-troposphere (Mobis and Stevens 2012), which reduces the
208 buoyancy of the convective plume. Assuming the sensitivity of the environmental saturated MSE
209 to f_{dp} is small, the depth of convection will depend on the boundary layer MSE and f_{dp} . De-
210 creasing f_{dp} will deepen convection for a constant boundary layer MSE, and reduce the minimum
211 boundary layer MSE at which deep convection occurs. Following weak-temperature gradient
212 arguments (e.g. Sobel et al. 2001) and assuming a small meridional gradient in free-tropospheric
213 tropical temperature, and hence a small gradient in the saturated MSE across the deep tropics,
214 the reduced minimum boundary layer MSE needed for deep convection strengthens convection
215 in off-equatorial tropical latitudes over cooler SSTs. Stronger off-equatorial deep convection
216 decreases equatorward low-level winds in the deep tropics, reducing equatorial boundary layer
217 MSE. Hence, decreasing f_{dp} is associated with a poleward ITCZ shift and promotes a double
218 ITCZ. Similar arguments can be made for higher f_{dp} promoting a single ITCZ.

219 The sensitivity of the ITCZ to f_{dp} is associated with AEI changes (Figure 1b), with a change
220 from a single (F1.13) to a double/split ITCZ (F0.28/F0.57) associated with a decrease in the $[AEI]_0$
221 (Figure 3d and e). Simulations with a single/double ITCZ in precipitation have a positive/negative
222 $[AEI]_0$ (Figure 1b), in agreement with Bischoff and Schneider (2014). Changes in cloudy-sky
223 radiation and latent heat flux are the dominant components of AEI changes (blue and orange
224 lines, respectively, in Figure 3). In F1.13 the total CRE peaks at approximately 60 Wm^{-2} at the
225 equator and reduces to zero around 15° latitude (blue line in Figure 3b). This equatorial warming
226 comes almost entirely from the longwave CRE, which dominates the total CRE equatorward

227 of 10° latitude (not shown). In the subtropics, 20° to 30° latitude, low clouds contribute to a
 228 negative CRE of $\approx 2 \text{ Wm}^{-2}$, as longwave cooling from boundary layer clouds is greater than
 229 the shortwave heating. Without the CRE contribution to the $[AEI]_0$ in F1.13, $[AEI]_0$ would be
 230 negative, suggesting that the CRE maintains the single ITCZ. Removing the CRE from the AEI
 231 in F1.13 would give an $[AEI]_0$ of -25.7 Wm^{-2} , assuming that no other AEI components change.
 232 Using Bischoff and Schneider (2016)'s framework, (4), with values for AEI once removing
 233 the CRE and assuming that $[\widehat{v\bar{h}}]_0 \simeq 0 \text{ Wm}^{-1}$, (associated with an hemispherically symmetric
 234 atmospheric circulation), predicts a double ITCZ at $\pm 5.6^\circ$ latitude.

235 The split ITCZ in F0.57 is associated with a substantially reduced equatorial CRE and an
 236 increased off-equatorial CRE (Figure 3d). We chose CRE profiles from one year of F0.57 and
 237 F1.13 for our prescribed CRE simulations (Table 2), as these two simulations show CRE profiles
 238 typical of a double and single ITCZ, respectively; these simulations are analysed in section 3d.
 239 As the Hadley circulation and ITCZ are associated with the AEI, and the CRE plays a substantial
 240 role in AEI changes when varying f_{dp} , we hypothesize that prescribing the CRE will reduce or
 241 remove the sensitivity of the AEI and ITCZ to f_{dp} .

242

243 *b. Sensitivity of the ITCZ to convective mixing with no cloud radiative effect*

244 To test our hypothesis above, we first analyse simulations with the CRE removed (Table
 245 1), similar to Harrop and Hartmann (2016). Figure 4a and Figure 5 show the zonal-mean
 246 precipitation and mass meridional streamfunction respectively in simulations with no CRE (Table
 247 1). Removing the CRE at $f_{dp} = 1.13$ (F1.13NC) leads to a switch from a single to a split ITCZ,
 248 and a $\approx 20\%$ weakening of the Hadley circulation (Figure 4a and 5b).

249 Similar to Harrop and Hartmann (2016), removing the CRE cools the tropical ($\leq 30^\circ$ latitude)

250 upper-troposphere, destabilizing the atmosphere and reducing the environmental saturated
251 MSE. For a fixed boundary layer MSE and convective mixing rate, removing the CRE deepens
252 convection as the buoyancy of a convective plume increases relative to the saturated MSE of the
253 environment. Hence, removing the CRE reduces the minimum boundary layer MSE for deep
254 convection, strengthening off-equatorial convection over cooler SSTs. Stronger off-equatorial
255 convection decreases equatorward low-level winds in the deep tropics, reducing equatorial
256 boundary layer MSE and promoting a double ITCZ. This mechanism is similar to that proposed
257 for the sensitivity of the ITCZ to f_{dp} (section 3a). However, when removing the CRE changes in
258 the environmental saturated MSE play the dominant role, whilst for the sensitivity of the ITCZ to
259 f_{dp} , changes in the convective parcel MSE dominate.

260 The weaker Hadley circulation and double ITCZ in precipitation in F1.13NC is consistent
261 with AEI changes. In F1.13NC removing CRE reduces the $[AEI]_0$ by $\approx 45 \text{ Wm}^{-2}$, leading to a
262 negative $[AEI]_0$, and increases the subtropical AEI by up to 15 Wm^{-2} (20 to 45° latitude) (Figure
263 6f). Across the deep tropics the AEI change is not equal to the CRE diagnosed from F1.13, due to
264 increased turbulent and clear-sky fluxes. These increased fluxes, associated with an equatorward
265 shift of the ITCZ, partially offset the reduction in $[AEI]_0$. Hence, the predicted location of the
266 double ITCZ in section 3a when removing the CRE overestimated the poleward shift of the
267 ITCZ. Removing the CRE reduces tropical-domain ($\leq 30^\circ$ latitude) AEI, which is associated
268 with increased AEI at higher latitudes to maintain equilibrium. Our simulations are consistent
269 with the suggested mechanisms proposed by Popp and Silvers (2017): the ITCZ is located at the
270 maximum boundary layer MSE, and a weaker meridional circulation is associated with a reduced
271 AEI gradient.

272 At all f_{dp} removing the CRE reduces the maximum precipitation rate, weakens the Hadley
273 circulation (comparing Figure 1a and 4a), and moves the latitude of peak precipitation poleward

274 (Figure 7a). The sensitivity of the ITCZ structure to removing the CRE depends on the convective
275 mixing rate: either a broader single ITCZ (F1.70NC), a poleward shift of a double/split ITCZ
276 (F0.28NC and F0.57NC), or a switch from a single to a split/double ITCZ (F0.85NC and
277 F1.13NC). Removing the CRE cools the upper troposphere and reduces the boundary layer MSE
278 required for deep convection. This increases the f_{dp} value at which the ITCZ transitions from
279 single to split/double.

280 Removing the CRE changes, but does not remove, the sensitivity of the ITCZ to f_{dp} . Quan-
281 tifying the apparent effect of the CRE on the sensitivity of the ITCZ to f_{dp} is difficult, as the
282 effect depends on both the range of f_{dp} considered and the metric used (Figure 7). When an
283 off-equatorial ITCZ is simulated in CRE-off simulations ($0.28 \leq f_{dp} \leq 1.13$), including the CRE
284 increases the sensitivity of the ITCZ location to f_{dp} by $\approx 30\%$ (comparing the slopes of the solid
285 regression lines in Figure 7a). However, because F1.70NC has a single ITCZ, including the CRE
286 cannot shift the ITCZ equatorward. Hence, when $0.28 \leq f_{dp} \leq 1.70$ the change in sensitivity
287 reduces to nearly zero (comparing the slopes of the dashed lines). The reduction in sensitivity
288 also depends on the chosen metric; for instance, the maximum precipitation rate has a negligible
289 sensitivity to f_{dp} in CRE-off simulations but a substantial sensitivity in CRE-on simulations
290 (Figure 7b), highlighting that the CRE has a positive feedback on convection as increasing f_{dp} is
291 associated with an increased CRE (Figure 8).

292 Increasing f_{dp} is associated with an increased tropical-domain CRE (Figure 8), which is
293 counter-intuitive as one might expect that increasing f_{dp} will lead to lower cloud tops and
294 hence a reduced CRE. However, the maximum cloud top height at the ITCZ is insensitive to
295 f_{dp} (not shown), but the minimum temperature where the cloud fraction goes to zero (cloud top
296 temperature) is sensitive to f_{dp} in both CRE-on and CRE-off simulations (Figure 8). The cloud
297 top temperature decreases as f_{dp} increases (Figure 8), associated with a cooler upper-troposphere.

298 Furthermore, the increase in SST at the ITCZ location, associated with equatorward contraction
299 of the ITCZ, also contributes to an increased CRE at higher f_{dp} .

300 Removing the CRE decreases the sensitivity of the AEI to f_{dp} (comparing Figure 1b and
301 Figure 4b). The reduced sensitivity of the AEI is associated with a reduced sensitivity of the
302 ITCZ. Latent heat flux variations account for most of the remaining AEI sensitivity to f_{dp}
303 (Figure 6). In simulations with a double ITCZ (F0.28NC, F0.57NC and F0.85NC), changes in
304 the latent heat flux and AEI have a bi-modal structure, indicating reduced latent heat flux at the
305 location of maximum precipitation in F1.13NC (Figure 6c-e). Changes in the latent heat flux
306 are predominantly controlled by alterations in near-surface wind speed rather than changes in
307 near-surface specific humidity (not shown).

308 Simulations so far agree with the association in Bischoff and Schneider (2016) between a neg-
309 ative $[AEI]_0$ and a double ITCZ. However, the negative $[AEI]_0$ in F0.57, F0.85NC and F1.13NC
310 requires an equatorward transport of energy at low latitudes, but the mean mass meridional
311 streamfunction suggests a poleward transport of energy (Figure 2b, 5c, 5d). In the following
312 subsection we discuss mechanisms for an equatorward energy transport.

313

314 *c. Mechanisms responsible for an equatorward energy transport*

315 To better understand the response of the mean circulation, associated with ITCZ changes, to
316 varying f_{dp} and removing the CRE, we partition the divergence of the MSE flux ($\partial_y[\widehat{v\hat{h}}]$) into
317 two components: the mean circulation ($\partial_y([\widehat{v}][\widehat{h}])$) and the eddy contribution ($\partial_y[\widehat{v\hat{h}}] - \partial_y([\widehat{v}][\widehat{h}])$).
318 In these simulations it has not been possible to close the atmospheric energy budget (1) due to
319 local energy conservation issues (discussed further in section 4), however the sign of the $[AEI]_0$
320 is consistent with the sign of the $\partial_y[\widehat{v\hat{h}}]$ in simulations so far. In all simulations the eddy con-

321 tribution to the meridional MSE flux is substantial across the tropics highlighting that the mean
 322 atmospheric circulation is not solely responsible for transporting energy. Furthermore, one should
 323 not necessarily assume a correspondence between the required MSE transport and the transport
 324 by the mean meridional circulation. In simulations with a single/double ITCZ, both the mean cir-
 325 culation and eddies transport energy poleward/equatorward at low latitudes. In F0.57, which has
 326 a negative $[AET]_0$ and a split ITCZ, equatorward transport of energy at low latitudes is achieved
 327 solely by eddies. When f_{dp} equals 0.85 and 1.13, a change in the sign of the energy transport by
 328 the mean circulation ($\partial_y([\hat{v}][\hat{h}])$) occurs at low latitudes when removing the CRE, however there is
 329 still equatorial ascent across most of the troposphere (Figure 5b, c). To understand the sensitivity
 330 of the mean circulation to removing the CRE at these convective mixing rates, we partition the
 331 change in the MSE flux ($[\hat{v}][\hat{h}]$) into mean circulation changes and MSE variations.

332 First, the meridional mass flux, denoted by V , in F1.13NC (V_e) is partitioned into two compo-
 333 nents:

$$\begin{aligned}
 V_e &= V_c(1 + \alpha) + V_r \\
 \text{where } \alpha &= \frac{V_e \cdot V_c}{V_c \cdot V_c} - 1
 \end{aligned}
 \tag{8}$$

334 Subscripts c and e represent the zonal-, time-mean value of the control and experiment simulation
 335 (in this case F1.13 and F1.13NC respectively). α is a globally uniform scaling term calculated
 336 using the dot product of the meridional mass fluxes in the tropics (30°N to 30°S). We account for
 337 variations in density in V . $V_c(1 + \alpha)$ represents a change in strength of the control circulation; V_r
 338 represents a change in circulation structure. Next, the MSE, $(c_p T + gz + Lq)$, in the experiment
 339 simulation (h_e) is written as:

$$h_e = h_c + h_p
 \tag{9}$$

340 where subscript p represents the zonal-, time-mean difference between the two simulations. The
 341 change in the MSE flux between the experiment and control simulation can therefore be written
 342 as:

$$V_e h_e - V_c h_c = \alpha V_c h_c + V_r h_c + V_c h_p + (\alpha V_c + V_r) h_p \quad (10)$$

343 Each term in (10) represents a mechanism by which vh can vary: $\alpha V_c h_c$ represents circulation
 344 intensity changes; $V_r h_c$ represents changes in circulation structure; $V_c h_p$ represents MSE profile
 345 changes; and $(\alpha V_c + V_r) h_p$, represents MSE profile changes correlated with changes in circulation
 346 structure and strength.

347 Three out of the four mechanisms are important in reducing the poleward MSE transport by
 348 the Hadley circulation in F0.85NC and F1.13NC (Figure 9): a reduction in Hadley circulation
 349 strength (Figure 9e); a shallower mean circulation (Figure 9f); and a reduced MSE export at
 350 the top of the Hadley circulation due to lower MSE associated with upper-tropospheric cooling
 351 (Figure 9g). MSE profile changes correlated with changes in circulation strength and intensity
 352 $[(\alpha V_c + V_r) h_p]$ are small compared to the other three mechanisms (Figure 9h). As changes
 353 in circulation strength ($\alpha V_c h_c$) cannot change the direction of energy transport, the reduced
 354 upper-tropospheric MSE ($V_c h_p$) and shallower Hadley circulation ($V_r h_c$) must be responsible
 355 for the change in energy transport direction by the mean circulation. At the equator, circulation
 356 strength changes ($\alpha V_c h_c$) contribute $\approx 16\%$ of the reduced $\partial_y([\hat{v}][\hat{h}])$; reduced MSE export by the
 357 upper branch of the mean circulation ($V_c h_p$) and a shallower Hadley circulation ($V_r h_c$) contribute
 358 $\approx 34\%$ and 50% respectively (not shown). Therefore, at certain convective mixing rates, in our
 359 case when $f_{dp} = 0.85$ and 1.13 , removing the CRE is not associated with a substantial double
 360 ITCZ in the mass meridional streamfunction, even though MSE is transported equatorward at

low latitudes and the $[AEI]_0$ is negative. Similar behaviour has also been concluded by Popp and Silvers (2017) who found that in certain simulations the zero mass meridional streamfunction remained at the equator even when the $[AEI]_0$ was negative.

Removing the CRE and varying f_{dp} are associated with substantial AEI changes which require MSE transport variations. In the two sets of simulations discussed so far, we identified three mechanisms to transport MSE equatorward at low latitudes; which mechanisms dominates depends on the CRE and f_{dp} . First, in F0.28, F0.28NC and F0.57NC, subsidence across the equatorial region is associated with an equatorward MSE flux at low latitudes (Figure 2e and Figure 5d, e). Secondly, eddy energy transport plays a role in the equatorward MSE flux in F0.28, F0.57, F0.28NC, F0.57NC, F0.85NC. Thirdly, in F0.85NC and F1.13NC a shallower Hadley circulation and reduced upper-tropospheric MSE reduces the MSE exported in the upper branches of the mean circulation, resulting in a net equatorward MSE transport. All other simulations (F0.85, F1.13, F1.70 and F1.70NC) have a single ITCZ associated with a positive $[AEI]_0$ and poleward MSE transport at low latitudes.

d. Sensitivity of the ITCZ to convective mixing with a prescribed cloud radiative effect.

To further understand the role of the CRE on the sensitivity of the ITCZ to convective mixing, we perform prescribed-CRE simulations and vary f_{dp} (Table 2). The prescribed CRE is diagnosed from single-year simulations with f_{dp} equal to 1.13 or 0.57 (section 2). The effect of prescribing the diurnal cycle of the CRE in a simulation with the same f_{dp} is minimal; for example, the ITCZ is similar in F1.13PC1.13 and F1.13 (Figure 1 and 10). Hence, we only discuss the mean circulation in F1.13PC0.57 and F0.57PC1.13 (Figure 11a and c).

Similar to CRE-off simulations, the sensitivity of the ITCZ to f_{dp} reduces in prescribed CRE

384 simulations (Figure 10a) compared to CRE-off simulations (Figure 1a), associated with a reduced
385 sensitivity of the AEI to f_{dp} (Figure 10b, 12a and c). The prescribed CRE heating acts as a fixed
386 MSE source, which requires an increase in MSE export and hence increased convective activity.
387 In PC1.13 simulations the CRE maximises at the equator, which is associated with increased
388 equatorial convective activity and a single ITCZ. In PC0.57 simulations on the other hand, the
389 CRE peaks off the equator and promotes a double ITCZ. The root mean squared difference of
390 tropical precipitation and the mass meridional streamfunction illustrates that prescribing the
391 CRE reduces the sensitivity of the ITCZ and Hadley circulation to f_{dp} by $\approx 50\%$ (Table 3).
392 Whilst the CRE plays a role in the sensitivity of the ITCZ to convective mixing (for example,
393 comparing F1.13PC1.13 and F1.13PC0.57 in Figure 10a), the ITCZ and Hadley circulation are
394 still sensitive to f_{dp} . For example, reducing f_{dp} (F0.57PC1.13) leads to a weakening in the upper
395 branch of the mean circulation whilst changing the prescribed CRE (F1.13PC0.57) intensifies
396 the upper branch of the Hadley circulation as the higher f_{dp} value is associated with a cooler
397 upper-troposphere, hence, an intensified upper branch of the mean circulation is required for
398 similar MSE transport (comparing F1.13 in Figure 2b to F0.57PC1.13 and F1.13PC0.57 in Figure
399 11c and a, respectively). The response of convection to changes in convective mixing is partially
400 offset by the effect of prescribing the location of the CRE.

401 As in CRE-off simulations, AEI changes in prescribed CRE simulations when varying f_{dp}
402 are predominantly driven by latent heat flux variations. For example, between F1.13PC1.13 and
403 F0.57PC1.13, the equatorial latent heat flux reduces whilst the off-equatorial latent heat flux
404 increases (Figure 12a). These changes are partially offset by changes in the clear-sky radiation,
405 associated with a decrease in the TOA outgoing longwave radiation, due to an increase in
406 atmospheric water vapour content. As changes in the ITCZ are associated with AEI changes,
407 we conclude that the remaining sensitivity of the ITCZ to f_{dp} in prescribed CRE simulations is

408 associated with latent heat flux variations. In simulations where the prescribed CRE is varied
409 but the same f_{dp} value is used, AEI changes are mostly associated with cloudy-sky radiation
410 (Figure 12b, d). However, latent heat flux variations are of the same order of magnitude as when
411 varying f_{dp} . Using the same technique described in section 3c, we conclude that a shallower,
412 weaker Hadley circulation is primarily responsible for changes in the MSE transport by the mean
413 circulation when reducing f_{dp} or changing the prescribed CRE from PC1.13 to PC0.57 (not
414 shown).

415 F1.13PC0.57 and F0.57PC1.13 have similar, split ITCZs (Figure 10a), yet very different
416 AEI profiles (Figure 10b, Figure 11b and d). F0.57PC1.13 highlights that a double ITCZ in
417 precipitation does not require a negative $[AEI]_0$ or an equatorward MSE transport (green and black
418 line respectively in Figure 11d), illustrating that a double ITCZ in precipitation is not necessarily
419 associated with an equatorward MSE flux at low latitudes. Instead a negative $[AEI]_0$ is a sufficient
420 but not a necessary condition for a double ITCZ in precipitation. Due to local energy conservation
421 issues, which are discussed further in section 4, it is challenging to understand F1.13PC0.57,
422 which shows a negative $[AEI]_0$ and a positive equatorial $\partial_y[\widehat{v}h]$ (Figure 11b), (contradicting (1) as
423 steady-state has been reached).

424

425 **4. Discussion**

426 We have analysed aquaplanet simulations with variations to convective mixing to show an
427 association between resultant variations in the AEI and characteristics of the ITCZ. Using the AEI
428 framework we have shown the importance of the CRE in the sensitivity of the ITCZ to convective
429 mixing. In a single ITCZ scenario (F0.85, F1.13 and, F1.70), the CRE is critical in maintaining a
430 positive $[AEI]_0$. For example, the $[AEI]_0$ would be negative without the CRE in F1.13 and F1.70,

431 associated with a double ITCZ. Changes in cloudy-sky radiation are the dominant cause of AEI
432 changes when varying the convective mixing rate, leading to our hypothesis that prescribing the
433 CRE would remove or reduce the sensitivity of the ITCZ to convective mixing. The fact that the
434 sensitivity of the ITCZ to f_{dp} remains in CRE-off and prescribed CRE simulations highlights the
435 importance of other AEI components, in particular the latent heat flux. All simulations, with the
436 exception of F0.57PC1.13, are consistent with Bischoff and Schneider (2016): a positive $[AEI]_0$
437 is associated with a single ITCZ and a negative $[AEI]_0$ with a double ITCZ.

438 CRE-off simulations illustrate that the CRE plays a substantial role in the structure and intensity
439 of the ITCZ. Similar to Harrop and Hartmann (2016), we observe that removing the CRE cools the
440 tropical upper-troposphere, reducing atmospheric stability and resulting in deep convection over
441 cooler SSTs. Stronger convection at higher latitudes reduces equatorial moisture convergence and
442 is associated with a double ITCZ. Removing the CRE also weakens the Hadley circulation which
443 is associated with a reduced AEI gradient between the tropics and sub-tropics, in agreement
444 with Popp and Silvers (2017). The sensitivity of the ITCZ to f_{dp} reduces when removing the
445 CRE, agreeing with our hypothesis that prescribing the CRE would either remove or reduce
446 the sensitivity of the ITCZ to convective mixing. Quantifying the reduction in sensitivity of
447 the ITCZ to f_{dp} when removing the CRE remains a challenge due to strong dependence on the
448 chosen metric and range of f_{dp} . It should also be noted that when removing the CRE other AEI
449 components change, such that the AEI change is not equal to the total CRE that is removed.

450 In prescribed CRE simulations, ITCZ characteristics are sensitive to both the prescribed CRE
451 and f_{dp} , however the sensitivity of the ITCZ to f_{dp} reduces by $\approx 50\%$ (Table 3). In prescribed
452 CRE simulations the response of convection to changes in convective mixing is offset by the effect
453 of prescribing the location of the CRE. Heating associated with the prescribed CRE is a MSE
454 source, therefore to increase the MSE exported, convective activity increases. The reduction in

455 sensitivity compliments work by Voigt et al. (2014), who found that prescribing the CRE reduced
456 the sensitivity of the ITCZ to hemispheric albedo perturbations to a similar degree. Thus, the
457 role of the CRE in the sensitivity of the ITCZ to both variations in the convection scheme and
458 boundary forcing appear similar, based on these two studies.

459 In both CRE-off and prescribed CRE simulations, latent heat flux alterations, associated with
460 circulation changes, are the predominant cause of AEI changes when varying f_{dp} . Circulation
461 changes when varying f_{dp} in CRE-off simulations are not associated with clear-sky flux variations,
462 consistent with Harrop and Hartmann (2016), which concluded that changes in the clear-sky
463 radiative cooling do not change the modelled circulation. Mobis and Stevens (2012) highlighted
464 the importance of surface fluxes in reducing the sensitivity of the ITCZ to the convective
465 parameterisation scheme when prescribing the wind speeds in the computation of surface fluxes.
466 Numaguti (1993) and Liu et al. (2010) also concluded that variations in surface evaporation are
467 associated with the ITCZ structure. We highlight that the sensitivity of the ITCZ to convective
468 mixing is predominantly associated with the surface fluxes in the absence of cloud feedbacks.

469 As noted earlier in sections 3c and 3d, the balance between the diagnosed AEI and diagnosed
470 $\partial_y[\widehat{vh}]$, equation (1), does not hold locally in MetUM. The mean of the maximum absolute
471 diagnosed imbalance across the tropics amongst simulations is 13.4 Wm^{-2} . More importantly,
472 the diagnosed equatorial energy imbalance ranges from 6.94 Wm^{-2} in F0.28NC to -20.63 Wm^{-2}
473 in F1.70 with a mean absolute error of 9.89 Wm^{-2} . For all of our simulations apart from
474 F1.13PC0.57, the sign of the equatorial $d_y[vh]$ and $[AEI]_0$ are the same, and therefore using $[AEI]_0$
475 as a proxy for the direction of energy transport at low latitudes is still valid. In F1.13PC0.57 the
476 difference between the diagnosed $d_y[vh]$ and $[AEI]$ is -16.9 Wm^{-2} ; the equatorial $d_y[vh]$ is positive
477 and $[AEI]_0$ is slightly negative (Figure 11b). Whilst the local energy imbalance is a concern for
478 F1.13PC0.57, we argue that in all other simulations the local energy imbalance does not affect

479 our conclusions. There are a number of possible reasons for the localised imbalance of the AEI
480 budget including: non-conservation associated with the semi-Lagrangian advection scheme in
481 MetUM; the use of dry and moist density in different components of the MetUM dynamics and
482 physics; errors in our diagnosis of the MSE budget, for example, not considering density changes
483 within a timestep; or, using an Eulerian approach for diagnosing the energy transport which is
484 inconsistent with the semi-Lagrangian advection scheme. It is worth noting that other studies
485 using the AEI framework have not shown that the MSE energy budget is locally closed, and this
486 problem may not be unique to our study. Nevertheless, the local energy imbalance has challenged
487 our interpretation of some simulations, and highlights that future modelling studies using an
488 atmospheric MSE budget should be cautious.

489 Variations in the CRE when varying f_{dp} can lead to a negative $[AEI]_0$ associated with a net
490 equatorward MSE energy transport at low latitudes. Whilst the predominant response to a negative
491 $[AEI]_0$ is a double ITCZ associated with equatorward energy transport at low latitudes by the
492 mean circulation (F0.28, F0.28NC and F0.57NC), F0.57, F0.85NC and F1.13NC have shown
493 that a net equatorward MSE transport can occur at low latitudes even with a poleward energy
494 transport by the mean flow at the tropopause. Two mechanisms can lead to this. Firstly, the MSE
495 flux due to eddies contributes a substantial proportion to the total MSE flux (as seen in Figure
496 11 12b and d), and this can support equatorward MSE transport. In F0.57, the MSE flux due to
497 eddies is responsible for a net equatorward energy transport in the deep tropics. This invalidates
498 the assumption that the energy flux equator is associated with zero MSE transport by the mean
499 circulation, as in Bischoff and Schneider (2016). This is also supported by the equatorward
500 displacement of the energy flux equator (from 2 and 4) relative to maximum precipitation in all
501 simulations except for F0.85NC and F1.70NC (Table 4). The second mechanism (F0.85NC and
502 F1.13NC) is a change in the MSE transport direction due to a shallower Hadley circulation and a

503 lower MSE in the upper-troposphere (section 3c). These changes reduce the MSE export in the
504 upper branch of the Hadley circulation, resulting in an equatorward MSE transport by the mean
505 circulation at low latitudes.

506 In our aquaplanet configuration SSTs are fixed which implies an arbitrary but varying oceanic
507 heat transport to maintain SSTs given a net surface heat flux imbalance. Thus, our aquaplanet
508 experiments may be viewed as energetically inconsistent. In Bischoff and Schneider (2014)
509 and Voigt et al. (2016) ocean heat transport, and hence the net downward flux at the surface, is
510 fixed, constraining the response of AEI components and potentially reducing the sensitivity of
511 the ITCZ to model perturbations. In reality the ocean circulation, and thus ocean heat transport,
512 is sensitive to changes in the surface wind stress. Therefore, both the SST and ocean heat
513 transport could change in response to tropical circulation changes from variations to f_{dp} or the
514 prescribed CRE. Recent work has shown that the ocean circulation plays an important role in
515 the meridional transport of energy (Green and Marshall 2017), and that sensitivities of the ITCZ
516 found in atmosphere-only simulations do not necessarily hold in a fully coupled model. For
517 example, coupling reduces the sensitivity of the ITCZ to an interhemispheric albedo forcing (e.g.
518 comparing Kay et al. (2016) and Hawcroft et al. (2017) to Voigt et al. (2014)). The radiative
519 effect of clouds on the surface and Ekman heat transport associated with a single ITCZ would be
520 expected to reduce the equatorial SST gradient, which would promote a double ITCZ (Numaguti
521 1995; Mobis and Stevens 2012) and may reduce the sensitivity of the ITCZ to convective mixing.
522 Coupled simulations with an interactive ocean are required to further investigate the sensitivity of
523 the ITCZ to the CRE and convective mixing.

524

525 **5. Conclusions**

526 The double ITCZ bias is a leading systematic error across a hierarchy of models (Li and Xie
527 2014; Oueslati and Bellon 2015). Inter-model variability in the ITCZ structure persists even
528 in a highly-idealised framework such as an aquaplanet with prescribed SSTs (Blackburn et al.
529 2013). This study confirms and extends previous research that variations in the convective
530 parameterisation scheme and convective mixing can alter the ITCZ (Figure 1a; Hess et al. 1993;
531 Numaguti 1995; Chao and Chen 2004; Liu et al. 2010; Mobis and Stevens 2012). Higher
532 convective mixing rates are associated with a single ITCZ whilst lower rates are associated with a
533 double ITCZ. As the convective mixing rate reduces, convection at higher latitudes strengthens,
534 decreasing equatorward low-level winds at low latitudes, promoting a double ITCZ structure.
535 The sensitivity of the ITCZ to convective mixing is associated with AEI changes, predominantly
536 caused by CRE variations. For example, the CRE plays an important role in maintaining a
537 positive equatorial AEI, and is therefore associated with a single ITCZ structure (consistent with
538 Harrop and Hartmann (2016) and Bischoff and Schneider (2016)'s framework). When removing
539 the CRE, the response of the ITCZ depends on the convective mixing rate. At low convective
540 mixing rates, where a double ITCZ is simulated with the CRE, precipitation bands shift poleward.
541 At high convective mixing rates the ITCZ broadens, whilst at certain convective mixing rates the
542 ITCZ structure changes from single to double. Quantifying whether the sensitivity of the ITCZ
543 to convective mixing reduces when removing the CRE is challenging, as the sensitivity depends
544 on the range of convective mixing rates and the chosen metric. Prescribing the CRE reduces
545 the sensitivity of the ITCZ to convective mixing by $\approx 50\%$. When removing or prescribing the
546 CRE other AEI components, in particular the latent heat flux, play a role in the sensitivity of
547 the ITCZ to convective mixing. Hence, simulations where the ocean heat transport is fixed,

548 thereby constraining surface fluxes, may underestimate the sensitivity of the ITCZ to changes
549 in model formulation. We have also shown two mechanisms responsible for a net equatorward
550 MSE transport even with no equatorial subsidence: MSE transport by eddies; and a reduced MSE
551 export in the upper branch of the mean circulation due to a shallower Hadley circulation. These
552 mechanisms highlight that caution should be taken when associating changes in the AEI to the
553 ITCZ structure.

554

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563

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705 TABLE 1. Simulations varying f_{dp} with cloud-radiation interactions on (CRE-on) and off (CRE-off). F1.13 is
 706 the default integration for GA6.0.

f_{dp}	CRE-on	CRE-off
0.28	F0.28	F0.28NC
0.57	F0.57	F0.57NC
0.85	F0.85	F0.85NC
1.13	F1.13	F1.13NC
1.70	F1.70	F1.70NC

707 TABLE 2. Simulations with a prescribed climatology of the CRE diurnal cycle. PC1.13 and PC0.57 represent
 708 the prescribed CRE diurnal cycle from a one-year simulation where f_{dp} equals 1.13 or 0.57 (respectively).

f_{dp}	PC1.13	PC0.57
1.13	F1.13PC1.13	F1.13PC0.57
0.57	F0.57PC1.13	F0.57PC0.57

709 TABLE 3. Root mean squared difference for tropical precipitation and mass meridional streamfunction be-
 710 tween two simulations. Tropical domain defined as 30°N to 30°S. Percentage value is the percentage reduction
 711 compared to F0.57 and F1.13.

Simulations	Precipitation (mm day ⁻¹)	Mass Meridional Streamfunction ($\times 10^{10}$ kg s ⁻¹)
F0.57 & F1.13	2.84	1.78
F0.57PC1.13 & F1.13PC1.13	1.18 (58%)	0.67 (62%)
F0.57PC0.57 & F1.13PC0.57	1.65 (42%)	0.96 (46%)

712 TABLE 4. AEI_0 , location of ITCZ and approximate energy flux equator (δ) using equation 2 or 4 in each
713 simulation. A single/double ITCZ is assumed when AEI_0 is positive/negative, respectively. Not applicable
714 (N/A) occurs when AEI_0 and $\partial_{yy}([AEI])_0$ are both negative and therefore the square root of $-\frac{6([AEI])_0}{\partial_{yy}([AEI])_0}$ has an
715 imaginary component.

Simulation	AEI_0 (W m^{-2})	ITCZ location ($^\circ$)	Energy Flux Equator (δ) location ($^\circ$)
F0.28	-18.1	8.13/8.13	6.85/-7.06
F0.57	-5.9	4.38/-4.38	0.84/-2.87
F0.85	33.4	1.88	-0.41
F1.13	36.7	0.63	0.22
F1.70	33.7	0.63	0.30
F0.28NC	-4.9	9.38/-9.38	N/A
F0.57NC	-12.2	8.13/-8.13	N/A
F0.85NC	-18.3	6.88/-5.63	6.48/-6.80
F1.13NC	-5.9	4.38/-4.38	3.21/-3.58
F1.70NC	2.0	1.88	2.73
F1.13PC1.13	33.6	0.63	0.16
F1.13PC0.57	-1.7	3.13/-3.13	0.19/-1.75
F0.57PC1.13	20.6	3.13	-0.12
F0.57PC0.57	-14.2	4.38/-4.38	2.70/-2.64

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746 line is printed in the legend. First value where $0.28 \leq f_{dp} \leq 1.13$ and second value where
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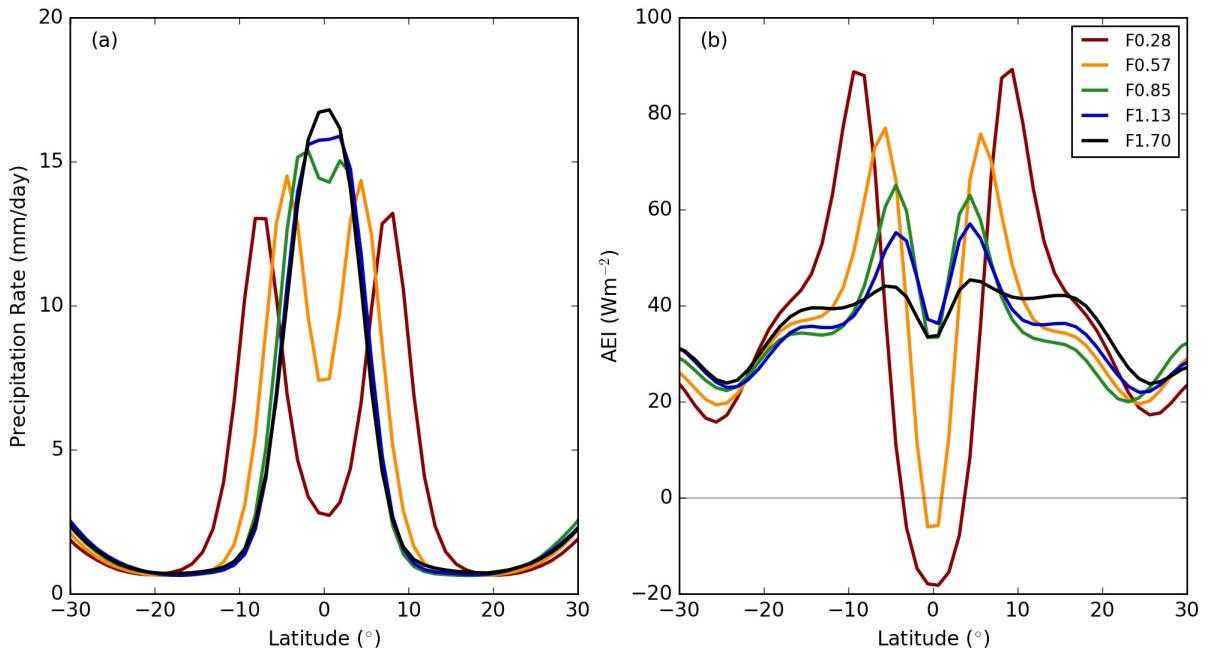
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756 changes in circulation structure and strength ($\alpha V_c + V_r$) h_p . Analysis explained in Section 3c. 48

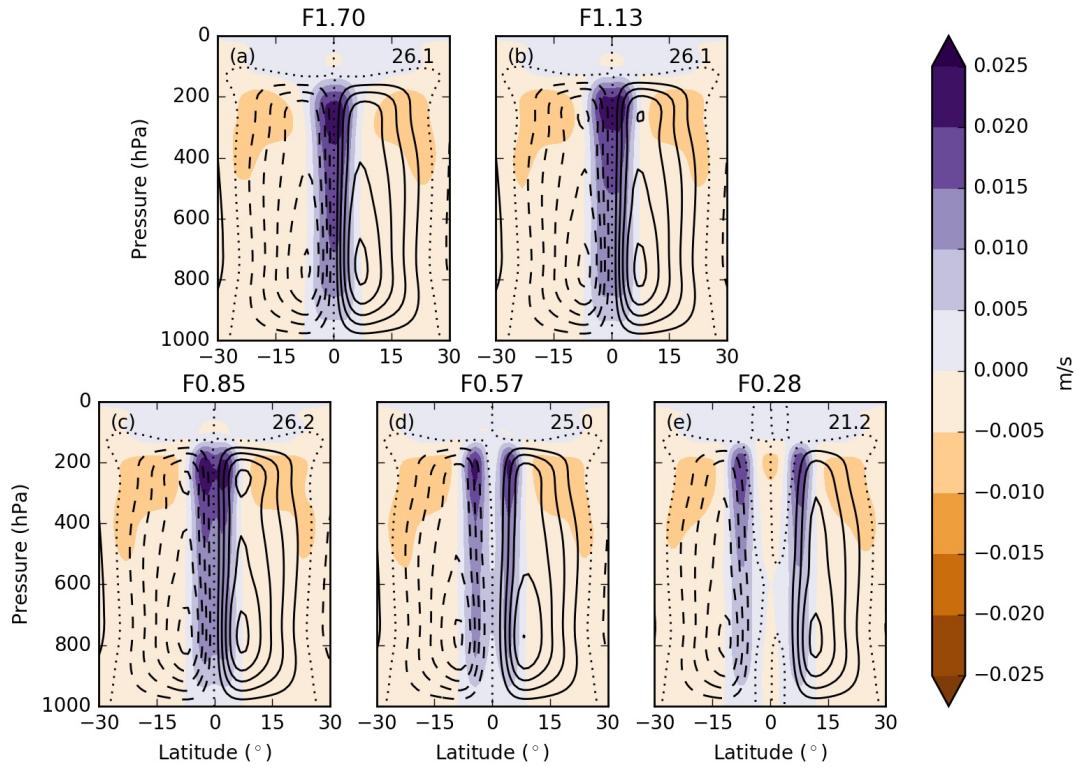
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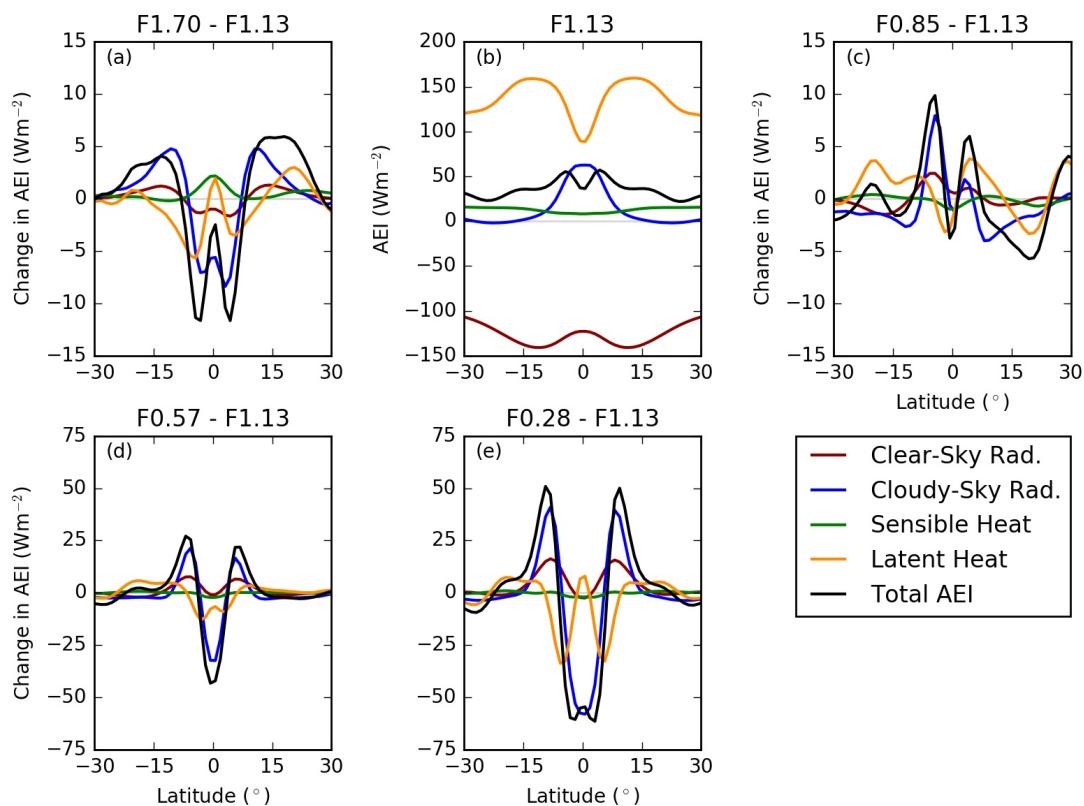
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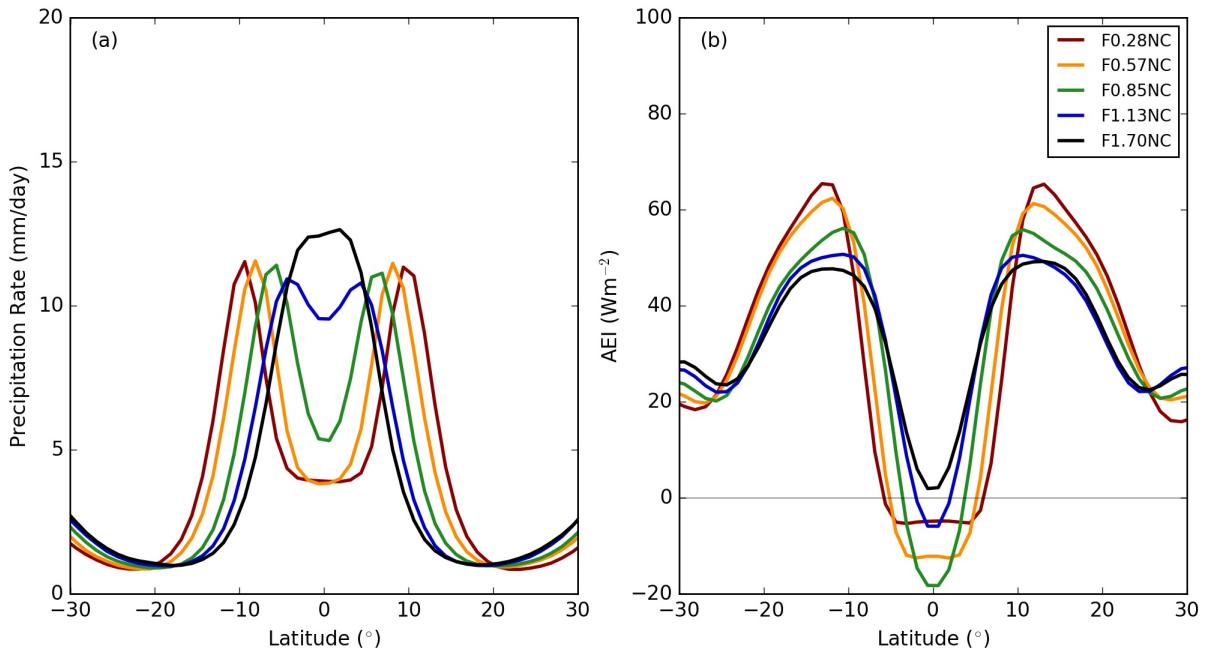
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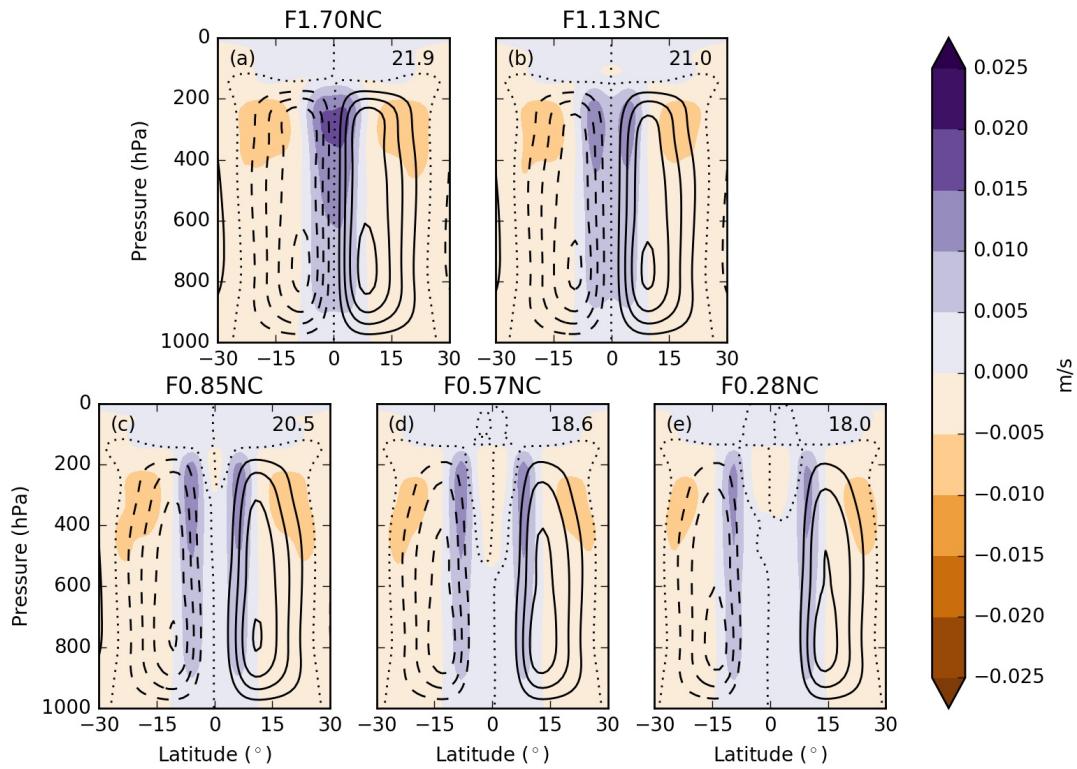
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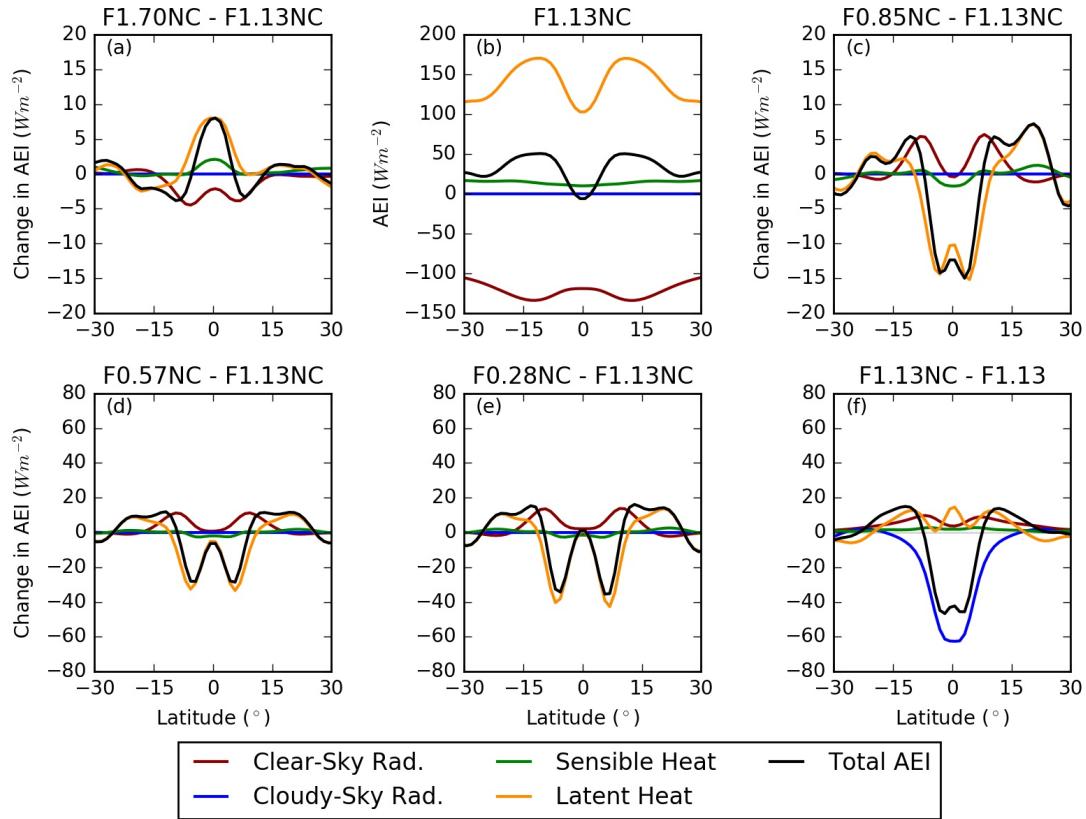
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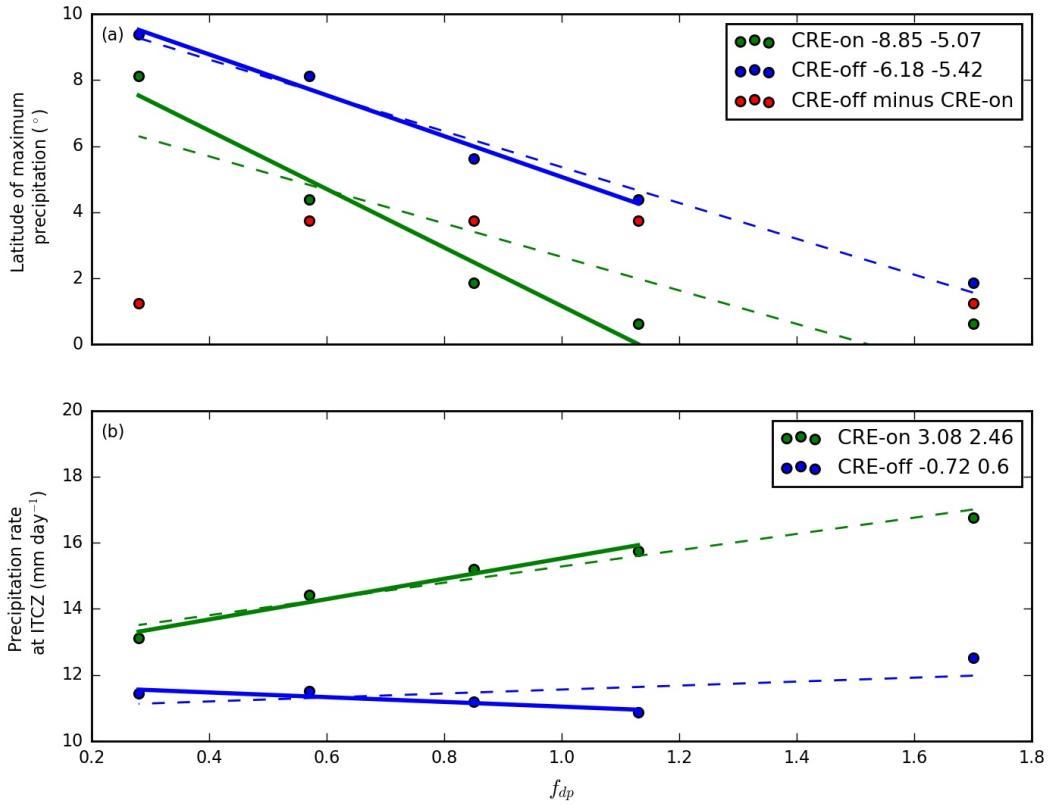
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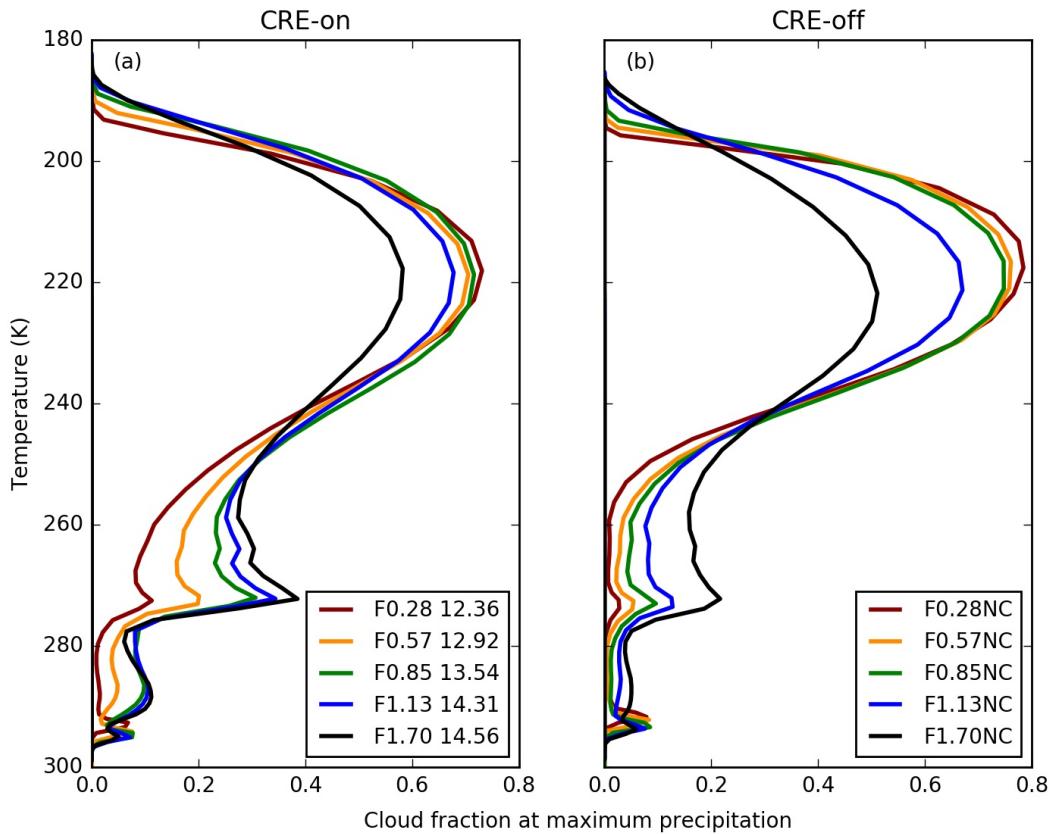
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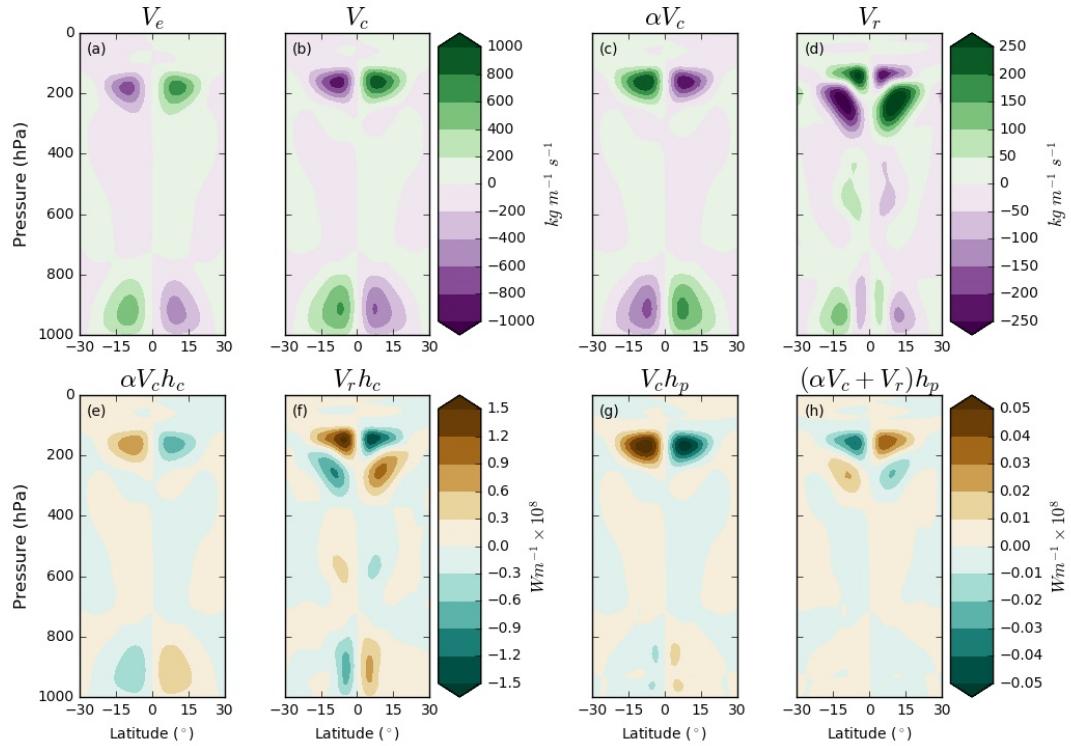
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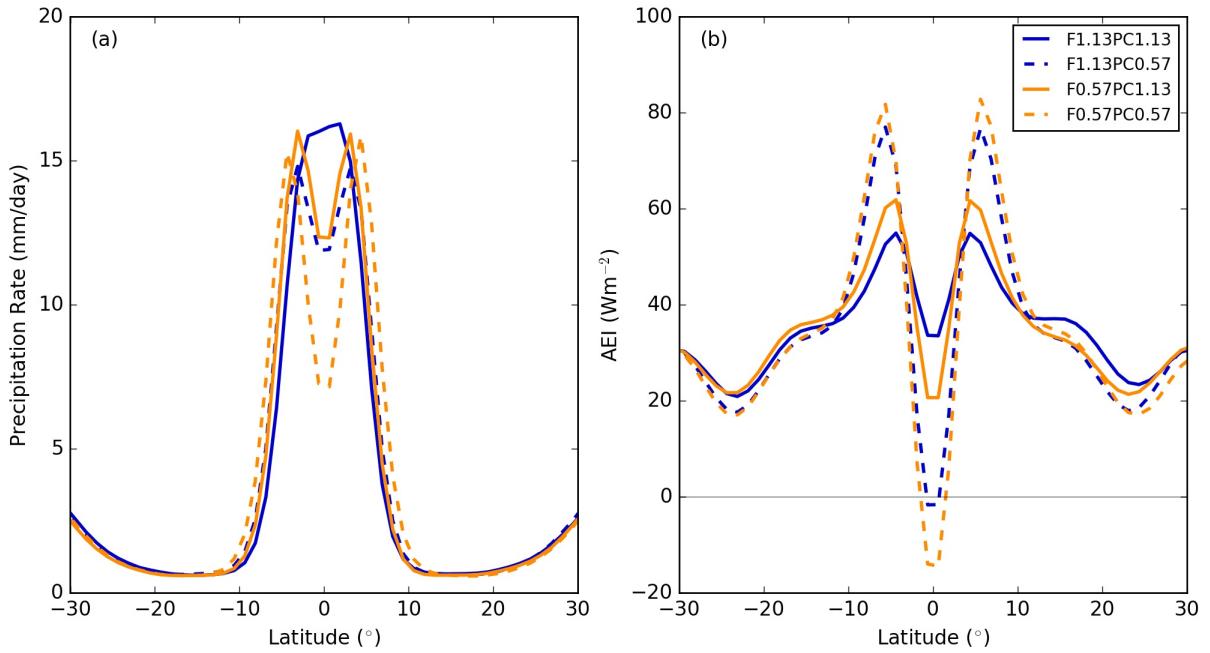
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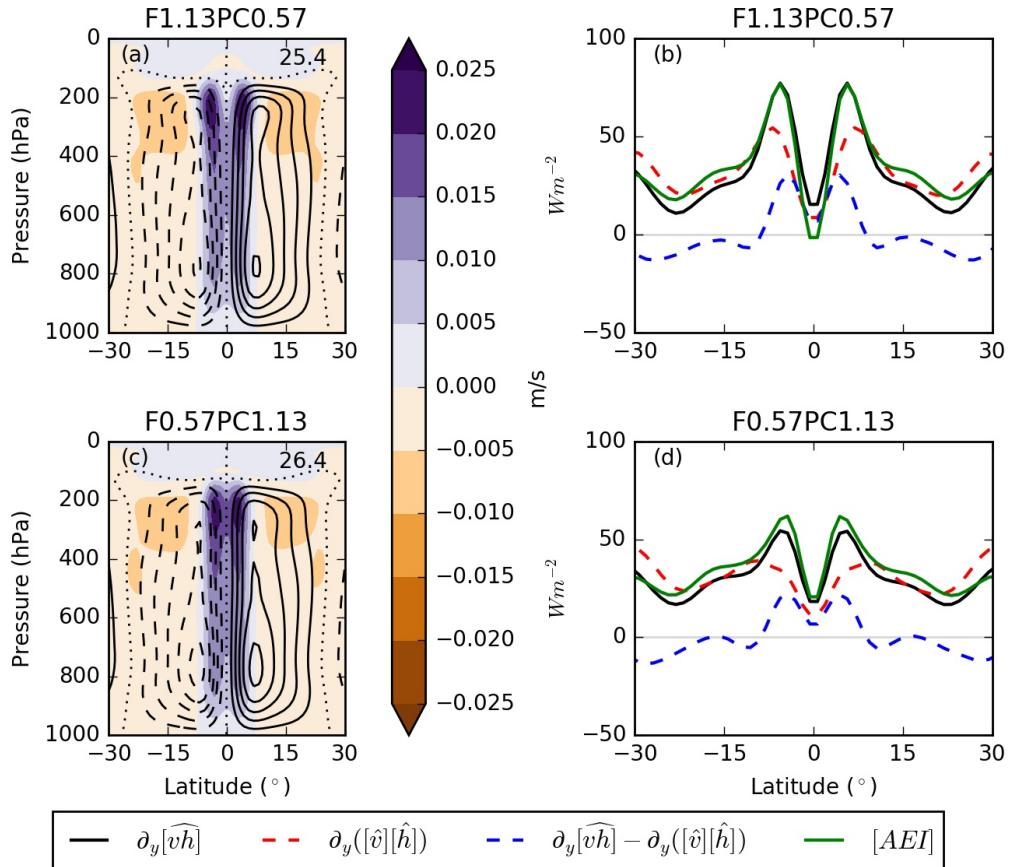
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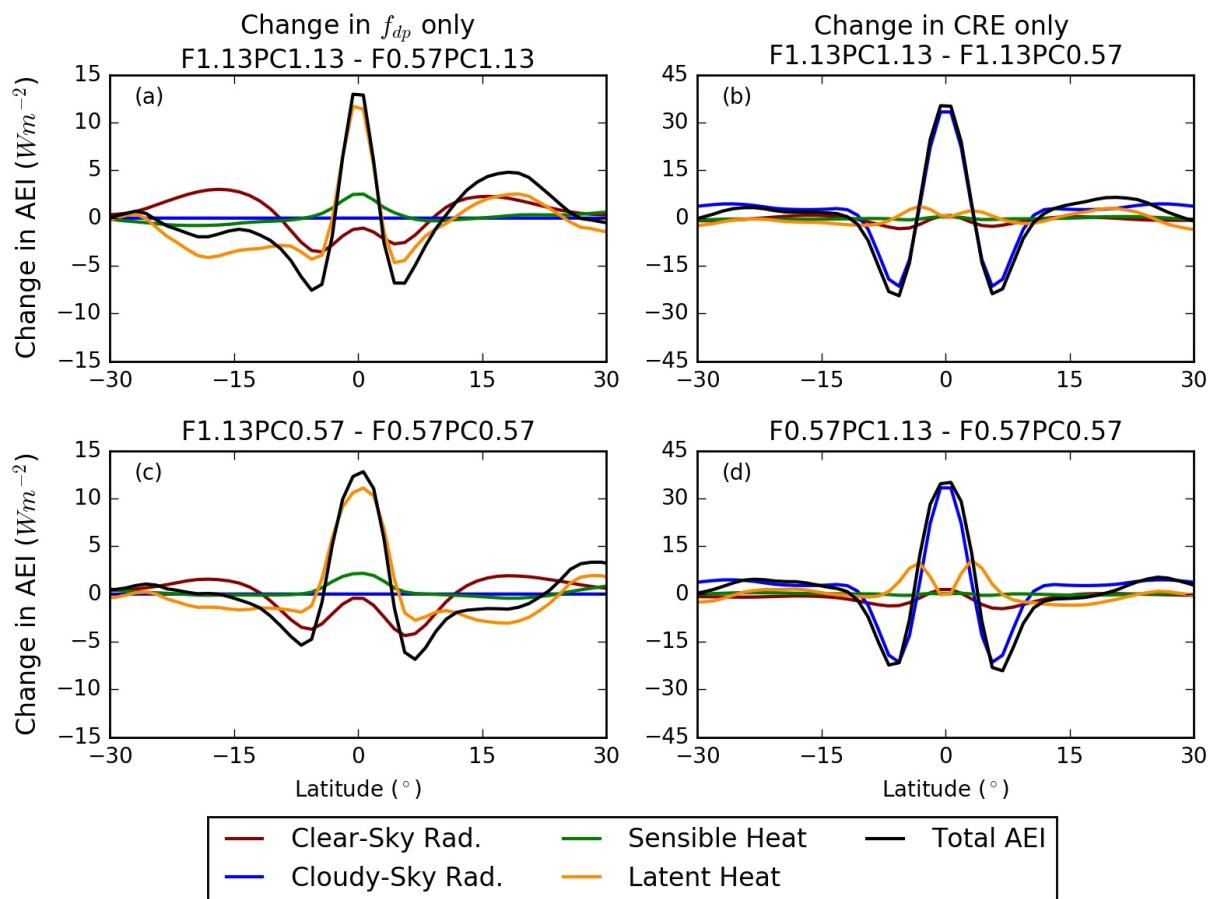
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